

Structure of the crust and upper mantle beneath Bass Strait, southeast Australia, from teleseismic body wave tomography

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We present new constraints on the lithospheric velocity structure of Bass Strait and the adjoining landmasses of mainland Australia and Tasmania in order to better constrain their geological and tectonic relationship. This is achieved by performing teleseismic tomography using data from fifteen deployments of WOMBAT and BASS transportable arrays, which span southeastern Australia. The starting model for the teleseismic tomography includes crustal velocity structure constrained by surface waves extracted from ambient seismic noise data and a Moho surface and broad-scale variations in 3-D upper mantle velocity structure from the Australian seismological reference Earth model (AuSREM). As a consequence, we produce a model with a high level of detail in both the crust and upper mantle. Our new results strengthen the argument for a low velocity upper mantle anomaly that extends down to ~150 km depth directly beneath the Newer Volcanics Province in Victoria, which is likely related to recent intra-plate volcanism. Beneath Bass Strait, which is thought to host the entrained VanDieland microcontinent, upper mantle velocities are low relative to those typically found beneath Precambrian continental crust; it is possible that failed rifting in Bass Strait during the Cretaceous, opening of the Tasman Sea, extension of VanDieland during Rodinian break-up and recent plume activity in the past 5 Ma may have altered the seismic character of this region. The data nevertheless suggest: (1) the velocity structure of the VanDieland microcontinent lacks continuity within its lithosphere; (2) the Moyston Fault defines an area of strong velocity transition at the boundary between the Cambrian Delamerian Orogen and the Cambrian-Carboniferous Lachlan Orogen; and (3) there is a rapid decrease in mantle velocity inboard of the east coast of Australia, which is consistent with substantial thinning of the lithosphere towards the passive margin.

KEY WORDS: teleseismic tomography, body waves, surface waves, upper mantle, Bass Strait, southeast Australia

1. Introduction

As a result of numerous studies carried out in southeastern Australia (Boger and Miller, 2004; Cawood, 2005; Cayley, 2011; White et al., 2013; Moresi et al., 2014; Pilia et al., 2015a, 2015b; Moore et al., 2015, 2016 - for earlier studies, see references in these publications), knowledge of the geological and tectonic relationship between mainland Australia and Tasmania, which are separated

39 by Bass Strait, has markedly improved over the past few decades (see Fig. 1 for geological map).
40 However, very little can be inferred about the lower crust and mantle lithosphere from these studies,
41 and consequently, a more complete picture of the region is yet to emerge.

42

43 The main difficulties in linking Tasmania and mainland Australia are the presence of Bass Strait,
44 which separates the two land masses, the presence of three thick Cretaceous to recent sedimentary
45 basins (Bass, Gippsland and Otway) and the widespread occurrence of recent intrusive and
46 extrusive volcanism, which have largely prevented direct observation of the underlying basement
47 rocks. Although gravity and magnetic data can provide useful insight into basement structures
48 (Cayley et al., 2011; Moore et al., 2015, 2016), the thickness of the failed-rift sedimentary basins
49 that formed during Australia-Antarctica breakup is a major obstacle for understanding the
50 Proterozoic-Palaeozoic connection between the lithosphere of Tasmania and mainland Australia. To
51 date, most models that attempt to connect Victoria and Tasmania either rely on the interpolation of
52 information between outcrops in both states, with some additional constraints provided by potential
53 field data (e.g. Leaman et al., 1994; Morse et al., 2009), geochemical investigations and seismic
54 reflection profiling (e.g. O'Sullivan et al., 2000; Drummond et al., 2000). One hypothesis that has
55 recently gained significant traction is that a fragment of Precambrian continental lithosphere, set
56 adrift by the break-up of the supercontinent Rodinia, became entrained in the convergent
57 accretionary system responsible for the formation of the Lachlan Orogen, and now underlies both
58 western Tasmania and southern Victoria (Cayley et al., 2002; Direen and Crawford, 2003; Glen,
59 2005; Cayley et al., 2011; Moresi et al., 2014; Moore et al., 2015, 2016). Ambient noise
60 tomography has also been employed to image crustal structure beneath Bass Strait, with initial
61 results pointing to the possible presence of a Precambrian continental fragment (Pilia et al., 2015a,
62 2015b). In other studies that are based on stratigraphic and geochemical considerations, Black et al.
63 (2004) found evidence for the presence of Precambrian lithosphere underlying Tasmania and King
64 Island, and Meffre et al. (2004) concluded that the Neoproterozoic mafic volcanics and intrusive
65 sequences of eastern King Island have many geochemical and stratigraphic characteristics that
66 indicate 579 Ma rifting associated with an extensive volcanic passive margin. All these studies
67 focus mainly on the crustal portion of the lithosphere; as such, the inclusion of data capable of
68 constraining the mantle lithosphere has the potential to shed new light on the structure and tectonic
69 evolution of Bass Strait.

70

71 In this paper, we apply teleseismic tomography to a dataset recorded by the high density WOMBAT
72 and BASS transportable seismic arrays, which comprise of over 620 stations, deployed during the
73 course of 15 separate array movements. These datasets are combined for the first time to constrain

74 3-D velocity perturbations in the upper mantle beneath Bass Strait. Teleseismic tomography utilises
75 relative arrival time residuals from distant earthquake sources, recorded across an array of seismic
76 stations, to image the seismic structure of the upper mantle beneath the array. A key assumption is
77 that relative arrival time residuals across the receiver array are largely unaffected by lateral
78 variations in the structure outside a local model volume defined beneath the array. Typically,
79 shallow structure to a depth approximately equal to the station spacing (Rawlinson et al., 2010) is
80 not constrained due to the near-vertical trajectories of impinging teleseismic rays. Here, we address
81 this limitation by incorporating an ambient noise crustal model in our starting model. Another
82 improvement, compared to traditional schemes, is the incorporation of an explicit Moho surface and
83 large-scale mantle velocity variation in the starting model, which are sourced from the AuSREM
84 model of the Australian continent (Kennett and Salmon, 2013).

85

86 Teleseismic body wave tomography has been frequently used in various parts of the world (Aki et
87 al., 1977; Oncescu et al., 1984; Humphreys and Clayton, 1990; Achauer, 1994; Steck et al., 1998;
88 Lippitsch et al., 2003; Rawlinson et al., 2006) but has never been applied to study the upper mantle
89 beneath Bass Strait. However, teleseismic tomography studies have been carried out in northern
90 Tasmania (Rawlinson et al., 2006) and southeastern mainland Australia (Graeber et al., 2002;
91 Rawlinson and Kennet, 2008; Rawlinson et al., 2011; Rawlinson and Fishwick, 2012; Rawlinson et
92 al., 2015; Rawlinson et al., 2016), but they have only been able to yield limited insight into the
93 relationship between the two landmasses. The inclusion of the BASS dataset, which spans Bass
94 Strait, together with adjacent data from the mainland and Tasmania, provides a unique opportunity
95 to image this region in detail using seismic techniques.

96

97 **2. Data and Method**

98 **2.1 Teleseismic data**

99 The teleseismic data for this study come from the WOMBAT transportable seismic array
100 deployments between 1998 and 2012 and the recent BASS array, which was deployed from 2011 to
101 2013 (Fig. 2). Each WOMBAT deployment involves between 20 to 70 stations deployed as a single
102 array, which after 6-12 months is transported to an adjacent location in order to achieve a high
103 density coverage over a large region. The average spacing between WOMBAT stations varies from
104 15 km in Tasmania to 50 km on the mainland. The seismometers used for this array are either 1 Hz
105 vertical component L4Cs (up until 2006), or 3-component 1Hz Lennartz LE-3Dlites. A number of
106 arrays (e.g. TIGGER and SEAL2) also incorporate several broadband seismometers (Guralp 40T or

107 3ESP).

108 In the case of the BASS array, the dataset consists of approximately 23 months (May 2011 to April
109 2013) of passive seismic data recorded by an array of 24 three-component broadband instruments
110 covering southern Victoria, several islands in Bass Strait (Flinders, King and Deal islands) and
111 northern Tasmania (Fig. 2). The instruments consist of 23 Guralp 40T and one Guralp 3ESP
112 broadband sensors. The 40T instrument has a flat response from 50Hz to 30s period and the 3ESP
113 has a flat response from 50Hz to 60s period. The instruments are connected to Earth Data PR6-24
114 portable field recorders.

115 The teleseismic dataset is characterised by dominant source regions to the north and east and a
116 much lower concentration of sources to the west and south of the study area. A total of 53,957 P-
117 wave relative arrival time residuals from 1558 events are used in the teleseismic tomography; since
118 we are using a transportable array (combined total of 624 receivers from 15 sequential
119 deployments), each event is only recorded by the sub-array which was operating at the time that it
120 took place. Teleseismic events with an angular distance $>27^\circ$ (from the centre of each array) and
121 $M_w \geq 5.5$ are considered. Figure 3 shows a plot of the locations of all teleseismic events used in this
122 study.

123 **2.2 Arrival time residual estimation and inversion**

124 The coherence of teleseismic arrivals across dense seismic arrays allows the relative onset times of
125 various global phases to be estimated with high accuracy. In this study, the adaptive stacking
126 scheme of Rawlinson and Kennett (2004) is used to extract residuals from the BASS array.
127 Residuals from the Tasmanian and mainland arrays were extracted using exactly the same
128 procedure described in previous studies (Rawlinson and Urvoy, 2006; Rawlinson et al., 2016).
129 Adaptive stacking exploits the similarity between waveforms to provide a rapid, accurate and robust
130 procedure for estimating the residual patterns across a network of seismic stations. Following any
131 pre-processing of the data (in our case, unifying instrument responses and band pass filtering
132 between 0.3-5.0 Hz), the first step of the scheme involves approximate alignment of all filtered
133 traces according to predictions from a propagation model (ak135), which are then stacked to form a
134 reference trace. Iterative improvement of the alignment, by comparison of the reference trace with
135 each station trace, leads to a direct estimate of the residuals relative to the 1-D propagation model.
136 This technique has proven to be fast and works well in the presence of noisy data and some
137 waveform distortion (Rawlinson and Kennett, 2004). Figure 5 shows an example of adaptive
138 stacking applied to traces from the BASS array experiment.

139 The arrival time residuals are mapped as 3D perturbations in P-wave velocity with respect to a
140 reference model using teleseismic tomography. The teleseismic tomography method assumes that
141 relative arrival time residuals across the array are largely unaffected by lateral structural variations
142 from outside the local 3-D model volume beneath the array. The software package FMTOMO was
143 used to perform the teleseismic tomography; FMTOMO uses the Fast Marching Method (FMM)
144 (Sethian, 1996; Sethian and Popovici, 1999; Rawlinson and Sambridge, 2004a, 2004b; de Kool et
145 al., 2006) for the forward prediction step and a subspace inversion scheme to adjust model
146 parameters in order to match observed and predicted traveltimes subject to damping and
147 smoothing regularisation (Kennett et al., 1988). The method is iterative non-linear in that the
148 inversion step assumes local linearity, but repeated application of FMTOMO and subspace
149 inversion allows the non-linear relationship between velocity and traveltimes perturbations to be
150 reconciled (see Rawlinson and Urvoy, 2006, and Rawlinson et al., 2016 for more information on
151 FMTOMO).

152 **2.3 3-D Starting model**

153 **2.3.1 Crustal model**

154 In order to mitigate a recognised weakness of teleseismic tomography- its inability to accurately
155 constrain shallow structure to a depth roughly equal to the station spacing (e.g. Rawlinson et al.,
156 2010) - a high resolution crustal model derived from ambient noise data that properly accounts for
157 the crustal contribution to the measured arrival time residuals was included in the starting model for
158 the inversion of WOMBAT and BASS teleseismic arrival time residuals (see supporting
159 information Text S1). As an alternative to the AuSREM crustal velocity model, the crustal model
160 derived from WOMBAT ambient seismic noise data contains well constrained features as small as
161 30–50 km horizontally and 2–5 km vertically. The robustness of this model is evident from the
162 features recovered (Fig. 4), which share a strong association with the surface geology and
163 constraints provided by potential field data in this region (Leaman et al., 1994; Morse et al., 2009).
164 As such, we believe that it is the better model to use for crustal corrections of the teleseismic
165 traveltimes residuals.

166 **2.3.2 Upper mantle model**

167 The AuSREM Moho model (Kennett et al., 2013) and broad scale mantle structure (Figs. 4c and 7
168 respectively) were also included in the starting model. It is well known that relative arrival time
169 residuals computed for adjacent arrays operating at different times are not directly comparable due
170 to the average velocity structures beneath them likely being different (e.g. Rawlinson et al., 2016).

171 Hence, unless we incorporate a robust 3-D reference model with a scale length less or equal to the
172 aperture of the smallest seismic array that is used, or have a backbone array of stations that record
173 throughout the durations of the transportable array experiment, then the solution model is likely to
174 contain artifacts, including the loss of long-wavelength structure. In our case, the AuSREM mantle
175 velocity model is sufficiently robust to allow the linking of the different WOMBAT and BASS
176 relative arrival time residual datasets in a joint inversion.

177 Seismic data exploited by AuSREM comes from over 12,000 km of deep crustal reflection profiles,
178 a number of major refraction experiments, and many different passive seismic deployments
179 (Kennett et al., 2013, Salmon et al., 2013). Due to moderate coverage of reflection data in this
180 region and constraints from receiver functions, the broadscale variations in AuSREM Moho
181 geometry are likely to be reasonably robust. However, this is not true for crustal P-wave velocity,
182 which is only constrained by a few 2-D refraction experiments in the model region (Salmon et al.,
183 2013); hence our use of an ambient noise crustal model in the 3-D starting model.

184 **2.4 Conversion of S to P velocity**

185 One limitation of the ambient noise model used for crustal correction is that it defines S-wave
186 velocities, whereas the teleseismic arrival time residuals come from global P-phases. We cannot use
187 S-phases in this study because the majority of WOMBAT stations are short period (~ 1 Hz corner
188 frequency), which makes teleseismic S-arrivals difficult to detect in comparison to P. In order to
189 convert from S-wave velocity to P-wave velocity, we apply the same approach as Rawlinson et al.
190 (2016), who use the AuSREM crustal V_p/V_s model (Salmon et al., 2013); although it features lower
191 lateral resolution compared to our new lithospheric model, the peak V_p/V_s perturbations are only
192 around $\sim 3\%$ which are unlikely to be a factor in the correction of P-wave arrival time residuals for
193 crustal structure.

194 **3. Results**

195

196 We represent the seismic structure of the southeast Australian lithosphere using a two-layer model
197 consisting of a crust and upper mantle separated by a Moho surface. Compared to previous
198 teleseismic studies that exploit WOMBAT data to image the lithosphere beneath southeast Australia
199 (Graeber et al., 2002; Rawlinson et al., 2006, 2011; Clifford et al., 2008; Rawlinson and Kennett,
200 2008; Fishwick and Rawlinson, 2012; Rawlinson et al., 2015, 2016), the work presented here uses a
201 geographically more extensive dataset that also covers Bass Strait and northern Tasmania.
202 However, the broad-scale variations in wave speed in regions that overlap with previous studies are

203 in general agreement; this is a reassuring sign when sequential studies are performed in the same
204 region using increasingly large datasets and/or different tomographic imaging schemes (Fishwick
205 and Rawlinson, 2012).

206 The inversion result was obtained by using six iterations of FMTOMO, with a starting model
207 defined by 459,000 velocity nodes and 7650 interface nodes. The data misfit component of the
208 objective function is based on the difference between observed residuals and model residuals. In the
209 latter case, these residuals are computed by taking the difference between predictions through the
210 recovered model (with receivers set at the correct elevation) at each iteration and ak135 reference
211 model predictions (with receivers set to zero elevation). Therefore, contributions to the predicted
212 residual from crustal velocity, Moho geometry and long wavelength mantle structure is accounted
213 for, as is the fact that adaptive stacking residuals were calculated using move-out corrections under
214 the assumption that all receivers were at the same elevation. It is important to note, therefore, that
215 the starting model for the inversion, which contains lateral structure in the crust, mantle and Moho,
216 is not the same as the laterally invariant model used to compute the reference arrival times. In order
217 to produce a model that does not greatly deviate from the starting model and contains no
218 unnecessary features, while at the same time having an acceptable fit to the data, damping and
219 smoothing values are chosen using the approach of Rawlinson et al. (2006) which is based on a
220 three step procedure using trade-off curves between data variance and model variance, and data
221 variance and model roughness (Fig. 6). This involves running a series of inversions with different
222 levels of damping and smoothing in order to identify the model with minimum perturbation and
223 roughness that still satisfies the data to an acceptable level. If we assume that the point of maximum
224 curvature on the trade-off curve provides the optimum compromise, then both the damping and
225 smoothing values are ~ 200 in our case (Fig. 6). However, it is well known that this approach is ad-
226 hoc, and following extensive testing, we ultimately prefer to use the same damping and smoothing
227 (75 in both cases) as Rawlinson et al. (2016) which produces a slightly underdamped and
228 undersmoothed version of the optimum model (according to the trade-off curves). We believe that
229 this model reveals features that are more consistent with crustal structure (from ambient noise
230 tomography) and broadscale surface geology. However, the differences between the two models are
231 not large, with our preferred model exhibiting a very similar pattern of anomalies, albeit with
232 slightly larger amplitude (see Supplementary Figs. S1-S3 for models produced using various
233 damping and smoothing values, including the “optimum” model according to the trade-off curves).
234 The final solution model, based on damping and smoothing values of 75, resulted in a variance
235 reduction in the traveltimes residual misfit of 73% relative to the starting 3-D model (and 78%
236 relative to ak135). Below, we present the results of a synthetic recovery test and the solution model

237 for mantle P-wave velocities.

238 **3.1 Synthetic test**

239 Because of solution non-uniqueness in seismic tomography problems a single data-satisfying model
240 may contain features that are not required by the data. Therefore, before interpreting our solution
241 model, the robustness of the data constraints is first investigated using a synthetic checkerboard test
242 (Glahn and Granet, 1993; Ritsema et al., 1998; Day et al., 2001; Rawlinson and Sambridge, 2003;
243 Rawlinson and Kenneth, 2008). This test involves using the identical sources, receivers and phase
244 types from the observational dataset to predict the arrival time-residuals for a synthetic structure
245 comprising an alternating 3D pattern of fast and slow velocities, separated by a narrow region of
246 zero perturbation (Fig. 8a and c). Here, we are only testing the resolving power of the teleseismic
247 dataset, because the crustal velocities are derived from ambient noise tomography alone and have
248 no dependence on the mantle velocities or teleseismic data. Hence, velocity anomalies are only
249 imposed on upper mantle structure, and both crustal velocity variations and the AuSREM Moho are
250 held fixed.

251 Prior to running an inversion for recovery of the synthetic checkerboard, Gaussian noise with a
252 standard deviation of 59 ms was added to the synthetic dataset. This is equal to the standard
253 deviation of the noise estimated by the adaptive stacking procedure used to extract the relative
254 arrival time residuals. The predicted data are then inverted using the same tomographic procedure
255 that is applied to the observations, including the same damping and smoothing values. Following
256 the inversion, the difference between the input and output models can be used as a measure of how
257 well structure can be recovered (Fig. 8b and d). This test is often favoured in seismic tomography
258 (e.g. Graeber and Asch, 1999; Rawlinson et al., 2011) because a continuous pattern makes
259 interpretation relatively straight forward; however, it cannot be regarded as a comprehensive test of
260 solution robustness (Lévéque et al., 1993; Rawlinson and Spakman, 2016).

261 The east – west cross-sections through the checkerboard inversion model (Fig. 9(A-A')) generally
262 exhibit good recovery, although there is some smearing of structure towards the edge of the array
263 caused by a predominance of paths with similar orientations. The north – south cross-sections (Fig.
264 9(B-B')) also recover some of the checkerboard pattern, although smearing is clearly evident
265 beyond the horizontal bounds of the array. The smearing is dominant in the north-south direction as
266 a consequence of source distribution which results in many more paths impinging on the model
267 from the north compared to the south. Similarly, the centre of Bass Strait is poorly recovered at
268 shallow depth, which can be attributed to an absence of receivers (such as ocean bottom

269 seismometers) in the marine environment. The minimum scale length of structure that can be
270 recovered by teleseismic tomography is typically equal to the station spacing (Rawlinson et al.,
271 2015), which in this case is around 50 km, similar to the mainland station spacing. In summary, the
272 upper mantle structure appears to be well constrained by the path coverage, although resolution
273 tends to decrease beyond the horizontal bounds of the seismic array (see Supplementary Fig. S4 for
274 additional checkerboard recovery tests).

275 **3.2 WOMBAT-BASS model**

276 The model recovered from inversion of WOMBAT and BASS teleseismic datasets is shown in
277 Figures 10, 11 and 12 (note that the grey mask denotes less resolved areas as determined by
278 inspection of checkerboard test results), and reveals a number of significant structural features. The
279 horizontal slice at 75 km depth (Fig. 10) shows an E-W (latitude 37.2° S) trending low velocity
280 anomaly in the upper mantle in southern Victoria (see dashed rectangle SV), which lies to the north
281 of a narrow high velocity anomaly located along the SE Victorian coast. Bass Strait appears as a
282 relatively low velocity zone (noting that it is not well resolved in the centre at this depth) while
283 upper mantle velocities in northern Tasmania vary from fast to slow to fast from west to east (see
284 dashed rectangle NT in Fig. 10). North of the line A-A' in Figure 10, the model is largely consistent
285 with the body wave tomography model of Rawlinson et al. (2016), which uses the same mainland
286 arrival time residual dataset from WOMBAT. At greater depths beneath the southern region of the
287 model, the pattern of velocity anomalies becomes less coherent, but the lower velocities beneath
288 southern Victoria and Bass Strait and the higher velocities to the north extend to well over 100 km
289 depth (B-B' in Fig. 11). In general, the pattern of residuals in the crust and Moho variations do not
290 appear to share an overall anti-correlation with uppermost mantle anomalies, which at least suggests
291 that crustal structure is not dictating upper mantle structure.

292 In order to reinforce the importance of a 3-D starting model in improving the recovery of mantle
293 structure, a comparative test was carried out. Here, the inversion of the data was carried out using a
294 purely 1-D starting model (ak135) in the crust and upper mantle, and a flat Moho consistent with
295 ak135. The recovered model in Figure 13 shows clear similarities and differences with the solution
296 produced using a 3-D starting model (Fig. 12). The lack of longer wavelength structures in Figure
297 13, particularly evident in the cross-sections, would make interpretation more challenging if this
298 model was considered in isolation. Furthermore, while there is general agreement on the shorter
299 wavelength structures, there are differences evident, particularly at shallow mantle depths. Of
300 particular relevance to this study is the presence of high velocity anomalies appearing to connect
301 Tasmania and Victoria across Bass Strait in the 75 km and 120 km depth slices of Figure 13. This

302 may lead one to conclude that the two share a structural connection in the mantle lithosphere, but as
303 Figure 12 demonstrates, these anomalies are more likely to be artifacts caused by the vertical
304 smearing of unresolved crustal structure.

305 **4. Discussion**

306 The inversion of the WOMBAT-BASS dataset has produced a high resolution lithospheric velocity
307 model of Bass Strait and adjoining southeast mainland Australia and Tasmania. Many intriguing
308 features are observed in our P-wave velocity model that could pertain to tectonic processes that
309 affected the study area. Several velocity anomalies imaged in our study are present in previously
310 determined models of northern Tasmania and southern Victoria. These include the high velocities
311 found in the upper mantle beneath the Rocky Cape Block of north-west Tasmania (Rawlinson and
312 Urvoy, 2006; Rawlinson and Kennett, 2008) and regions of low mantle velocity beneath the Tamar
313 Fracture System (TFS) and Newer Volcanics Province (NVP) (Rawlinson and Fishwick, 2012;
314 Rawlinson et al., 2015; Rawlinson et al., 2016). Further north, higher velocities associated with
315 cratonic regions and lower velocities near the eastern seaboard, which are likely due to lithospheric
316 thinning towards the passive margin, can be observed (Rawlinson et al., 2016). A recent study that
317 also shows consistency with our results in terms of lower velocities underneath the Lachlan Fold
318 Belt and Bass Strait is that of Sun et al. (2016). In their study, they apply 3-D P_n travel time
319 tomography to improve structural constraints on the uppermost mantle across the Australian
320 continent. They find some of the lowest velocities beneath the southern Lachlan Fold Belt, the
321 Newer Volcanics province and Bass Strait, which largely agrees with our teleseismic body wave
322 tomography model, and supports the idea that recent intraplate volcanism, coupled with lithospheric
323 thinning associated with failed rifting, have strongly influenced the seismic structure of the
324 lithosphere in this region.

325 Our new tomographic model exhibits a distinct north dipping low velocity anomaly that extends
326 from just below the crust to ~ 150 km depth in the upper mantle directly beneath the NVP (see Fig.
327 12 cross-section A-A'). Previously, Davies et al. (2015) identify the world's longest continental hot-
328 spot track (over 2000 km total length), which begins in north Queensland, and extends southward,
329 possibly as far as NW Tasmania. Their study suggests that the plume source of the hotspot track
330 passed by the eastern edge of the NVP shortly before volcanism initiated nearly 5 Ma. In another
331 study, Rawlinson et al. (2017) suggest that an interaction between this passing plume and pre-
332 existing Edge Driven Convection (EDC) (King and Anderson, 1998) and Shear Driven Upwelling
333 (SDU) (King and Ritsema, 2000) (caused by a huge central cavity in the lithosphere, which
334 corresponds to the central low velocity zone north of the NVP in our model) was responsible for the

335 NVP volcanism. A more recent study by Wei et al. (2018) using 3-D S_n traveltimes tomography
336 finds strong heterogeneities in S wave speed in the upper mantle across the entire Australian
337 continent. They also include Pn data to determine the uppermost mantle V_p/V_s ratio across the
338 whole continent. Their study finds distinctive localised low velocity anomalies beneath the eastern
339 margin of Australia which they interpret as evidence of mantle plume remnants. Our model shows a
340 low velocity zone beneath Bass Strait (Fig. 12 cross-section C-C'), but below 200 km depth these
341 lower velocities disappear entirely. A recent study has suggested that the plume was waning as
342 Australia gradually moved north (Jones and Verdel, 2015), so it is possible that there is no longer
343 any clear evidence for it at greater depth. Nonetheless, it may have contributed to the lower
344 velocities observed beneath Bass Strait.

345

346 Our seismic tomography results appear to be consistent with the findings of previous studies
347 regarding the boundary between the Lachlan and Delamerian Orogens (Rawlinson et al., 2015;
348 Rawlinson et al., 2016). The mantle model of Rawlinson et al. (2016), which covers mainland
349 southeast Australia, shows a clear transition from high velocities in the west to lower velocities in
350 the east. This same velocity transition is clearly imaged in our model (Fig. 11). Moreover, the
351 location and dip direction of the velocity transition (marked “DLB” in section B–B' of Fig. 12) is
352 consistent with the Moyston Fault at the surface (Cayley and Taylor, 1998; VandenBerg, 1999;
353 Cayley et al., 2002), which is a widely accepted candidate for the Delamerian–Lachlan boundary in
354 the crust. However, while the velocity transition in the mantle is unlikely to represent a continuation
355 of this fault, it may be the signature of an W-E contrast from Precambrian to Palaeozoic mantle
356 lithosphere.

357 A long standing source of debate in southeast Australian tectonics is the possible connection
358 between the mainland Delamerian Orogen and the Tyennan Orogen in Tasmania, where there have
359 been two schools of thought. Crawford et al. (2003) and Reed et al. (2002) proposed a model
360 depicting the two orogens connected directly. This connection is reinforced by several studies that
361 have examined the age and geochemistry of various igneous rocks in both regions, and found strong
362 similarities. For example, Direen and Crawford (2003), have found paired suites of unusual ~580
363 Ma olivine rich mafic volcanics, together with rare ~515 Ma boninites in both western Victoria and
364 the mainland Delamerian Orogen. However, subsequent studies (e.g. Cayley et al., 2011; Moore et
365 al., 2015, 2016) argued that the two orogens are not connected directly. In particular, Moore et al.
366 (2015) found that the Tyennan and Delamerian Orogenies occurred at the same time but are
367 displaced along strike. In their model, the Tyennan Orogeny takes place further to the east and is
368 controlled by the master subduction zone to the Pacific plate.

369 In light of the above discussion, it is interesting to note that the velocity transition “DLB” observed
370 on the mainland does not show any extension southwards through Bass Strait to Tasmania.
371 However, one of the challenges in trying to interpret structure that was emplaced in the distant past
372 from present day geophysical images is that subsequent deformation and alteration will mask pre-
373 existing features and make any meaningful interpretation difficult. Notable tectonic events that
374 post-date the formation of the Delamerian and Lachlan Orogens include the break-up of Australia
375 and Antarctica and opening of the Tasman Sea, which resulted in the formation of the associated
376 failed rift basins that accommodate thick sedimentary sequences (Gaina et al., 1998). These events
377 also resulted in substantial thinning of the lithosphere towards the passive margin. The lower
378 velocities observed beneath the Bass Strait-Tasmania region (Fig. 12 cross-section C-C'), may be a
379 result of elevated temperatures associated with the rifting and the presence of a thinner lithosphere,
380 as well as the influence of the plume associated with the Cosgrove track. Hence the interpretation of
381 Cayley et al. (2011) and Moore et al. (2015, 2016) may be consistent with our observation of the
382 DLB not extending south into Bass Strait, implying no direct connection between the Delamerian
383 and Tyennan Orogens.

384 Recent studies have strengthened the argument for the presence of a Proterozoic continental
385 fragment, sourced from the break-up of Rodinia, beneath the Melbourne Zone in southern Victoria
386 (the so-called Selwyn Block), which extends south through Bass Strait to form the Precambrian
387 basement of western Tasmania (Cayley, 2011; Moresi et al., 2014; Pilia et al., 2015a, 2015b; Moore
388 et al., 2015, 2016). Following the earlier work of Cayley (2011), the existence of an entrained
389 Precambrian continental fragment has been reinforced by using 3-D geodynamic modeling (Moresi
390 et al., 2014). This study suggests that an exotic continental fragment became entrained in the proto-
391 Pacific convergent margin of eastern Australia, which resulted in a complex sequence of processes
392 including differential subduction roll-back and southward transfer of material through continental
393 transform faulting to form a large orocline that now underlies the Lachlan Orogen. The models
394 illustrate how significant curvature of the orogenic system develops, as well as the mechanism for
395 tectonic escape of the back arc region.

396 Moore et al. (2015, 2016) provide robust geological and potential field evidence for the processes
397 that led to the formation of Tasmania and how it is connected to mainland Australia. Their studies
398 suggest that an extensional event along the Proto-Terra Australis margin began Rodinian breakup at
399 approximately 830 Ma. As the extension continued, rifting took place along the Rodinian margin at
400 around 600 Ma, which resulted in the separation of Laurentia, leaving vanDieland (comprising of
401 the Selwyn Block, western Tasmania and the western part of the Bass Strait) attached to a thinned
402 Gondwana margin. By 580 Ma vanDieland was completely rifted from Rodinia, forming three or

403 four weakly connected cratonic ribbons separated by thinned continental or back-arc oceanic crust
404 and a completely detached King Island. At the beginning of Ross Orogeny subduction, the rifted
405 fragments of vanDieland drifted northwards at approximately 560 Ma. By 510 Ma or shortly
406 thereafter, trench advance during the Tyennan Orogeny caused closure between the cratonic blocks
407 bringing most of them together. Finally, the micro continental ribbons became accreted within a
408 closing back arc system and later, together with eastern Tasmania and the Lachlan Orogen,
409 coalesced with the Gondwana Craton in the Middle Devonian. By 495 Ma all subduction processes
410 had stalled and post collisional extension was underway. Moore et al. (2015, 2016) also
411 demonstrate how the accreted vanDieland ribbons can be traced from present day Tasmania across
412 Bass Strait into the Selwyn Block, thereby providing evidence that vanDieland is embedded in the
413 Lachlan Orogen in southeast mainland Australia.

414 Pilia et al. (2015a) studied Bass Strait and adjoining land masses by applying a transdimensional
415 Bayesian inversion scheme to WOMBAT and BASS data to retrieve group velocity maps and 3-D
416 shear wavespeed. The result is an image of the entire crust beneath Bass Strait that reveals a high
417 velocity anomaly connecting southern Victoria and western Tasmania in the mid-lower crust, which
418 is suggestive of a Proterozoic geological connection. This high velocity anomaly is consistent with
419 recent gravity and magnetic maps and results from a seismic reflection survey through the
420 Melbourne Zone in southern Victoria (Cayley et al., 2002, Cayley et al., 2011).

421 Our new model shows that high velocity anomalies in the upper mantle tend to underlie crust of
422 Precambrian continental origin, for example the Curnamona Province and Gawler Craton in Figure
423 12. Contrary to this, apart from the narrow sliver beneath the Melbourne zone, there is no evidence
424 of a high velocity body in the upper mantle connecting mainland Australia and Tasmania.
425 However, as noted previously, this does not necessarily rule out the existence of Precambrian
426 continental material. Apart from the influence of recent plume activity, two possible explanations
427 include delamination of the mantle lithosphere beneath the fragment triggered by edge driven
428 convection and shear driven upwelling (Rawlinson et al., 2017), or so-called rift-induced
429 delamination (Wallner and Schmeling, 2010). Interestingly, mantle velocities below 200 km depth
430 beneath Bass Strait (Fig. 12, cross-section C-C') are anomalously high compared with mantle
431 beneath the Australian mainland, which potentially supports this hypothesis. However, the lack of
432 any apparent uplift in Bass Strait, which would normally be a required response to the loss of a
433 dense lithospheric mantle layer, means that this is not very likely. Furthermore, the deep high
434 velocity zone is in part inherited from AuSREM (Fig. 7d), which may not be well constrained in
435 this region. A third possible explanation is that the micro-continent was subjected to intense
436 lithospheric rifting during 250 Myr of Neoproterozoic extension which lead to Rodinian break-up,

437 as suggested by Moore et al. (2015). This resulted in the development of continental ribbons
438 (boudins), which by 510 Ma or shortly thereafter were brought together by the Tyennan Orogeny.
439 This Rodinian extension is likely to have modified and thinned any high velocity mantle, resulting
440 in a similar velocity signature to present day rifted margins.

441 Previous studies (e.g. Pilia et al., 2015a, Rawlinson et al., 2016) have also suggested that any high
442 velocity associated with a lithospheric keel beneath an entrained Precambrian continental fragment
443 may have been overprinted by failed rifting associated with the break-up of Australia and Antarctica
444 and the opening of the Tasman Sea between 80 and 90 Ma (Gaina et al., 1998). Both of these events
445 would have resulted in substantial thinning of the lithosphere in the Bass Strait region (marked 'BS'
446 in Fig. 12, cross-section C-C'), and potentially elevated upper mantle temperatures. The rapid
447 eastward decrease in upper mantle velocities towards the east coast of Australia (marked 'LTZ' Fig.
448 12) is consistent with lithospheric thinning associated with the opening of the Tasman Sea. This
449 change is due to the proximity of the passive margin and may be accentuated by the presence of
450 asymmetric rifting, with the upper-plate margin underlying the Australian east coast (Lister et al.,
451 1986; Lister and Etheridge, 1989; Lister et al., 1991).

452 **5. Conclusions**

453 In this study, a P-wave velocity model of the crust and upper mantle was used to investigate the
454 broad-scale geology of the Bass Strait region. The inversion of teleseismic arrival time residuals
455 from WOMBAT and BASS transportable array data for mantle velocity variations was improved by
456 embedding a detailed ambient noise crustal model and AuSREM Moho and mantle velocity
457 variations in the starting model. Resolution tests show that the new model of the Bass Strait region
458 is generally well constrained by the data set beneath the horizontal bounds of the array, although
459 there is a lack of coverage beneath the centre of Bass Strait to a depth of approximately 100 km due
460 to the use of only land-based stations. North of Victoria, the pattern of anomalies is consistent with
461 previous studies that have exploited the WOMBAT teleseismic dataset. We also image a distinct
462 low velocity zone beneath the NVP in western Victoria, and show that (i) it is restricted to the upper
463 mantle, and appears to die out at around 150 km depth; and (ii) it lies at the northern end of a
464 pervasive low velocity zone in the upper mantle that stretches between Victoria and Tasmania. As
465 such, we do not find any evidence to support the existence of a Precambrian microcontinent that
466 joins Tasmania to the mainland, as has been suggested in recent studies. However, we cannot rule
467 out the possibility that hyper-extension during Rodinian break-up, failed rifting in Bass Strait
468 associated with the separation of Australia and Antarctica, the opening of the Tasman Sea and
469 mantle sources of recent intra-plate volcanism may have conspired to reduce seismic wavespeeds

470 via processes such as lithospheric stretching, delamination, changes in bulk composition and
471 heating.

472

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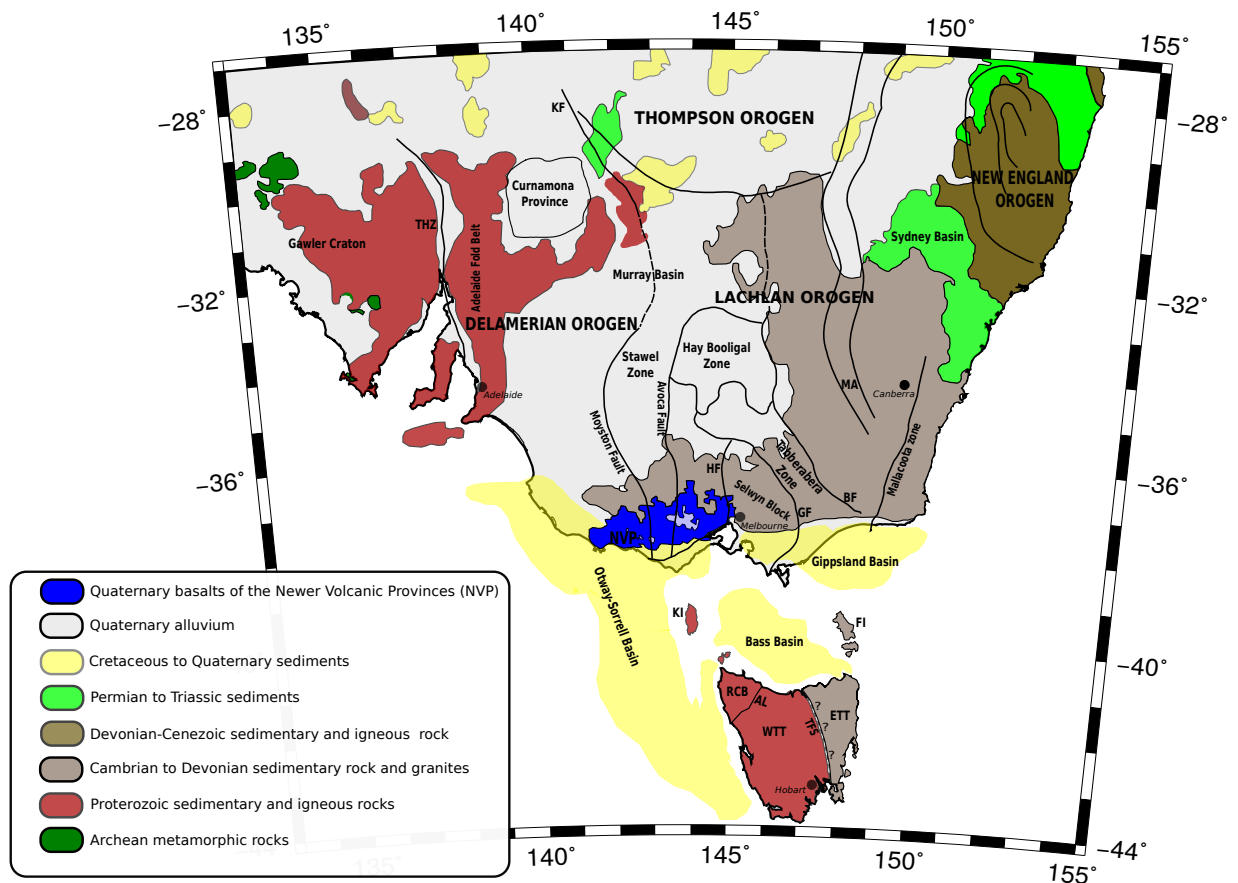


Fig. 1: Simplified geological map of southeastern Australia showing observed and inferred geological boundaries and main tectonic features mentioned in the text. Thick black lines show locations of structural boundaries. KF = Koonenberry Fault; HF = Heathcote Fault; GF = Governor Fault; BF = Bootheragandra Fault; THZ = Torrens Hinge Zone; NVP = Newer Volcanic Province; MA = Macquarie Arc; KI = King Island and FI = Flinders Island in Bass Strait; WTT = Western Tasmania Terrane; ETT = Eastern Tasmania Terrane; AL = Arthur Lineament; TFS = Tamar Fracture System and RCB = Rocky Cape Block.

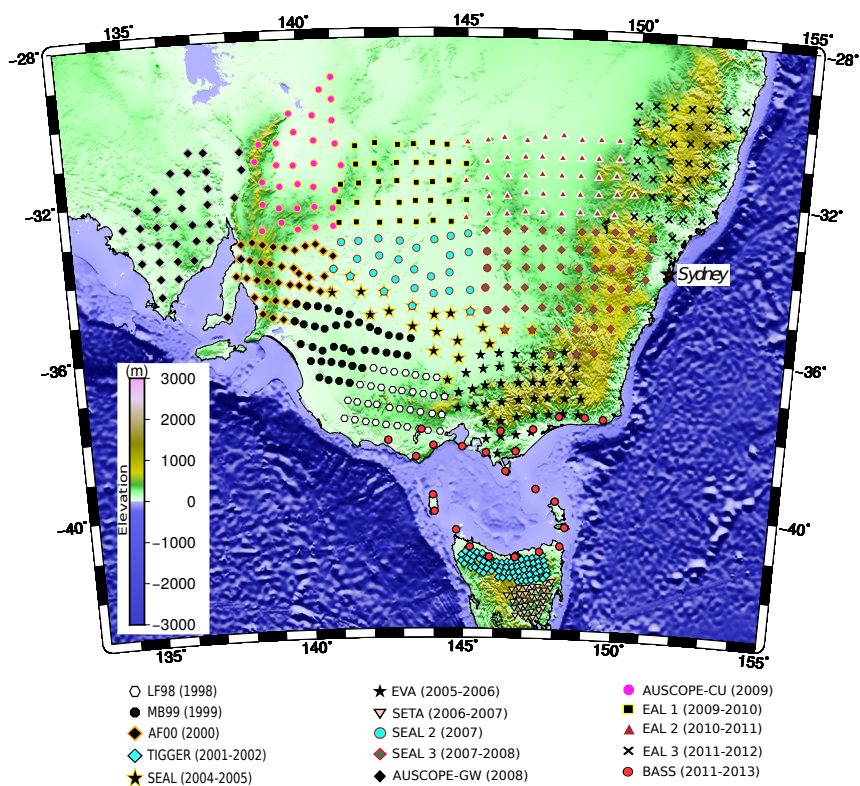


Fig. 2: Topographic/bathymetric map of the study area showing locations of the fifteen different arrays represented by different symbols.

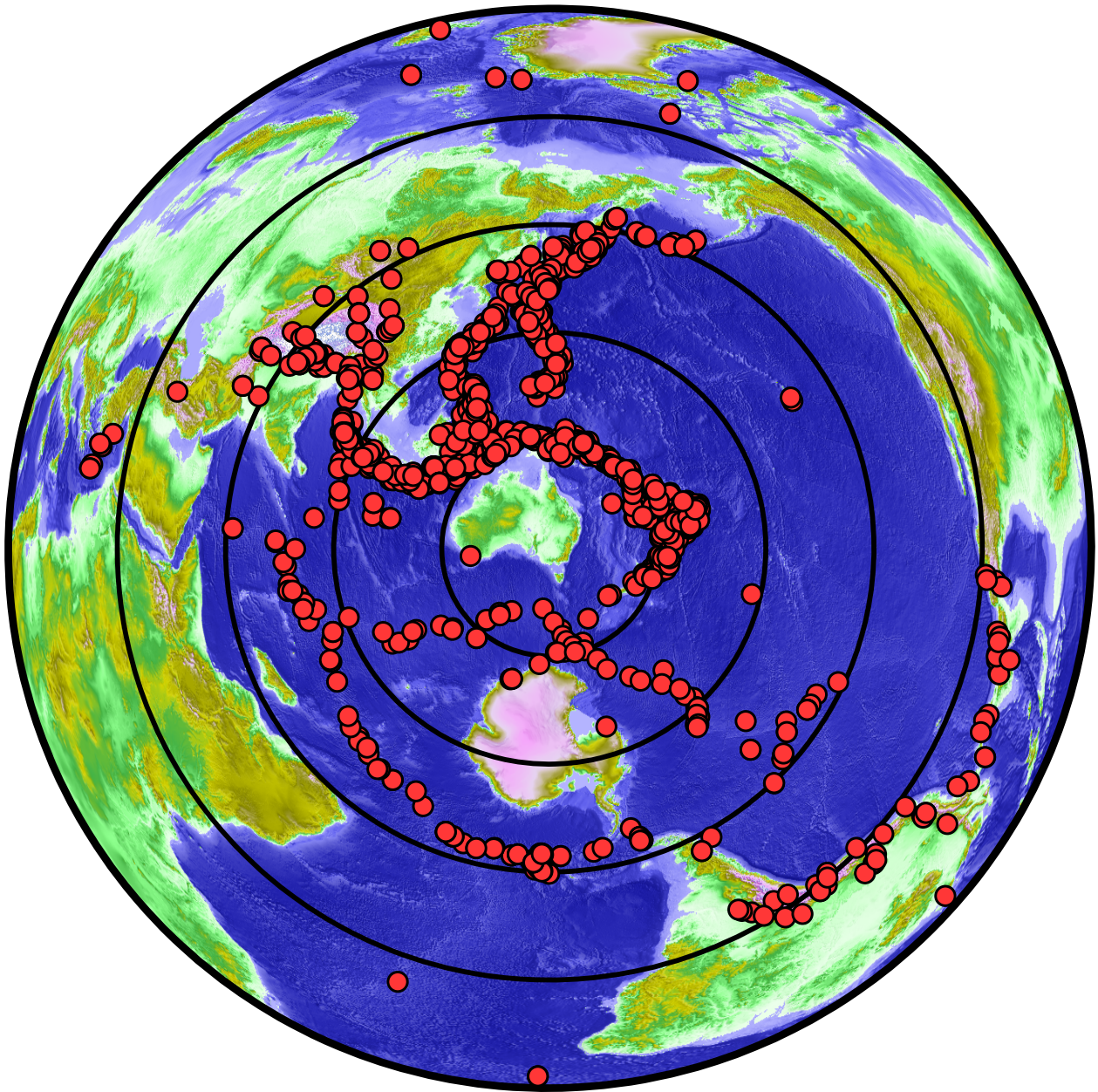


Fig. 3: Distribution of teleseismic events used to constrain the 3-D P wave velocity model. Concentric circles are plotted at 30° intervals from the centre of Bass Strait.

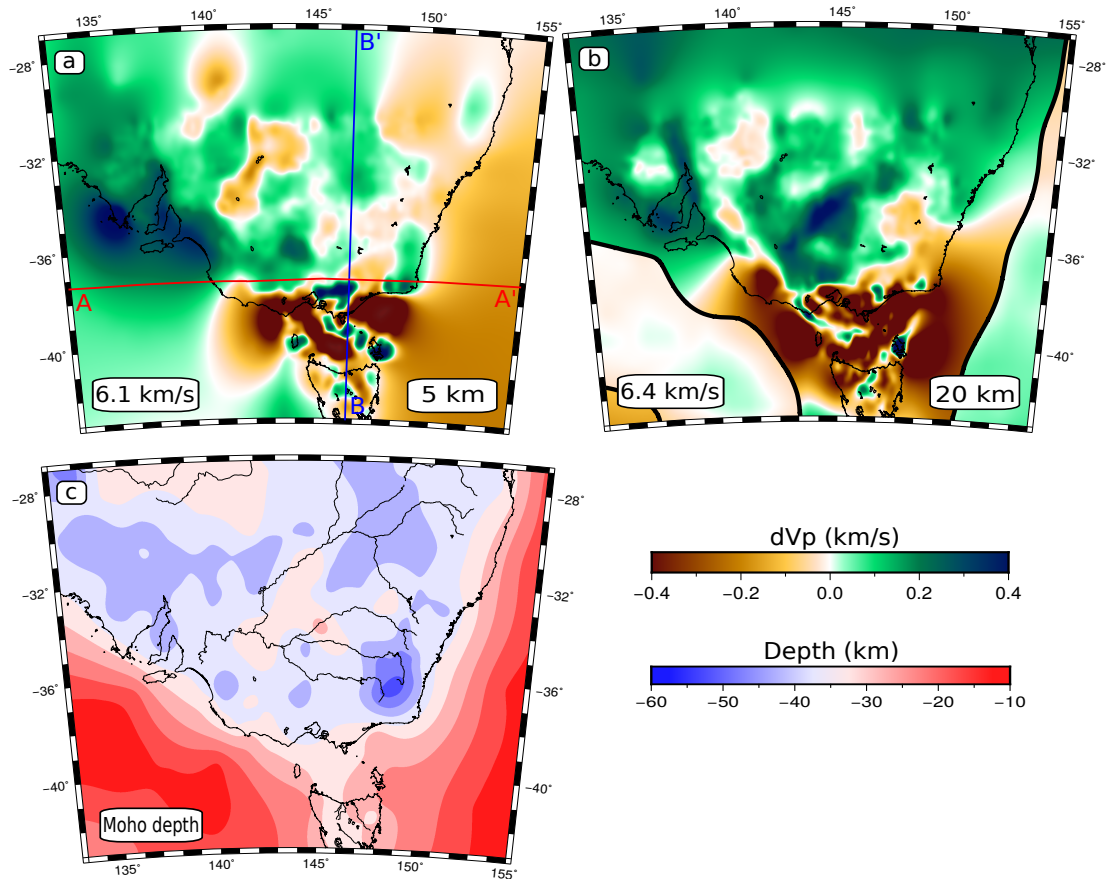


Fig. 4: (a,b) Two depth slices through the new ambient noise crustal model at 5 km and 20 km depth. Since it is the relative velocity variations that are important for teleseismic tomography, perturbations from a laterally averaged depth-dependent 1-D version of the model (see bottom left corner of a,b for reference velocity values) are shown. Thick black lines (see 20 km depth slice in (b)) denote the intersection of the depth slice with the Moho. (c) AuSREM Moho depth model for southeast Australia, which is also used in the starting model of the teleseismic tomography. We have converted the original S-wave velocity model to P-wave velocity, which is what is shown here.

Fiji Islands; 16/05/2012; Mw=6.6

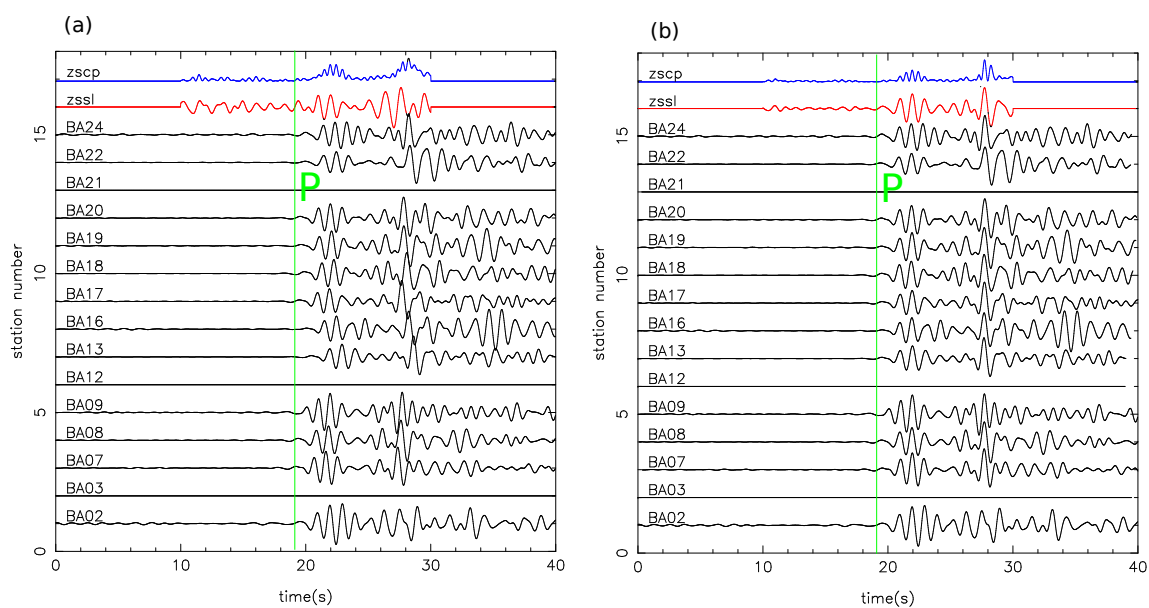


Fig. 5: Example of adaptive stacking of recorded P waves at BASS stations. (a) are the filtered traces before stacking; (b) the filtered traces after stacking. The top blue and red coloured traces are the quadratic stack (zscp) and linear stack (zssl) respectively. The vertical green lines indicate predictions from the global reference model ak135, to which all traces have been aligned in (a).

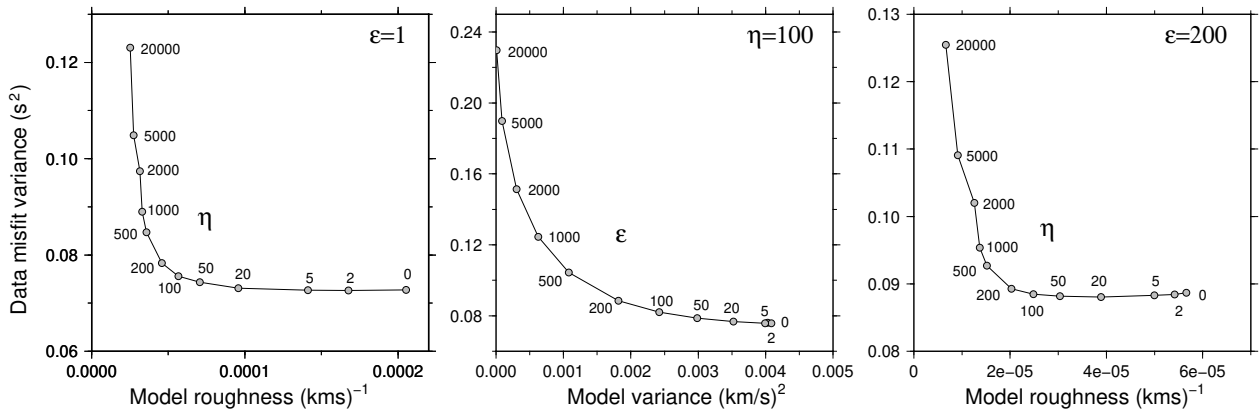


Fig. 6: Trade-off curves used to find optimum damping and smoothing parameters. (a) Vary smoothing parameter while holding damping parameter fixed at $\epsilon = 1$. (b) Vary damping parameter while holding smoothing parameter fixed at $\eta = 100$. (c) Vary η while holding the damping parameter fixed at the new value of $\epsilon = 200$. In all the three cases, $\epsilon = \eta = 200$ appears to be an acceptable choice.

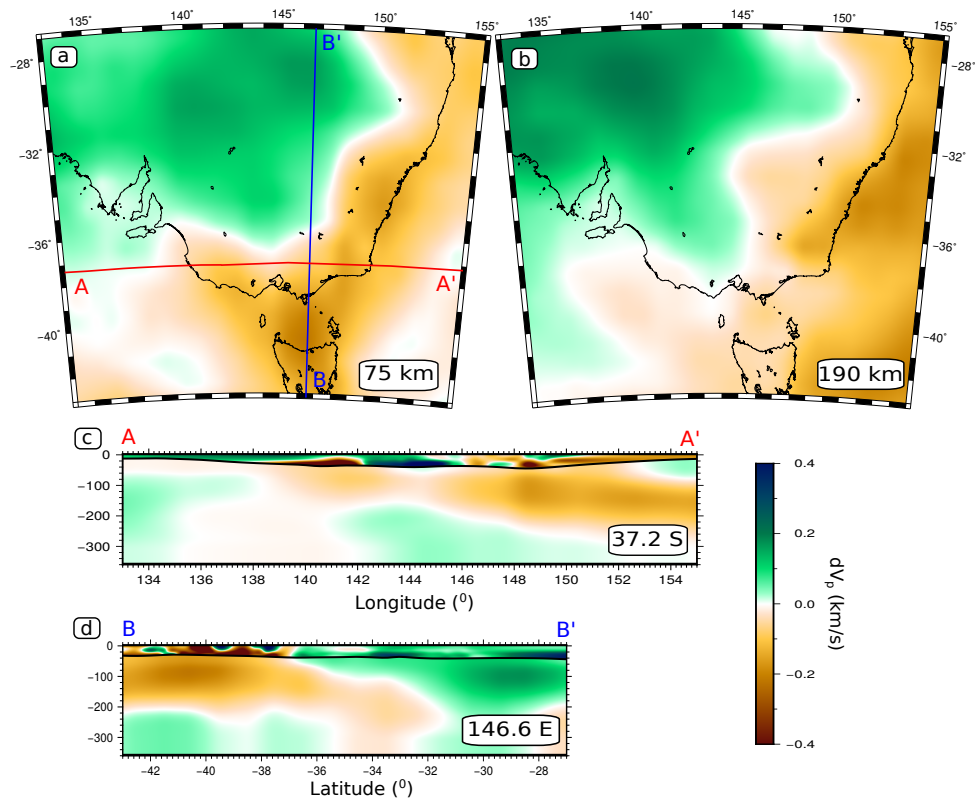


Fig. 7: A selection of slices through the starting model used for teleseismic tomography. (a) 75 km depth slice; (b) 190 km depth slice; (c) EW cross-section at 37.2°S; (d) NS cross-section at 146.6°E. In all cases, velocities are plotted relative to a reference one-dimensional model. The horizontal red line and vertical blue line in (a) denote the cross-section locations of (c) and (d) respectively.

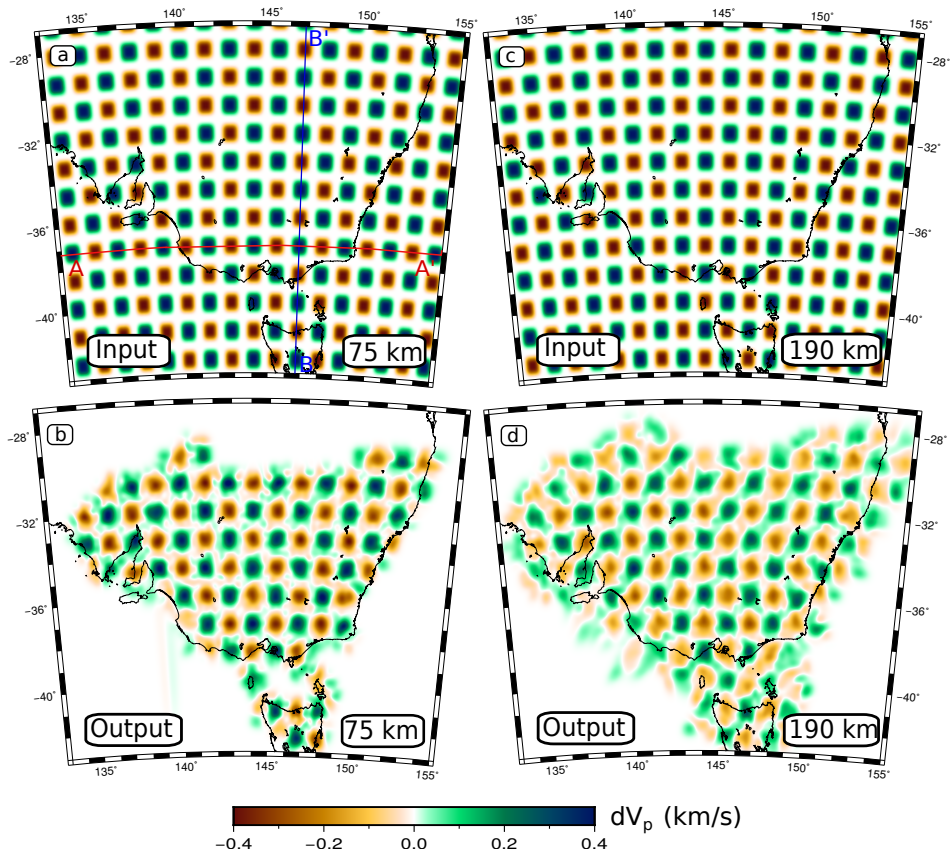


Fig. 8: Depth sections illustrating the synthetic resolution test results. (a) Input test model at 75 km depth; (b) recovered model at 75 km depth; (c) input test model at 190 km depth; (d) recovered model at 190 km depth. The horizontal red and vertical blue lines in (a) denote the cross-section locations of Fig. 9a and b respectively.

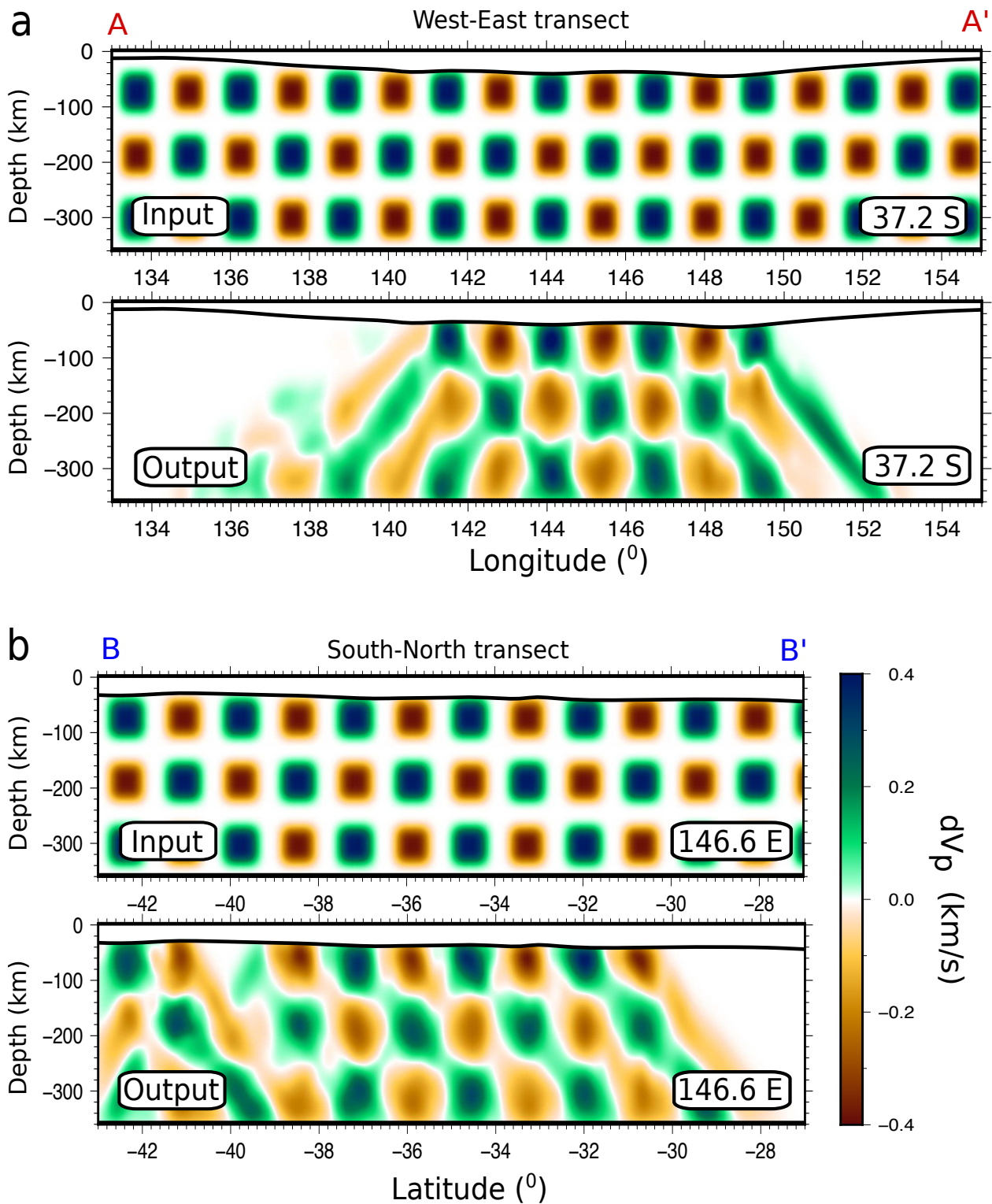


Fig. 9: Cross-sections illustrating the synthetic resolution test results. (A-A') Input and output test models at 37.2°S (see red line in Fig. 8a for location); (B-B') input and output test models at 146.6°E (see blue line in Fig. 8a for location).

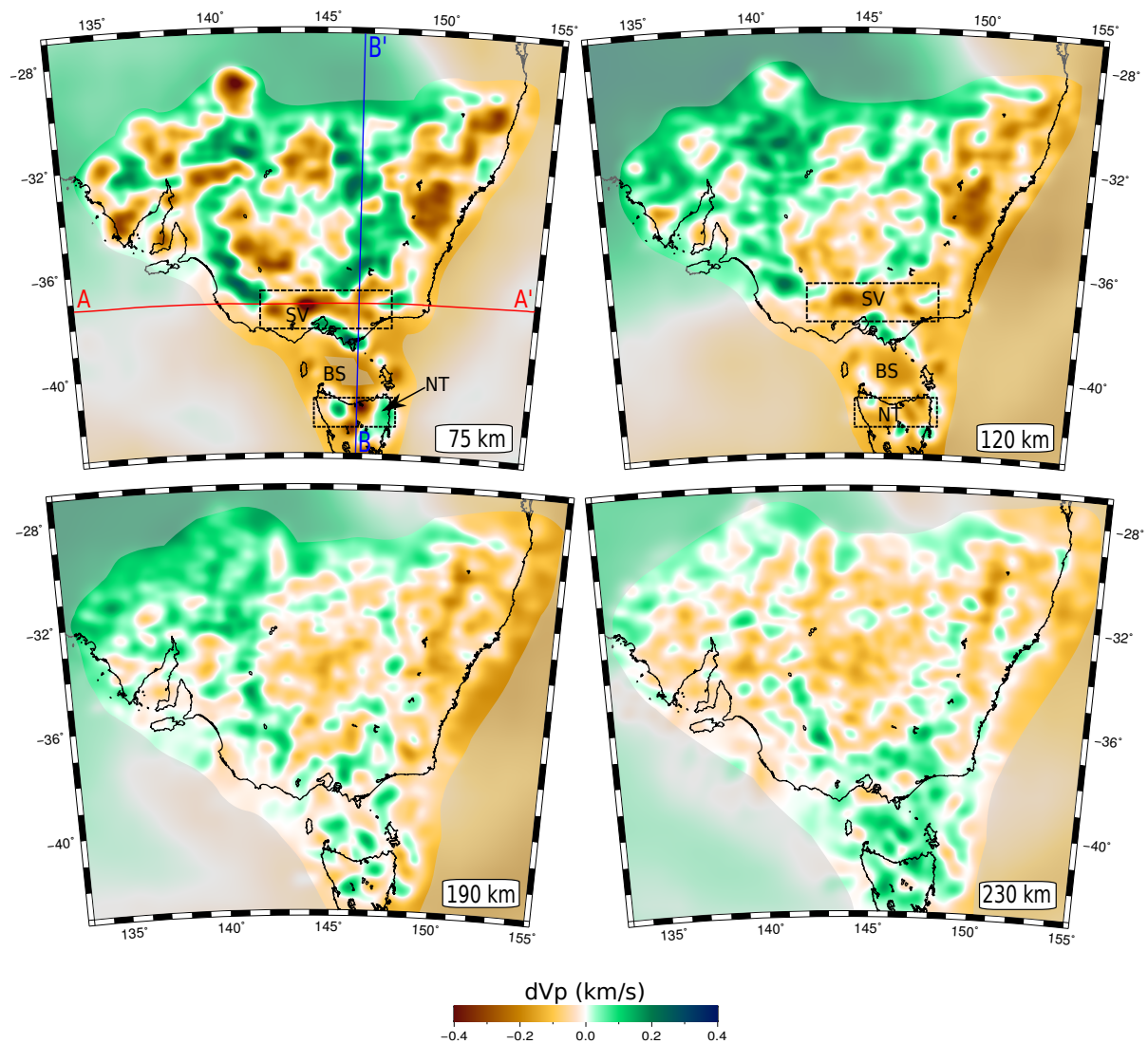


Fig. 10: Horizontal slices through the solution model obtained by inversion of relative arrival time residual data from the WOMBAT and BASS arrays. P-wave speed perturbations are relative to a 1-D depth-averaged version of the model. The gray mask denotes less resolved areas as determined by inspection of checkerboard test results for each depth slice. Several features are highlighted which are discussed in detail in the text: SV = Southern Victoria; NT = Northern Tasmania; BS = Bass Strait.

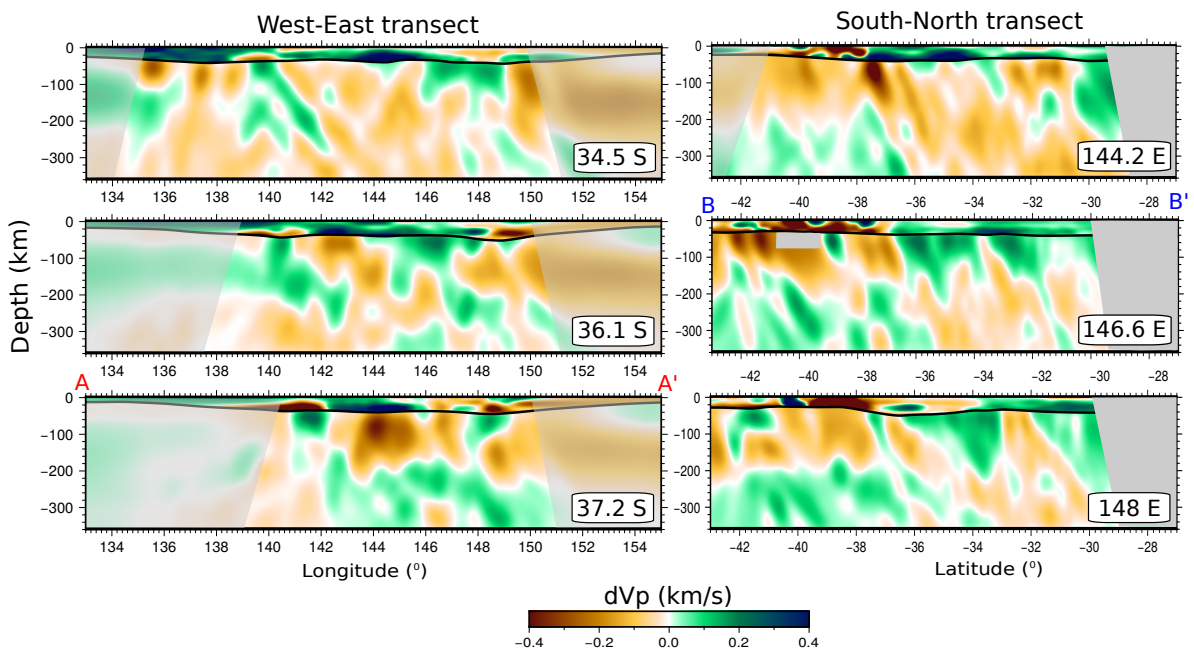


Fig. 11: Vertical slices through the solution model obtained by inversion of relative arrival time residual data from the WOMBAT and BASS arrays. As in Fig. 10, P-wave speed perturbations are relative to a 1-D depth-averaged version of the model. The gray mask denotes less resolved areas as determined by inspection of checkerboard results for each profile. See blue and red lines in Fig. 10 for locations of AA' and BB'.

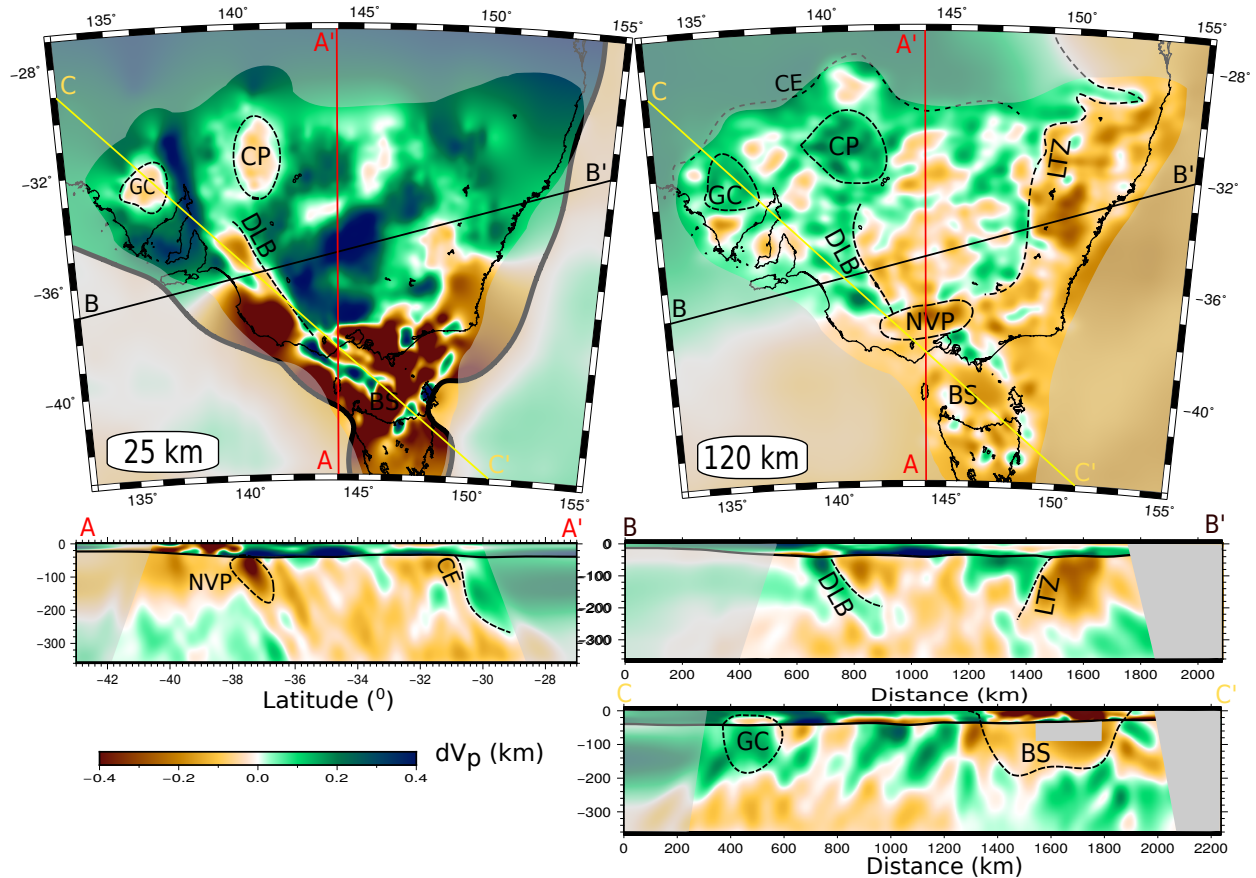


Fig. 12: Slices through the 3-D crust and upper mantle model of southeast Australia which uses the ambient noise crustal model and AuSREM Moho and upper mantle model in the starting model for the teleseismic data inversion. The locations of the three cross sections are denoted in both depth sections as coloured lines. Velocities are plotted relative to a reference one-dimensional model. The gray mask denotes less resolved areas as determined by inspection of checkerboard test recovery for each profile and depth slice. Several features are highlighted which are discussed in detail in the text: DLB = Delamerian-Lachlan boundary; LTZ = Lithosphere Transition Zone; NVP = Newer Volcanics Province; BS = Bass Strait; CE = Craton Edge; CP = Curnamona Province; GC = Gawler Craton. The velocity anomalies identified by the dashed lines are inferred to represent a signature at depth of what is observed at the surface.

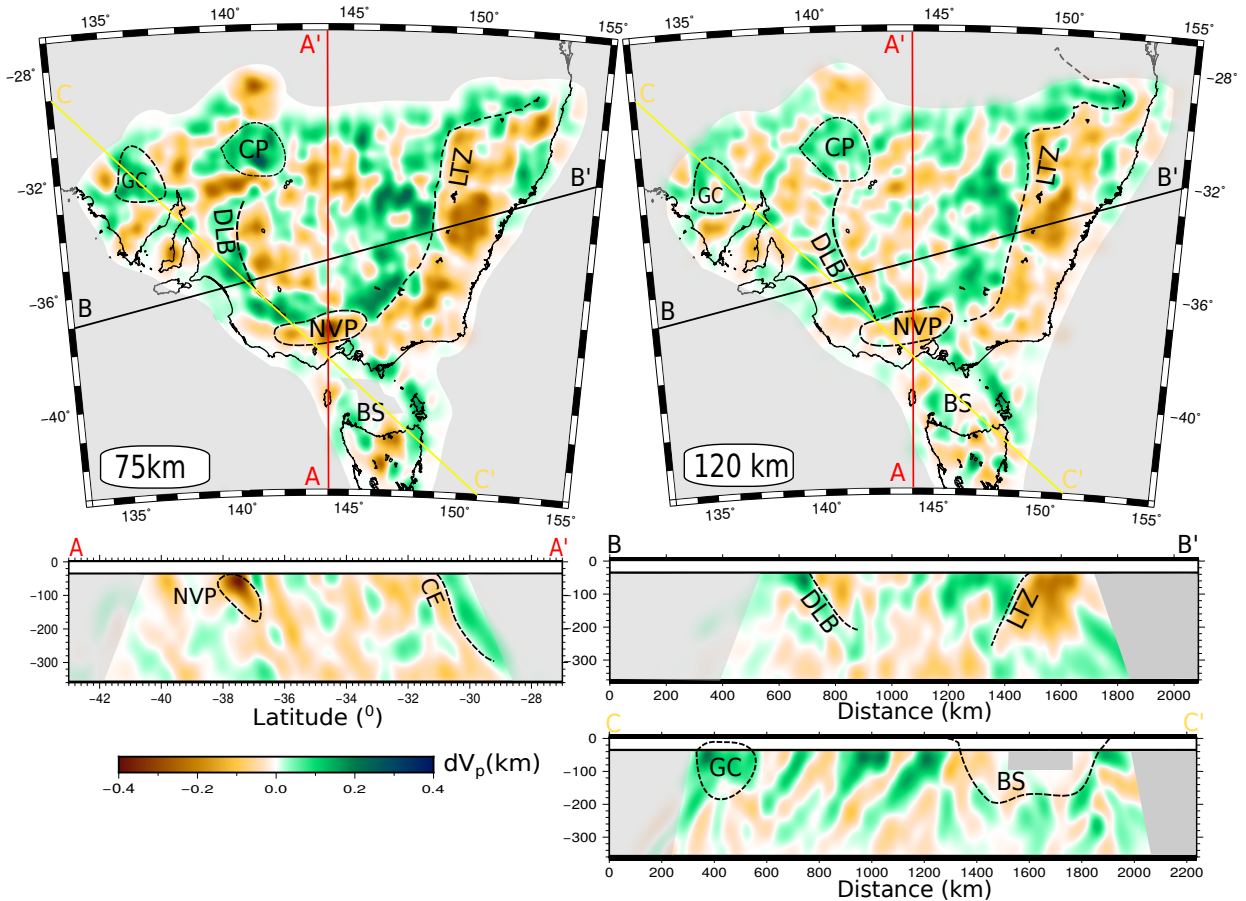


Fig. 13: Same as Figure 12 but now using a 1-D starting model (ak135) and a flat Moho for the teleseismic data inversion. The locations of the three cross sections are denoted in both depth sections as coloured lines. The gray mask denotes less resolved areas as determined by inspection of checkerboard test recovery for each profile and depth slice. Note the poor resolution in upper parts of the Bass Strait. Several features are highlighted which are discussed in detail in the text: DLB = Delamerian-Lachlan boundary; LTZ = Lithosphere Transition Zone; NVP = Newer Volcanics Province; BS = Bass Strait. CE = Craton Edge; CP = Curnamona Province; GC = Gawler Craton. The velocity anomalies identified by the dashed lines are inferred to represent a signature at depth of what is observed at the surface.