

# 1 Gondwana break-up related magmatism in the Falkland Islands

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8 Jurassic dykes (c. 182 Ma) are widespread across the Falkland Islands and exhibit considerable  
9 geochemical variability. Orthopyroxene-bearing NW-SE oriented quartz-tholeiite dykes  
10 underwent fractional crystallization > 1 GPa, and major element constraints suggest that they  
11 were derived by melting of pyroxenite-rich source. They have  $\epsilon\text{Nd}_{182}$  in the range -6 to -11 and  
12  $^{87}\text{Sr}/^{86}\text{Sr}_{182} > 0.710$  and therefore require an old lithospheric component in their source. A suite  
13 of basaltic-andesites and andesites exhibit geochemical compositions transitional between Ferrar  
14 and Karoo magma types, and are similar to those seen in the KwaZulu-Natal region of southern  
15 Africa and the Theron Mountains of Antarctica. Olivine-phyric intrusions equilibrated at < 0.5  
16 GPa, and have isotopic compositions ( $\epsilon\text{Nd}_{182}$  1.6-3.6 and  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  0.7036-0.7058) that require  
17 limited interaction with old continental lithosphere. A suite of plagioclase-phyric intrusions with  
18  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  c. 0.7035 and  $\epsilon\text{Nd}_{182}$  c. +4, and low Th/Ta and La/Ta ratios (c. 1 and c. 15  
19 respectively) also largely escaped interaction with the lithosphere. These isotopically depleted  
20 intrusions were probably emplaced synchronously with Gondwana fragmentation and the  
21 formation of new oceanic lithosphere. Estimates of mantle potential temperature from olivine  
22 equilibration temperatures do not provide unequivocal evidence for the presence of a plume  
23 thermal anomaly beneath the Falkland Islands at 182 Ma.  
24

25 The Early Jurassic (c. 180 Ma) Karoo and Ferrar large igneous provinces (LIP) were associated  
26 with Gondwana break-up. Igneous rocks of the Karoo province occur predominantly in South  
27 Africa but extend into Dronning Maud Land (Antarctica) with the main phase of activity taking  
28 place in the interval 182-183 Ma (Svensen *et al.* 2012). The Ferrar Province, which is  
29 contemporaneous with the magmatism in the Karoo province (Burgess *et al.* 2015), is typified  
30 by the low TiO<sub>2</sub> Jurassic igneous rocks of the Transantarctic Mountains and Tasmania (Hergt *et*  
31 *al.* 1989; Fleming *et al.* 1995). It has also been established that the Karoo and Ferrar provinces  
32 have areas of geographical overlap, most notably in the KwaZulu area of South Africa (Sweeney  
33 *et al.* 1994; Riley *et al.* 2006) and in the Theron Mountains of Antarctica (Brewer *et al.* 1992).  
34 In the latter, at least four suites of low TiO<sub>2</sub> igneous rocks have been recognized, and it has been  
35 suggested that there is a transition from one province to the other rather, than a strict  
36 geographical delineation between the two provinces (Brewer *et al.* 1992).

37 Elliott & Fleming (2000) argued that the principal focus of magmatism for both the Karoo  
38 and Ferrar provinces was the Weddell Triple Junction (Fig. 1) which was within the envelope of  
39 a plume-related thermal anomaly associated with Gondwana break-up. The near perfect  
40 matching of the stratigraphy, structural geology and outcrop pattern of the of the Falkland  
41 Islands with that of the eastern Karoo basin, led Adie (1952) to propose a pre-Jurassic plate  
42 reconstruction which placed the islands adjacent to the Transkei area of South Africa (Fig. 1a),  
43 forming the south-east corner of the Karoo basin with its adjoining Palaeozoic fold belts.  
44 Evidence from palaeomagnetic studies (e.g. Mitchell *et al.* 1986; Taylor & Shaw 1989) similarly  
45 place the Falkland Islands on the extension of the Cape Fold Belt of South Africa, on the eastern  
46 flank of the Lebombo Rift (Fig. 1; Macdonald *et al.* 2003; Stone *et al.* 2008; 2009; Richards *et*  
47 *al.* 2013). Post-180 Ma, there was major re-organization of crustal blocks in Patagonia, the  
48 Falklands Plateau and west Antarctica, which included the clockwise rotation of the Falklands  
49 crustal block in an overall extensional regime, the early stages of this process being dominated  
50 by strike-slip motion between the Antarctic and African plates (Taylor & Shaw 1989). The 180°

51 rotation of the Falkland Islands from their pre-180 Ma position was complete by 165 Ma  
52 (Macdonald *et al.* 2003), and by this time the islands had migrated to the west along the  
53 extension of the Aghulas Fracture zone to a position well to the west of the WTJ (Richards *et al.*  
54 2013).

55 Consequently, the Falkland Islands may have been very close to the focus of break-up related  
56 magmatism, and it is logical to assume that the geochemical composition of any igneous rocks  
57 found in the islands should reflect the diversity of magmatism in the Jurassic Gondwana LIP as  
58 a whole. In this paper, new data are presented that show that the dykes and minor intrusions of  
59 the Falkland Islands exhibit variability in mineralogy, major element, trace element and Sr-, Nd-  
60 and Pb-isotopic compositions that is nearly as large as that seen in the entire Jurassic Gondwana  
61 LIP, even though the Falkland Islands themselves represent only an extremely small area  
62 compared to the total distribution of Jurassic igneous rocks of the region. Intrusions with major  
63 and trace element characteristics most similar the Ferrar dolerites of the Transantarctic  
64 Mountain are juxtaposed with intrusions which are nearly identical to some Karoo basalts of  
65 South Africa and Antarctica.

## 66 **Falkland Islands Dyke Swarm**

67 Dolerite dykes, mostly of Jurassic age, are widespread in West Falkland and rather sparse in  
68 East Falkland (Fig. 2; Greenway 1972; Mussett & Taylor 1994; Thistlewood *et al.* 1997;  
69 Mitchell *et al.* 1999; Stone *et al.* 2008, 2009; Richards *et al.* 2013). Distinct sub-swarms of  
70 dykes have been recognized based on azimuth of exposed intrusions and aeromagnetic  
71 anomalies (Mitchell *et al.* 1999; Stone *et al.* 2009). Prominent dolerite dykes, tens of metres  
72 wide and oriented NE-SW, are present in both West and East Falkland and are reversely  
73 magnetized. This suite corresponds to the N-S suite of Mitchell *et al.* (1999), and is of Jurassic  
74 age (c. 178-190 Ma; Mussett & Taylor 1994; Stone *et al.* 2009) although the older of these ages  
75 were generated by the Ar-Ar method on whole-rock samples and have large errors (e.g R1790  
76 190±4 Ma; Mussett & Taylor 1994). E-W oriented olivine-dolerite dykes occur locally in the

77 south of West Falkland, and they form part of a larger suite of intrusions that Stone *et al.* (2009)  
78 suggest has a partially radial disposition. In addition, Richards *et al.* (2013) noted that there is a  
79 suite of about 40, N-S oriented magnetic anomalies, that may represent intrusions, and these  
80 occur across the entire Falkland Islands. Exposed examples from Teal Creek and Peat Banks  
81 (Fig. 2) yield  $^{40}\text{Ar}/^{39}\text{Ar}$  ages in the range 133-138 Ma and these dykes are likely to be members  
82 of the Etendeka suite of south-western Africa (Stone *et al.* 2009; Richards *et al.* 2013). During  
83 the current study,  $^{40}\text{Ar}/^{39}\text{Ar}$  step-heating analysis was carried-out on separated plagioclase  
84 feldspar phenocrysts from three samples, but only one of these yielded useful information.  
85 Sample WI-5, a NE-SW oriented dyke from Weddell Island (Fig. 2), which is also within the  
86 area of the radial swarm identified by Richards *et al.* (2013), contains abundant plagioclase  
87 phenocrysts, and yielded a precise age of  $182.3\pm 1.5$  Ma (Fig. 3). This confirms a Jurassic age  
88 for some of the Falkland Islands intrusions, and it is within error of the  $178.6\pm 4.9$  Ma  
89 determined by Stone *et al.* (2008) for an aphyric NE-SW dyke from Port Sussex Creek, East  
90 Falkland (Fig. 2).

91 Selected major and trace element abundances *versus* weight % MgO for 139 intrusions from  
92 the Falkland Islands, including 109 from this study and 30 from Mitchell *et al.* (1999), are  
93 shown in Fig. 4 and representative analyses are given in Table 1. Mitchell *et al.* (1999) divided  
94 the intrusions of the Falkland Islands into two main N-S and E-W suites based on azimuth, field  
95 occurrence, petrography, mineral chemistry and whole-rock geochemical data. A subsidiary  
96 three magma types were also tentatively identified by Mitchell *et al.* (1999), including ‘evolved  
97 N-S’, Lively Island and Mount Alice types. The reassessment of the spatial distribution,  
98 orientation and age of the dyke swarms by Stone *et al.* (2009) and Richards *et al.* (2013), along  
99 with the much enlarged data set for the igneous rocks of the Falkland Islands generated for this  
100 study, now allows the identification of five individual geochemical types of intrusions. The  
101 criteria used to separate the different groups of intrusions are given in Table 2 and are illustrated  
102 in Figs 4 to 9. A description of each suite is given below.

103 **Port Sussex Creek-type intrusions (PST).** All the N-S dykes of Mitchell *et al.* (1999) are  
104 included in this suite of intrusions, with the exception of the ‘evolved type’ described by  
105 Mitchell *et al.* (1999) which will be discussed under the Dyke Island Type (samples NHF17 and  
106 NGF15). PST intrusions are widely distributed across both East Falkland and West Falkland, all  
107 are sub-vertical with an azimuth of NE-SW, and they are consistently between 8 and 10 m in  
108 thickness. A typical example occurs at Port Sussex Creek, East Falkland, (MHF1; Table 1, Fig.  
109 2), and is an 8m wide, sub-vertical, medium-grained, spheroidally-weathering dolerite dyke with  
110 an azimuth of 45° (NE-SW) and an age of 178.6±4.9 Ma (Stone *et al.* 2008). The texture is  
111 equigranular and intersertal. Pyroxene is enstatite (En<sub>70</sub>Wo<sub>4</sub>Fs<sub>26</sub>), pigeonite (En<sub>51</sub>Wo<sub>13</sub>Fs<sub>36</sub>) and  
112 augite (Fig. 5) and the feldspar is labradorite (An<sub>70</sub>). All PST intrusions contain both augite and  
113 pigeonite, with more mafic samples containing orthopyroxene. Olivine (Fo<sub>50-71</sub>) is rare in this  
114 suite of rocks and is restricted to intrusions with Mg# >58 (e.g. FAR1503 and NGF16; Table 1,  
115 Fig. 4). Whole-rock MgO contents vary from 5.9-9.5 wt% (Mg# 50-62) and SiO<sub>2</sub> (53-55wt%) is  
116 higher for a given MgO concentration than any of the other Falkland Islands intrusions (Fig. 4).  
117 The PST intrusions are characterized by low CaO (8.1-9.8wt%) for a given MgO content  
118 compared to other Falkland Islands intrusions. TiO<sub>2</sub> abundances (0.9-1.2wt%) are typical of the  
119 low TiO<sub>2</sub> Gondwana break-up related LIPs of the southern hemisphere and distinguishes them  
120 from the high TiO<sub>2</sub> (>2.5wt%) suite of break-up related magmas (e.g. Brewer *et al.* 1992).  
121 Abundances of Cr are unusually high (up to 648ppm) for samples with SiO<sub>2</sub> in their range, and  
122 are reflected in the high Cr content of orthopyroxene. Abundances of Nb and Y are restricted to  
123 2-5 and 19-23ppm respectively. PST intrusions are LREE enriched (Fig. 6) with [La/Yb]<sub>N</sub> in the  
124 range 3.2-3.9 and samples lack any significant Eu anomaly (Eu/Eu\* 0.89-0.97). La/Ta and  
125 Th/Ta are the highest of any of the Falkland Islands samples analysed (44-52 and 5.9-8.6  
126 respectively), and consequently, on ORB-normalized multi-element diagrams (Fig. 7), samples  
127 exhibit a marked trough in the abundances of Ta and Nb relative to Th, U, K and La. [Ta/Yb]<sub>N</sub>  
128 is in the range 2.0 to 2.6, the lowest values for any of the Falkland Islands intrusions. Ti/Zr and

129 P/Zr (55-60 and 4.5-6.3 respectively) are such that all PST intrusions exhibit a minor trough at  
130 Ti and P relative to adjacent elements on ORB-normalized diagrams.  $\epsilon\text{Nd}_{182}$  varies from -5.5 to -  
131 11.0 and is accompanied by radiogenic Sr-isotopic compositions ( $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  0.7070-0.7134),  
132 although Sr-Nd isotope covariations are rather scattered (Fig. 8). Pb-isotopic compositions form  
133 an array that is close to the Geochron ( $^{207}\text{Pb}/^{204}\text{Pb}$  =15.55-15.65), and extends to  $^{206}\text{Pb}/^{204}\text{Pb}$   
134 ratios of up to 18.40.  $^{207}\text{Pb}/^{204}\text{Pb}$  exhibits a negative correlation with  $\epsilon\text{Nd}_{182}$  for PST intrusions  
135 (Fig. 8). Marked negative correlations between  $1/\text{Sr}$  and  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$ ,  $\epsilon\text{Nd}_{182}$  and Th/Ta and a  
136 positive correlation between MgO and  $\epsilon\text{Nd}_{182}$  (Fig. 9) suggests that PST dykes underwent  
137 interaction with a high  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  ( $> 0.714$ ), low  $\epsilon\text{Nd}_{182}$  ( $< -12$ ) component that had Th/Ta  $> 9$ ,  
138 and that interaction was concomitant with crystallization

139 ***E-W intrusions.*** The E-W samples reported by Mitchell *et al.* (1999) are from a single intrusion,  
140 approximately 10m wide, which can be traced for more than 30 km from Fox Bay West to  
141 Queen Charlotte Bay (Fig. 2). These samples are generally medium-grained olivine-phyric  
142 dolerites ( $\text{Fo}_{82}$  at 11wt% MgO in the whole-rock), the only pyroxene present being augite (Fig.  
143 5). During the current study, intrusions with similar petrographic and mineralogical  
144 characteristics were found around South Harbour and on Weddell Island (Fig. 2). E-W  
145 intrusions are distinguished from the PST (Fig. 10) by their lower  $\text{SiO}_2$  contents (48-52wt%)  
146 and higher Ti/Zr (80-95) for a similar range in  $\text{TiO}_2$ , and MgO content (1.0-1.4 and 4.8-11.4wt  
147 % respectively). E-W intrusions have  $[\text{La}/\text{Yb}]_N$  in the range 2.1-4.0 (Fig. 6) and no  
148 appreciable Eu anomaly ( $\text{Eu}/\text{Eu}^*$  0.89-1.0).  $[\text{Ta}/\text{Yb}]_N$  ratios are in the range 2.8 to 4.7, and all  
149 samples exhibit a negative Nb, Ta trough relative to the LILE (La/Ta, 16-27, Th/Ta 2.2-2.8) but  
150 this is not as pronounced as that for the PST intrusions (Fig. 7). E-W intrusions have isotopic  
151 compositions that are close to, or slightly depleted relative to the Chondritic Uniform Reservoir  
152 ( $\epsilon\text{Nd}_{182} = -0.4$  to  $3.0$ ;  $^{87}\text{Sr}/^{86}\text{Sr}_{182} = 0.7036$ - $0.7058$ ) and have Pb-isotopic compositions that lie  
153 just above the NHRL (Fig. 8).

154 **Lively Island intrusion.** A single 30m thick intrusion which is exposed on Lively Island has  
155 noticeably lower TiO<sub>2</sub> for a given MgO content than any other of the Falklands Islands  
156 intrusions (TiO<sub>2</sub> = 0.8wt% at 6wt% MgO) and the data falls close to the compositional trend for  
157 low TiO<sub>2</sub> Ferrar dolerites from the Transantarctic Mountains (Fig. 11). Characteristic  
158 mineralogical features are the presence of sparse, Mg-rich biotite and rare Ca-poor groundmass  
159 pyroxene (Fig. 5). The intrusion has a LREE-enriched REE profile ([La/Yb]<sub>N</sub> = 3.2; Fig. 6)  
160 which lacks a significant negative Eu anomaly (Eu/Eu\* = 0.87). La/Ta and Th/Ta (20.4 and 3.2  
161 respectively) are similar to E-W intrusions and considerably lower than for PST intrusions  
162 (Table 2). The Lively Island intrusion contains radiogenic Nd and unradiogenic Sr (εNd<sub>182</sub> -0.5,  
163 <sup>87</sup>Sr/<sup>86</sup>Sr<sub>182</sub> c. 0.7060) compared to PST intrusions.

164 **Dyke Island Type (DIT).** The greatest concentration of DIT intrusions is on aptly-named Dyke  
165 Island (Fig. 2). Sample WI-5 which yielded the Ar-Ar age of 182.3±1.5 Ma which crops out on  
166 Weddell Island (Fig. 2) is of this type. In addition, the evolved N-S samples described by  
167 Mitchell *et al.* (1999) are of this type (e.g. NHF17; Fig. 2). Intrusions are generally <50 cm  
168 thick, they may contain abundant plagioclase ± augite phenocrysts (samples WI-5, MHF14.9  
169 and FAR338; Fig. 5), or more commonly they are medium- to fine-grained aphyric basaltic-  
170 andesites and andesites with rare rhyolite sheets occurring locally. DIT intrusions represent an  
171 expanded fractionation series with MgO varying from 5.6 to <0.1wt%, over a range of 51-  
172 75wt% SiO<sub>2</sub>. Ti/Zr is in the range 32-55 for samples with 4.0-5.6wt% MgO, and for samples  
173 with <1wt% MgO, Ti/Zr falls to <5 (Fig. 4). All DIT intrusions have higher concentrations of  
174 the incompatible elements Zr, Nb and Y than any of the other intrusions from the Falkland  
175 Islands, and exhibit strong positive linear correlations between these elements. On a plot of TiO<sub>2</sub>  
176 *versus* MgO (Figs 4 & 11) DIT intrusions can be divided into three distinct series; i) a low TiO<sub>2</sub>  
177 series which forms and extension of the data array for PST intrusions; ii) a high TiO<sub>2</sub> series with  
178 MgO in the range 2.5-4.0wt% MgO with TiO<sub>2</sub> >1.7wt%; and iii) acid intrusions with <2wt%  
179 MgO.

180 DIT intrusions are LREE-enriched ( $[La/Yb]_N = 4.1-6.6$ ; Fig. 6) and exhibit stepwise  
181 increases in both LREE and HREE abundances with decreasing MgO with the most evolved  
182 sample (MHF41.3, 0.06wt% MgO) having  $La_N = 190$  and  $Yb_N = 44$ . The development of a  
183 progressively larger negative Eu anomaly ( $Eu/Eu^* = 0.85-0.71$ ) with decreasing MgO, attests to  
184 the importance of plagioclase fractionation during their petrogenesis. Th/Ta and La/Ta for the  
185 DIT intrusions (2.4-3.4 and 17-26 respectively) are similar to those for the E-W intrusions.  
186 Multi-element diagrams (Fig. 7) show that DIT intrusions exhibit troughs at Ti and P relative to  
187 adjacent elements, and a progressively larger negative Sr anomaly is developed with decreasing  
188 MgO. Plagioclase-phyric samples WI-5 and FAR338 exhibit a positive Sr spike in Fig. 7, which  
189 is presumably a result of accumulation of plagioclase feldspar, although neither sample exhibits  
190 a Eu anomaly. The distribution of trace elements in DIT intrusions bears a strong resemblance to  
191 those for Ferrar dolerites from the Transantarctic Mountains (Fig. 7). DIT intrusions have  
192  $^{87}Sr/^{86}Sr_{182}$  in the range 0.7055-0.7170 but all samples with  $^{87}Sr/^{86}Sr_{182} > 0.7090$  contain  $< 2$   
193 weight % MgO.  $\epsilon Nd_{182}$  falls in the range -2.8 to +0.6, and there is no systematic variation  
194 between MgO,  $\epsilon Nd_{182}$  or  $^{87}Sr/^{86}Sr_{182}$  (Figs 8 & 9).

195 **Mount Alice Type intrusions (MAT).** MAT intrusions are restricted to the south-western area of  
196 West Falkland, around South Harbour, Dyke Island and Cape Orford (Fig.1) and are early  
197 Jurassic in age ( $188 \pm 2$  Ma for sample MA3; Mussett & Taylor 1994). MAT intrusions are  
198 generally  $< 1$  m thick and are characterized by plagioclase  $\pm$  augite  $\pm$  olivine phenocrysts in a  
199 fine-grained groundmass. Since the area in which the MAT intrusions occur is within the region  
200 of the radial dyke swarm described Richards *et al.* (2013), azimuths cannot be used as one of  
201 their classification criteria, although in the region of Cape Orford, MAT dykes are generally  
202 oriented E-W (Mussett & Taylor 1994; Thistlewood *et al.* 1997). A typical example (MFH15.2)  
203 contains sparse, scattered phenocrysts of olivine ( $Fo_{80}$ ), calcic augite ( $En_{31}Fs_{25}Wo_{44}$ ; Fig. 5) and  
204 labradorite ( $An_{60}$ ). MgO varies from 6-12wt%, and the MAT intrusions have the lowest  $SiO_2$   
205 (46-50wt%) for a given MgO content of any of the Falklands Islands samples (Fig. 4).  $TiO_2$



206 abundances (1.3-2.0wt%) overlap with those for both the PST and E-W intrusions but Ti/Zr is in  
207 the range 90-150 (Fig. 10) which is considerably higher than any other of the Falkland Islands  
208 intrusions. MAT intrusions have  $[La/Yb]_N$  in the range 1.8-3.1 (Fig. 6) and flat to slightly  
209 LREE-depleted REE profiles for elements La to Sm ( $[La/Sm]_N$  0.9-1.5). On multi-element  
210 diagrams (Fig. 7), MAT intrusions exhibit a positive Sr spike relative to N-ORB, but otherwise  
211 have smooth profiles from elements Nd to Lu, with Ti/Zr and P/Zr (98-130 and 8.0-10.4  
212 respectively) in the range for normal ORB (Ti/Zr c. 100, P/Zr c. 6.9; Sun & McDonough 1987).  
213 Unlike all the other Falkland Islands intrusions, the MAT have Th/Ta and La/Ta (0.7-1.0 and  
214 13-17 respectively) which are also within the range for normal ocean ridge basalts and  
215 asthenosphere-derived basalts (Sun & McDonough 1987). Sr- & Nd-isotopic compositions fall  
216 in the upper-left quadrant of Fig. 8, with  $\epsilon Nd_{182} > 5$  and  $^{87}Sr/^{86}Sr_{182} < 0.7040$ . Pb-isotopic  
217 compositions fall just above the NHRL with  $^{206}Pb/^{204}Pb$  in the range 18.2-18.5 (Fig. 8).

## 218 **Petrogenesis of the Falkland Islands intrusions**

219 The diversity of major and trace element geochemistry and isotopic compositions of the  
220 Falkland Islands intrusions requires an equally diverse range of petrogenetic histories. In  
221 particular, the observed ranges of isotopic compositions described above are likely to require  
222 variable interaction between isotopically depleted melts from mantle peridotite with melts  
223 derived from continental lithosphere, which in some cases, must of considerable antiquity.  
224 Consequently, before attempting to make regional comparisons between the Falkland Islands  
225 intrusions and other Gondwana break-up related low  $TiO_2$  suites within the southern  
226 hemisphere, an assessment of the petrogenetic history of each of the suites of Falkland Island  
227 intrusions will be made in turn below.

### 228 ***PST intrusions***

229 PST and low  $TiO_2$  DIT intrusions exhibit variations in major element compositions that fall  
230 along the same fractionation trend as low  $TiO_2$  basaltic-andesites and andesites from the Theron

231 Mountains (Fig. 11). However, PST and DIT intrusions cannot be related to one another by  
232 simple crystal fractionation because their Sr-, Nd- and Pb-isotopic compositions differ  
233 significantly from one another (Fig. 8). The high  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  ( $>0.710$ ) and unradiogenic Nd-  
234 isotopic compositions of the PST is a feature they share with Ferrar Province igneous rocks.  
235 Fleming *et al.* (1995) and Molzhan *et al.* (1999) demonstrated that elevated  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  (0.7090-  
236 0.7112) of MFCT basalts was partly a function of Rb and Sr mobility during a Cretaceous (97-  
237 125 Ma) hydrothermal event. However, the observed range in  $\epsilon\text{Nd}_{182}$  in the same samples (-4.8  
238 to -5.6) is unlikely to be the result of alteration, and Fleming *et al.* (1995) concluded from  
239 analysis of phenocryst phases, that prior to alteration, MFCT basalts must have had  
240  $^{87}\text{Sr}/^{86}\text{Sr}_{182} \geq 0.7090$ . Alteration by the same regional hydrothermal thermal event cannot be used  
241 as an explanation for variability in the Sr-isotopic compositions of PST intrusions because the  
242 Falkland Islands would have already broken-away from the Antarctic continent by this time. In  
243 addition, the range of  $\epsilon\text{Nd}_{182}$  from -6 to -12 over the limited range of  $^{87}\text{Sr}/^{86}\text{Sr}_{182} = 0.7110$ -  
244 0.7115, for the PST intrusions, requires potential contaminants that had a range of Nd-isotopic  
245 compositions and were therefore probably of differing ages.

246 **Interaction with the continental lithosphere.** For the PST intrusions, the variations shown in  
247 Fig. 9 indicate that Sr- and Nd-isotopic variations were imposed on the magmas concomitant  
248 with fractional crystallization, by assimilation with fractional crystallization (AFC) or a similar  
249 process. The relationships shown in Fig. 9 require that Sr behaved incompatibly during  
250 fractional crystallization of the PST suite. The crystal cumulate formed during AFC cannot,  
251 therefore, have been plagioclase-rich. To generate the range of Sr-isotopic compositions seen in  
252 the PST intrusions requires a source with  $^{87}\text{Sr}/^{86}\text{Sr}_{182} \leq 0.7075$ , and contaminants with  
253  $^{87}\text{Sr}/^{86}\text{Sr}_{182} > 0.7130$  and a range of  $\epsilon\text{Nd}_{182}$ , which must be  $\leq -6.0$  for all samples. Least-squares  
254 modelling of the extract and evolved liquid from a starting composition with 9.6wt% MgO  
255 (NGF16) to evolved composition with 6.78wt% MgO (MHF5.1; Table 3) requires 21%  
256 crystallization of an assemblage of orthopyroxene (74.7%), plagioclase (18.9%) and minor

257 augite (6.4%). With only 18.9% of the fractionating assemblage being plagioclase,  $D_{Sr}$  would  
258 have been  $<1$  which is consistent with the relationship between  $1/Sr$  and  $^{87}Sr/^{86}Sr_{182}$  in Fig. 9.  
259 In addition, Demarchi *et al.* (2001) showed that orthopyroxene is on the liquidus of Ferrar  
260 tholeiites at 1.0-1.5 GPa, suggesting that magmatic differentiation of the PST intrusions  
261 occurred at depths  $\geq 30$  km.

262 Sr- and Nd-isotopic compositions for PST intrusions fall in an intermediate position between  
263 the data for CT1 basalts of Dronning Maud Land and the Karoo (Fig. 12). Luttinen & Furnes  
264 (2000) argued that the extreme Nd-isotopic compositions ( $\epsilon Nd_{182} \leq -17$ ) of CT1 basalts were the  
265 result of interaction between a mantle-derived magma and Archean (3.0 Ga) Grunehogna  
266 cratonic lithosphere (Fig. 1). Riley *et al.* (2006) used AFC and energy-constrained recharge  
267 AFC to model the isotopic compositions of Karoo basaltic rocks using an ORB-like source and  
268 an assimilant with  $\epsilon Nd_{182} = -4$  and  $^{87}Sr/^{86}Sr_{182} = 0.710$ , and showed that the observed isotopic  
269 variability in the basalts could be explained partly by these processes. In Fig. 12 three AFC  
270 trajectories are plotted and the parameters used to generate the curves are given in Table 4.  
271 These are not designed to fully explain the isotopic diversity in Gondwana low  $TiO_2$  continental  
272 flood basalts, they have been generated in an attempt to constrain possible and impossible  
273 petrogenetic processes. The starting composition has been kept constant, and is based on that of  
274 largely uncontaminated low  $TiO_2$  basalts with  $\epsilon Nd_{182} = 2$  and  $^{87}Sr/^{86}Sr_{182} = 0.7035$ . For all three  
275 modelled AFC trends, the ratio of the country rock assimilated to crystal cumulate formed,  $R$ ,  
276 has been set at 0.40, a value that is appropriate for crystallization in the middle- to upper-crust  
277 (Riley *et al.* 2006; Hole *et al.* 2015).  $D_{Sr}$  and  $D_{Nd}$  are set at 0.5 and 0.1 respectively, to simulate a  
278 cumulate with approximately 25% plagioclase, and 75% ferromagnesian minerals. This means  
279 that all three AFC trajectories approach the composition of the most contaminated magma for  $\leq$   
280 20% AFC (Table 4). Increasing the value of  $R$  to 0.5 for any of the models does not significantly  
281 change the shape of the trajectories, but decreases the amount of AFC that is needed to reach the  
282 target compositions to  $\leq 12\%$ , and conversely, decreasing  $R$  to 0.3 requires  $\leq 25\%$  AFC. For the

283 CT1 AFC model, the contaminant represents 3.0 Ga Grunehogna Craton (Fig. 1) felsic granulite,  
284 with  $\epsilon\text{Nd}_{182} = -50$  and  $^{87}\text{Sr}/^{86}\text{Sr}_{182} = 0.712$  (felsic xenolith sample X4-AVL of Luttinen & Furnes  
285 2000). The PST-1, mixing line intersects the lowest  $\epsilon\text{Nd}_{182}$  samples in the PST suite, and the  
286 contaminant represents a 2.2 Ga Palaeoproterozoic felsic granulite with  $\epsilon\text{Nd}_{182} = -20$  and  
287  $^{87}\text{Sr}/^{86}\text{Sr}_{182} = 0.720$  (Luttinen & Furnes 2000). The PST-2 mixing line, which also intersects the  
288 majority of data for Karoo basalts and lowest  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  ( $\sim 0.7090$ ) samples of Ferrar igneous  
289 rocks, representing mixing between a mantle-derived magma and 1.0-1.5 Ga felsic crust with  
290  $\epsilon\text{Nd}_{182} = -10$  and is the same contaminant as that suggested by Riley *et al.* (2006) for Karoo  
291 basalts. Plate reconstructions place the Falkland Islands mainly within the 1.0-1.5 Ga  
292 Namaqualand-Natal-Maudheim-Mozambique belt (Thistlewood *et al.* 1997) and on the  
293 continuation of the Cape Fold Belt (Fig. 1). Mesoproterozoic crust is therefore a likely candidate  
294 for basement to the Falkland Islands, although there are no isotopic data for the Cape Meredith  
295 metamorphic complex. What is also clear is that cratonic basement like that involved in the  
296 petrogenesis of the CT1 basalts affected neither the PST intrusions nor Karoo low  $\text{TiO}_2$  basaltic  
297 rocks. AFC models with geologically reasonable parameters and appropriate ages of potential  
298 basement contaminants can therefore produce the observed variations in the Sr- and Nd-isotopic  
299 characteristics of the PST intrusions for  $<20\%$  AFC.

300 **Pyroxenite versus peridotite sources.** PST intrusions with  $\text{MgO} > 8$  weight % have lower CaO  
301 abundances (c. 8.5 weight %) than any other of the other Falkland Islands intrusions (Figs. 4 and  
302 13). Such compositions are uncommon in continental flood basalts provinces. Orthopyroxene  
303 was the dominant fractionating phase during crystallization of the PST (Table 3) and estimates  
304 of more primitive compositions can be calculated by incrementally adding enstatite to a mafic  
305 PST composition. Addition of 30% enstatite to sample NGF16 yields magma with  $\sim 15\text{wt}\%$   
306  $\text{MgO}$  and  $\sim 7.5\text{wt}\%$   $\text{CaO}$ . Compositions such as these are also found in the CT1 basalts of  
307 Dronning Maud Land (Fig. 13a). An unusual feature of the PST intrusions is their Si-

308 oversaturated nature and high Cr content (Fig. 13b) which is also reflected in unusually high Cr  
309 content of component orthopyroxene (e.g. enstatite in MHF3.2 has 0.74wt% Cr<sub>2</sub>O<sub>3</sub> at  
310 MgO/FeO = 2.8). In terms of major element compositions, PST intrusions bear strong  
311 similarities with magnesian andesite from continental subduction settings (e.g. Baker *et al.*  
312 1994; Sato *et al.* 2014). For example, high-Mg andesites from Mt Shasta have a similar range of  
313 MgO to PST intrusions (Fig. 13a) which is accompanied by SiO<sub>2</sub> = 51.5-54.0wt%, Cr = 245-695  
314 ppm, Ni = 99-235 ppm, TiO<sub>2</sub> = 0.6-0.8wt% and CaO = 8.6-9.6wt%. One mechanism that has  
315 been suggested for the production of high-Mg andesite is the interaction of slab-derived adakitic  
316 melts with mantle peridotite during subduction (Heinonen *et al.* 2014). A link to the previous  
317 subduction history of the mantle source from which Ferrar and Karoo basaltic rocks were  
318 derived has been made by a number of workers (e.g. Brewer *et al.* 1992; Storey *et al.* 1992;  
319 Heinonen *et al.* 2014) and in particular, the characteristic trough at Nb and Ta relative to  
320 adjacent elements (Fig. 7) has been interpreted as an inherited subduction signature.

321 Herzberg & Asimow (2008) note that primary magmas derived from the melting of  
322 pyroxenite will exhibit relative CaO depletion compared to melts from a peridotite source  
323 because of the dominance of residual clinopyroxene in the source region during partial melting  
324 of pyroxenite. Given the position that data for the PST occupy in Fig. 12, it seems clear that  
325 their major element compositions are not consistent with an origin by melting of mantle  
326 peridotite. It is well established that pyroxenite can be formed at the base of the lithosphere as a  
327 result of accumulation of mafic phases during basaltic magmatism (e.g. Downes *et al.* 2007).  
328 Such accumulative pyroxenite can yield magma by partial melting at some later stage, promoted  
329 by a new phase of mafic magmatism and by interaction with peridotite-derived melts (Lambart  
330 *et al.* 2013). The generation of silica-enriched pyroxenite melts is possible, which can yield Si-  
331 oversaturated melts like those of the PST intrusions (Lambart *et al.* 2013). It is therefore  
332 suggested that the PST were derived from a pyroxenite-rich source that was emplaced at the  
333 base of the lithosphere during the prolonged subduction history of Gondwana. Metasomatism of

334 the pyroxenite by slab-derived fluids and melt, imparted a subduction signature to the  
335 pyroxenite. When subjected to the high mantle potential temperatures associated with the mantle  
336 plume beneath Dronning Maud Land at c. 180 Ma ( $T_P$  up to 1600°C; Heinonen *et al.* 2010), the  
337 pyroxenite underwent partial melting and produced the primary melt precursor to the PST  
338 intrusions. These melts then interacted with fusible, felsic continental crust to produce the  
339 geochemical composition of the more evolved PST compositions by AFC, or a related process.  
340 Extrapolation of the MgO -  $\epsilon Nd_{182}$  trend for the PST to higher MgO contents (Fig. 9a), suggests  
341 that a primitive composition with 15 weight % MgO might have had  $\epsilon Nd_{182} \sim 0$ , and the  
342 correlation between  $1/Sr$  and Sr-isotopic compositions requires the source to have  $^{87}Sr/^{86}Sr_{182} \leq$   
343 0.7075.

#### 344 ***E-W intrusions***

345 Least-squares modelling of the extract and evolved liquid from a starting composition ECF12 to  
346 more evolved sample ECF44 (Table 3) requires crystal fractionation of 57% olivine and 40%  
347 plagioclase feldspar with little contribution from augite (c. 2.6 %) which is a typical  
348 crystallizing assemblage for tholeiitic melts at pressures  $\leq 0.5$  GPa (e.g. Hole & Morrison 1992;  
349 Villiger *et al.* (2007), requiring E-W intrusions to have last equilibrated at  $\sim 15$  km depth,  
350 within the crust. This is in contrast to the  $>1.0$  GPa equilibration required by PST intrusions.  
351 The isotopic compositions of E-W intrusions require derivation from a depleted mantle source  
352 (Figs 8), but they have higher Th/Ta (2.2-2.8) and La/Ta (16.4-26.1) than would be expected if  
353 they were derived from asthenospheric mantle (Th/Ta  $\sim 1.0$  and La/Ta  $\leq 18$ ; Sun & McDonough  
354 1989) and they also exhibit a noticeable trough at Nb and Ta relative to adjacent elements in  
355 Fig. 7, a feature that is most often attributed to interaction with continental lithosphere.  
356 However, the extent of this interaction must either have been limited, or the source from which  
357 the E-W intrusions were derived had a Th/Ta  $> 2.0$  and La/Ta  $> 26$ . The low pressure  
358 equilibration of E-W intrusions, coupled with their depleted isotopic compositions may suggest

359 that they were emplaced during a period of crustal attenuation and were thus able to escape  
360 interaction with continental lithosphere.

### 361 ***DIT and Lively Island intrusions***

362 In contrast to the PST intrusions, the sub-horizontal arrays delineated by DIT intrusions in Fig.  
363 9a, suggests that AFC or a similar process was not important during their petrogenesis.  
364 However, a negative correlation between Th/Ta and  $\epsilon\text{Nd}_{182}$  for the DIT intrusions (Fig. 9b) may  
365 require minor modification by a crustal component with  $\text{Th}/\text{Ta} \geq 3.0$ . A characteristic feature of  
366 the DIT samples is that they have  $\epsilon\text{Nd}_{182}$  in range -2.8 to +0.6, but with only a single analysed  
367 sample (NHF17) having  $\epsilon\text{Nd}_{182} < -1$ . In addition, the Lively Island dyke, which falls close to the  
368 fractionation trend for the MFCT basaltic rocks of the Transantarctic Mountains (Fig. 11), has  
369 Sr- and Nd-isotopic compositions ( $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  c. 0.7052,  $\epsilon\text{Nd}_{182} = -0.5$  to  $-1.4$ ) that do not  
370 require the significant isotopic enrichment seen in the Ferrar dolerites ( $\epsilon\text{Nd}_{182}$  in the range  $-3.3$   
371 to  $-5.3$ ; Fleming *et al.* 1995; Hergt *et al.* 1989). The source of the low  $\text{TiO}_2$  DIT magmas could,  
372 therefore, have had  $\epsilon\text{Nd}_{182} > 0$ ,  $^{87}\text{Sr}/^{86}\text{Sr}_{182} < 0.7050$ ,  $\text{Th}/\text{Ta} < 2.5$  and  $\text{La}/\text{Ta} < 20$ .

### 373 ***MAT intrusions***

374 The positive  $\epsilon\text{Nd}_{182}$  (2.7-3.6) and low Th/Ta, La/Ta and  $[\text{La}/\text{Yb}]_N$  (0.8-1.0; 12.8-17.3 and 1.9-  
375 3.7 respectively) of MAT intrusions suggests that they were derived from an asthenospheric  
376 source, and escaped significant interaction with lithosphere. The most satisfactory explanation  
377 for the geochemical compositions of these rocks is that they were generated by decompression  
378 melting of the asthenosphere during the rifting stage of Gondwana break-up. In this respect they  
379 have similar geochemical compositions to the ORB-like Rooi Rand basaltic dykes of the  
380 southern Lebombo (Cox & Bristow 1984, which post-date the main magmatic phases in the  
381 region by about 5 Myr (Jourdan *et al.* 2007).

### 382 ***Cretaceous intrusions***

383 Until more data are forthcoming, the origin and affinity of the Cretaceous Teal Creek intrusion  
384 reported by Stone *et al.* (2009) remains somewhat obscure. Major element data for the intrusion  
385 plot close to the Theron Mountains low TiO<sub>2</sub> trend in Fig. 6, but the intrusion has higher Fe<sub>2</sub>O<sub>3</sub>  
386 (c. 15.9wt%) at 5.7wt% MgO than any of the data for the intrusions presented here. What is  
387 clear, is that there is an extensive suite of low TiO<sub>2</sub> basalts within the Etendeka Province (e.g.  
388 Gibson *et al.* 2005; Thompson *et al.* 2001) from which it could be related. However, none of the  
389 groups of intrusions described here carries a similar signature to that presented by Stone *et al.*  
390 (2009) for the Teal Creak dyke.

### 391 **Provinciality and chemical affinities of Falkland Islands intrusions**

392 A number of authors have noted that there are spatially constrained variations in major and  
393 trace element compositions within the Jurassic Gondwana break-up related flood basalts  
394 provinces of the Southern Hemisphere (e.g. Brewer *et al.* 1992; Luttinen & Furnes 2000; Riley  
395 *et al.* 2006). Figs 11 and 14 illustrate the variability in abundances of MgO, TiO<sub>2</sub>, SiO<sub>2</sub> and  
396 Ti/Zr for Falkland Island intrusions, along with igneous rocks which are defined as being either  
397 Karoo or Ferrar magma types, or those which are considered to be transitional between the two  
398 magma types (Brewer *et al.* 1992; Luttinen & Furnes 2000; Riley *et al.* 2006). In the following  
399 sections, we will examine the geochemical affinities of the Falklands Islands intrusions in  
400 relationship to other early Jurassic flood basalts of Gondwana.

#### 401 ***PST intrusions***

402 PST intrusions exhibit strong similarities to CT1 basalts of Dronning Maud Land (e.g. Fig. 14),  
403 and show some overlap with the compositional field for samples from the Theron Mountains  
404 and Transantarctic Mountains. In terms of incompatible trace elements, PST dykes exhibit  
405 almost identical multi-elements patterns to sample SA.6.1 (Fig. 7; Riley *et al.* 2006) from  
406 KwaZulu-Natal, with which they share also unradiogenic Nd-isotopic compositions (e.g. SA.6.1  
407  $\epsilon\text{Nd}_{182} = -8.9$ ). Within the CT1 basalts of Dronning Maud Land, very similar compositions to  
408 PST intrusions can be found (e.g. B70-AVL; Luttinen & Furnes 2000) again with many basalts



409 in the CT1 suite having unradiogenic Nd-isotopic compositions (Fig. 12). Furthermore, both  
410 CT1 and PST magmas required derivation from a pyroxenite-rich, CaO deficient, source region  
411 (Fig. 13). It is with some confidence that we conclude that the PST, CT1 and some KwaZulu-  
412 Natal basalts were derived from very similar source regions, had similar petrogenetic histories  
413 and represent the same phase of pre-break-up magmatism.

#### 414 ***E-W intrusions***

415 These igneous rocks exhibit a strong geochemical affinity with basalts from Kirwanveggan  
416 (Harris *et al.* 1990) and Schirmaker Oasis (Sushchevskaya *et al.* 2009), Dronning Maud Land  
417 (Figs 11 and 14). A notable characteristic of all the above samples is that they have radiogenic  
418 Nd-isotopic compositions ( $\epsilon_{\text{Nd}_{182}} = 2-6$ )  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  in the restricted range 0.7040-0.7060 and  
419  $^{206}\text{Pb}/^{204}\text{Pb}$  in the range 17.97-18.49 and plot close to the NHRL in Fig. 8. These isotopic  
420 characteristics, coupled with  $\text{Ti}/\text{Zr} = 70-90$  and  $\text{SiO}_2 = 47.0-52.8\text{wt}\%$ , clearly separates  
421 Kirwanveggan, Schirmaker Oasis and Falkand Islands E-W intrusions from PST intrusions. E-  
422 W intrusions also fall within the compositional field for basalts from the Central Karoo of South  
423 Africa in Fig. 14, and intrusions with compositions similar to those of Falkland Islands E-W  
424 intrusions are prevalent in the Golden Valley Sill Complex (Neumann *et al.* 2011).

#### 425 ***DIT and Lively Island intrusions***

426 In Fig. 14b, DIT intrusions exhibit almost complete overlap with the MFCT Ferrar dolerites of  
427 the Transantarctic Mountains. However, consideration of Fig. 11, shows that DIT intrusions are  
428 not of the same low  $\text{TiO}_2$  (<0.5-1.2wt%) lineage of the MFCT. However, DIT intrusions do  
429 have compositions that overlap with those for low  $\text{TiO}_2$  Theron Mountains basalts and samples  
430 from KwaZulul-Natal that Riley *et al.* (2006) and Brewer *et al.* (1992) argued were transitional  
431 between Ferrar and Karoo magma types. We concur with this hypothesis, and given the position  
432 that the Falkland Islands occupied in southern Gondwana at the time of Karoo and Ferrar  
433 magmatism, this seems entirely reasonable.

434 The Lively Island intrusion is the only member of the intrusive suite of rocks of the Falkland  
435 Islands that falls within the compositional range for MFCT samples from the Transantarctic  
436 Mountains in Figs 11 and 14. However, the Lively Island intrusion has considerably lower  
437 La/Ta and Th/Ta (20.5 and 3.2 respectively) than the majority of the MFCT dolerites (La/Ta 19-  
438 47; Th/Ta 4.3-23.3), and the Lively Island dyke also has considerably more radiogenic Nd  
439 ( $\epsilon\text{Nd}_{182} = -0.5$  to  $-1.4$ ) than MFCT dolerites ( $\epsilon\text{Nd}_{182} = -4.7$  to  $-5.7$ ). It is important to note, that  
440 the Lively Island intrusion is a single body of igneous rock around 30m thick, and consequently  
441 is not a volumetrically significant part of the Falkland Islands dyke swarm.

#### 442 ***MAT intrusions***

443 Ti/Zr $>90$  coupled with SiO<sub>2</sub> of 45.7-51.0wt% are characteristics that MAT intrusions share  
444 with CT2 and CT3 basalts from Dronning Maud Land and samples from the Rooi Rand dyke  
445 swarm of the southern Lebombo area of southern Africa (Cox & Bristow 1984).. The  
446 unradiogenic Sr-isotopic compositions ( $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  c. 0.7040) and radiogenic Nd-isotopic  
447 compositions ( $\epsilon\text{Nd}_{182} = 2.5$ -4.0) of E-W intrusions also characterize Rooi Rand and CT2 and  
448 CT3 basalts, although the Dronning Maud Land samples have a rather more extended range of  
449 isotopic compositions, which Luttinen & Furnes (2000) attribute to conservative amounts of  
450 interaction with upper-crustal felsic contaminants.

#### 451 **Mantle potential temperature, rifting and magmatism**

452 Fig. 15a summarizes the available data for olivine equilibration temperatures ( $T_{\text{OL}}$ ) for MAT  
453 and E-W basalts and picrites from Dronning Maud. MAT and E-W basalts yield olivine  
454 equilibration temperatures of 1245°C and 1330°C respectively, using the method of Putirka *et*  
455 *al.* (2007), whilst olivine in picrites from Dronning Maud Land yield  $T_{\text{OL}}$  up to 1450°C.  
456 Converting equilibration temperatures to  $T_{\text{P}}$  is problematical if the pressure and extent of  
457 melting cannot be independently determined (Herzberg & Asomow 2008; Herzberg & Gazel  
458 2009; Hole 2015), which they cannot for the MAT and E-W samples. However, since olivine  
459 equilibration temperature increases with increasing pressure of crystallization, synthetic olivine

460 liquidi can be calculated for any given temperature and pressure (Herzberg & Gazel 2009). Fig.  
461 15b shows the inferred temperature-pressure conditions at which fractional melting terminated  
462 for calculated primary magmas from Dronning Maud Land, the Karoo Province of southern  
463 Africa, Ferrar dolerites of Antarctica. (Hole 2015). In Fig. 15c, data for basalts from the  
464 Cretaceous Etendeka Province of SW Africa (Kieding *et al.* 2011) are given, for which  
465 estimates of  $T_{OL}$ , estimates of  $T_P$  from melt inclusions in ultra-magnesian olivine, and estimates  
466 of  $T_P$  from the PRIMEL2 model of Herzberg & Asimow (2008) have all been calculated on the  
467 same samples. Using the Herzberg & Asimow (2008) model yields  $T_P = 1500-1550^\circ\text{C}$  and final  
468 pressures of melting ( $P_f$ ) between 1.5 and 4.0 GPa (Fig. 13c).  $T_P$  from melt inclusions is 1300-  
469 1520 $^\circ\text{C}$ , whilst  $T_{OL}$  is in the range 1250-1400 $^\circ\text{C}$  and there is an empirical relationship between  
470  $T_{OL}$  and melt inclusion  $T_P$  which approximates to  $T_P = 1.443 \times T_{OL} - 501$  for the Etendeka plume  
471 system (Fig. 15). Therefore it seems that within a single plume system, basalts may be  
472 generated over ranges of  $T_P$  that are larger than the  $\pm 50^\circ\text{C}$  error inherent in the calculation  
473 methods (Herzberg & Asimow 2009; Hole 2015). Direct application of this empirical  
474 observation to the Dronning Maud Land picrites suggest maximum  $T_P \sim 1550^\circ\text{C}$ , a temperature  
475 that is considered to be associated with ‘hot’ mantle plumes such as Iceland at 60 Ma (Fig. 15;  
476 Herzberg & Gazel 2009). For Falkland Islands E-W basalts,  $T_{OL} \sim 1330^\circ\text{C}$ , which implies  
477  $T_P \sim 1400^\circ\text{C}$  and for olivine-phyric MAT basalts,  $T_{OL} \sim 1250^\circ\text{C}$  implying  $T_P \sim 1300^\circ\text{C}$ . These  $T_P$   
478 estimates for the Falkland Islands E-W and MAT basalts may therefore be reconciled with a  
479 model involving melting of mantle with near-ambient temperature ( $T_P \geq 1350^\circ\text{C}$ ), but would  
480 require intersection of the dry peridotite solidus at  $\sim 2.1$  GPa ( $\sim 70\text{km}$ ) and all melting to take  
481 place in the spinel stability field of the mantle; the most mafic MAT and E-W intrusions have  
482  $[\text{La}/\text{Yb}]_N < 2.0$  which does not preclude such an origin. Additionally, near-ambient  $T_P$  melting  
483 would require the continental lithosphere to be thinned substantially and perhaps to  $< 50\text{km}$ , to  
484 allow decompression melting to take place. The depleted isotopic compositions of the E-W and  
485 MAT intrusions, along with the  $< 0.5$  GPa equilibration of the E-W magmas, provides additional

486 evidence to suggest that these intrusions were emplaced during a period of crustal stretching,  
487 possible coeval with the initiation of Gondwana. Nevertheless, whilst there is no primary  
488 evidence to suggest  $T_P$  was  $>1450^\circ\text{C}$  beneath the Falkland Islands at 180 Ma, it is possible that  
489 high-MgO large melt fractions requiring substantially higher  $T_P$  exist in the region, but have not  
490 been sampled, remains a possibility.

491 The diversity of magma types found in the Falkland Islands, and the position in Gondwana  
492 which the islands occupied during magmatism (Fig. 1) is entirely consistent with their being  
493 close to the focus of magmatism during continental break-up. We concur with Brewer *et al.*  
494 (1992) and Riley *et al.* (2006) that there is considerable overlap in the geographical distribution  
495 of the Ferrar and Karoo LIPs, which is most obvious in the Theron Mountains and Falkland  
496 Islands. It is also clear, that despite the wealth of geochemical data available for the  
497 Transantarctic Mountains and Tasmania, there is no evidence to suggest that volcanic rocks with  
498 affinities to the Karoo LIP occur in those areas. With the exception of the ORB-like basalts of  
499 the Rooi Rand dyke swarm (Marsh *et al.* 1997; Mitchell *et al.* 1999) which are likely to  
500 represent syn-break-up magmas, basaltic rocks with Karoo-type geochemical compositions only  
501 extend as far south as the overlap zone in the Theron Mountains.

## 502 **Conclusions**

503 The Jurassic (c. 182 Ma) intrusions of the Falkland Islands exhibit a broad range of geochemical  
504 compositions and at least four main petrogenetic lineages are recognized. PST intrusions were  
505 derived by melting of an isotopically-enriched pyroxenite-rich source, followed by  
506 orthopyroxene-dominated crystal fractionation at  $\geq 1$  GPa. Pyroxenite-derived PST magmas  
507 subsequently interacted with 'old' ( $\geq 2.2$  Ga) fusible continental lithospheric components by  
508 AFC or a related process. The geochemical compositions of DIT intrusions bear striking  
509 similarities to igneous rocks of Kwazulu-Natal and the Theron Mountains, which are considered  
510 to be transitional in composition between those of the Ferrar and Karoo magma types. A  
511 significant number of mafic (Mg# 50-62) E-W and MAT intrusions possess radiogenic Nd- and

512 unradiogenic Sr-isotopic compositions ( $^{87}\text{Sr}/^{86}\text{Sr}_{182} < 0.7050$  and  $\epsilon\text{Nd}_{182} > 2.5$ ), also have Th/Ta  
513 and La/Ta ( $< 3.0$  and  $< 25$  respectively) that require little input from the continental lithosphere.  
514 In addition, E-W intrusions carry mineralogical and chemical fingerprints of equilibration at  $<$   
515  $0.5$  GPa. E-W and MAT basalts were probably emplaced during rifting and continental break-up  
516 and are likened to the Rooi Rand dykes of the Southern Lebombo of Africa. However, there is  
517 currently no evidence to suggest that the Falkland Islands intrusions were derived by melting  
518 above a significant mantle thermal anomaly. Early Jurassic plate reconstructions place the  
519 Falkland Islands close to the Weddell Triple Junction, perhaps explaining the diversity of  
520 igneous rock compositions found in a relatively limited geographical region.

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525

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- 714

715 **Figure Captions**

716 Figure 1. a) Reconstruction of Southern Gondwana showing the position of the Falkland Islands  
717 relative to south-eastern Africa prior to continental break-up. The three arrows represent the  
718 main dyke trends on the Falkland Islands rotated back to their pre-180 Ma orientation. After  
719 Trewin *et al.* (2002). b) Southern Gondwana, in the Middle Jurassic, showing the distribution of  
720 Jurassic break-up related magmas in southern Africa and Antarctica. MEB, Maurice Ewing  
721 Bank; EWM, Ellsworth-Whitmore Mountains; AP, Antarctic Peninsula; SA, South Africa;  
722 SAM, South America; ANT, Antarctica. Position of the Weddell, Limpopo and Lower Zambesi  
723 triple junctions are from Elliot & Fleming (2000). Ar-Ar ages, this study and Stone *et al.* (2009).  
724 After Macdonald *et al.* (2003). Key to shading for Fig. 1b is the same as that for Fig. 1a.

725 Figure 2. Map of the Falkland Islands showing the distribution of magnetic anomalies and main  
726 trends of dyke swarms. Solid or pecked lines do not necessarily represent continuous exposure  
727 of dykes. Inset; azimuths of Dykes in the South Harbour area of West Falkland. The rectangle at  
728 South Harbour is the area covered by the map in the supplementary material, which gives the  
729 sample locations and geochemical type to which the dykes belong in that area. After Stone *et al.*  
730 *et al.* (2009) and Richards *et al.* (2013). Ar-Ar ages (this study, Stone *et al.* 2008; 2009) and  
731 sample locations which are mentioned in the text, are indicated, along with the geochemical  
732 group to which the intrusions belong, given by the following abbreviations; PST, Port Sussex  
733 Type; MAT, Mount Alice Type, E-W, East-West Type of Mitchell *et al.* (1999); DIT, Dyke  
734 Island Type. Identifying characteristics of each type of intrusion are discussed in detail in the  
735 text.

736 Figure 3. Ar-Ar step-heating spectrum for plagioclase in sample WI-5. Full data are given in the  
737 supplementary materials.

738 Figure 4. Major (wt%) and trace element (in ppm) variations *versus* MgOwt% in Falkland  
739 Islands dykes. Filled dots, Port Sussex Creek type (PST) NE-SW two-pyroxene dolerites; open  
740 triangles, E-W olivine dolerite dykes; open squares, Lively Island dyke; filled squares, Mount  
741 Alice-type (MAT) dykes; open dots, low TiO<sub>2</sub> DIT intrusions; filled diamonds, high TiO<sub>2</sub> DIT  
742 intrusions; open diamonds, evolved sheets from the South Harbour-Dyke Island transect (Dyke  
743 Island Type; DIT); crosses, Pony's Pass N-S Cretaceous dyke (Stone *et al.* 2008). Data from this  
744 study, Mitchell *et al.* (1999) and Thistlewood *et al.* (1997).

745 Figure 5. Pyroxene end-member compositions represented in the quadrilateral system Enstatite -  
746 Ferrosilite – Wollastonite for Falkland Islands intrusions (this study and Mitchell *et al.* 1999)  
747 and dolerites from the Transantarctic Mountains (Elliot 1995; Demarchi *et al.* 2001). MFCT,

748 Mount Fazio Chemical Type; SPCT, Scarab Peak Chemical Type; NVL, Northern Victoria  
749 Land.

750 Figure 6. Chondrite-normalized REE profiles for representative samples of a) DIT intrusions  
751 and b) PST, MAT and E-W intrusions.

752 Figure 7. a) to d) Multi-element ORB-normalized (Sun & McDonough 1989) variation diagrams  
753 for Falkland Islands dykes. Comparable basalts from other regions of the low TiO<sub>2</sub> Gondwana  
754 LIP are shown by grey lines. Sample SA.6.1 (South Africa), Riley *et al.* (2006); VF111-85, CT3  
755 basalt, Dronning Maud Land (Luttinen & Furnes 2000); 47206-3, low TiO<sub>2</sub> tholeiite from  
756 Schirmacher Oasis, Dronning Maud Land (Sushchevskaya *et al.* 2009); Average MFCT from  
757 Elliot *et al.* (1995).

758 Figure 8. a)  $\epsilon\text{Nd}_{182}$  versus  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$ ; b)  $^{207}\text{Pb}/^{204}\text{Pb}$  versus  $^{206}\text{Pb}/^{204}\text{Pb}$  for Falkland Islands  
759 dykes. c)  $\epsilon\text{Nd}_{182}$  versus  $^{207}\text{Pb}/^{204}\text{Pb}$  for Falkland Islands intrusions. Symbols as for Fig. 4. Data  
760 sources this study, Mitchell *et al.* (1999) and Thistlewood *et al.* (1997).

761 Figure. 9 a)  $\epsilon\text{Nd}_{182}$  versus MgO; b)  $\epsilon\text{Nd}_{182}$  versus Th/Ta and c)  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  versus 1/Sr for  
762 Falkland Islands intrusions. Symbols as for Fig. 4 except grey dots are for the lowest reported  
763 Th/Ta for Ferrar dolerites (Fleming *et al.* 1995). Parameters for the AFC mixing line are given  
764 in Table 4 with % AFC given on the mixing line.

765 Figure 10. Ti/Zr versus SiO<sub>2</sub> for Falkland Islands dykes. Symbols as for Fig. 4.

766 Figure 11. TiO<sub>2</sub> versus MgO, for Falkland Islands dykes the Ferrar LIP and igneous rocks  
767 considered to be transitional between the compositions of Ferrar and Karoo magmas. Data  
768 sources; Hergt *et al.* (1989), Brewer *et al.* (1992), Elliot *et al.* (1995), Fleming *et al.* (1995),  
769 Molzahn *et al.* (1996), Wilhelm & Worner (1996), Antonini *et al.* (1999), Elliot *et al.* (1999),  
770 Elliot & Fleming (2004), Riley *et al.* (2006). Falkland Islands samples symbols as for Fig. 4.

771 Fig. 12.  $\epsilon\text{Nd}_{182}$  versus  $^{87}\text{Sr}/^{86}\text{Sr}_{182}$  for Falkland Islands PST intrusions (filled dots) Karoo low  
772 TiO<sub>2</sub> volcanic rocks (open circles), Dronning Maud Land CT1 (open triangles), CT2 (filled  
773 diamonds) and CT3 (filled triangles) basalts. Details of the parameters used in generating the  
774 three AFC mixing lines (CT1, PST-1 and PST-2) are given in Table 4. Each cross represents 1%  
775 AFC. Data sources for Karoo Province; Galerne *et al.* (2008); McClintock *et al.* (2008);  
776 Neumann *et al.* (2011).

777 Figure 13. a) CaO versus MgO (weight %) for Falkland Islands intrusions (black dots PST; grey  
778 dots, DIT; grey squares MAT; triangles, E-W) and Dronning Maud Land high MgO, silica-

779 oversaturated CT1 basalts (circles). The dividing line between melts derived from peridotite  
780 and pyroxenite sources is taken from Herzberg & Asimow (2008). Lines with crosses and  
781 arrows represent the effect of accumulation of the phase indicated on the composition of PST  
782 basalt NEF9, with each cross representing 5% accumulation. b) Cr (ppm) *versus* SiO<sub>2</sub> for  
783 Falkland Islands intrusions (symbols as for Fig. 4) and high-Mg andesites from Mt Shasta  
784 (crosses; Baker *et al.* 1994).

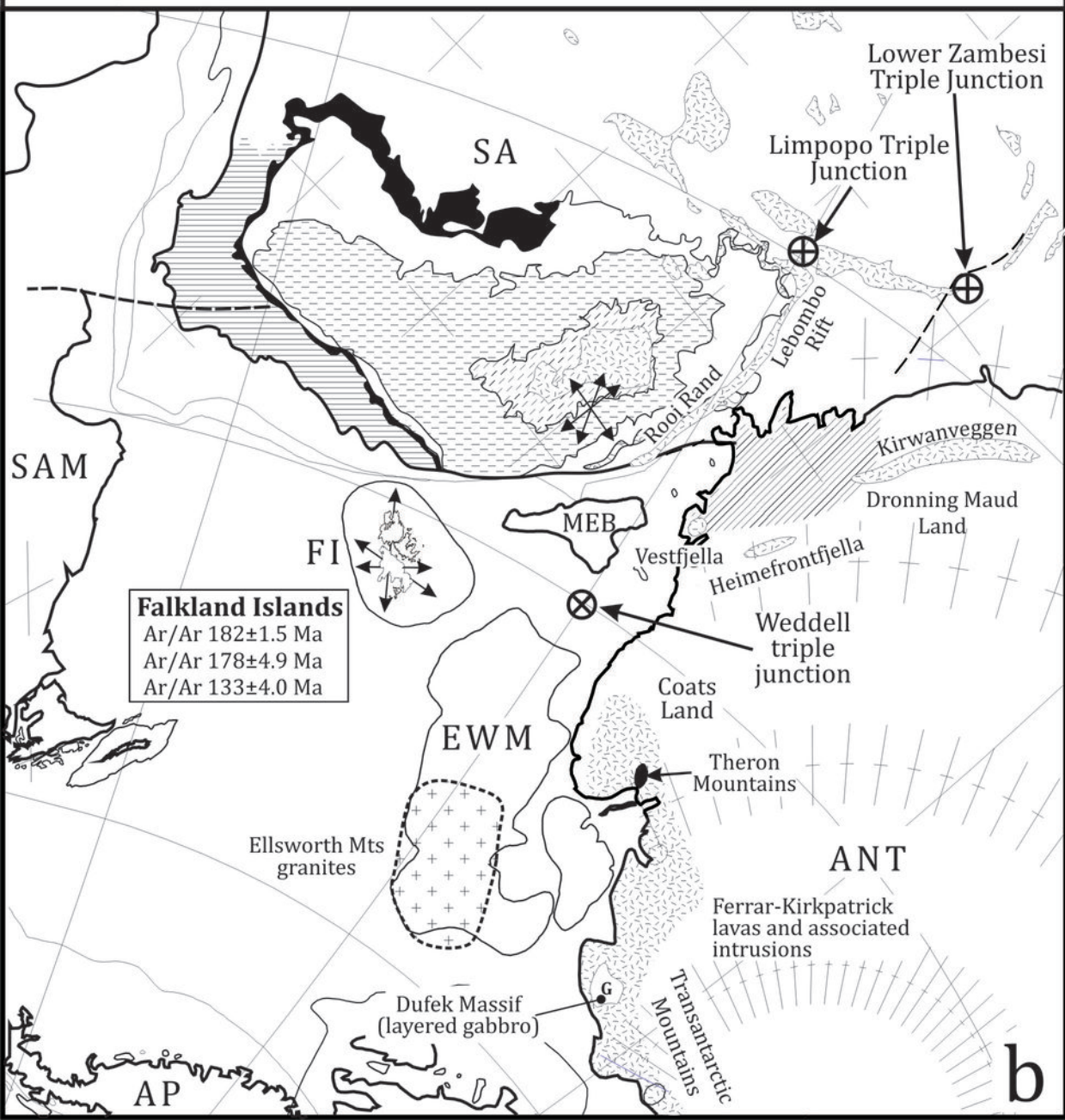
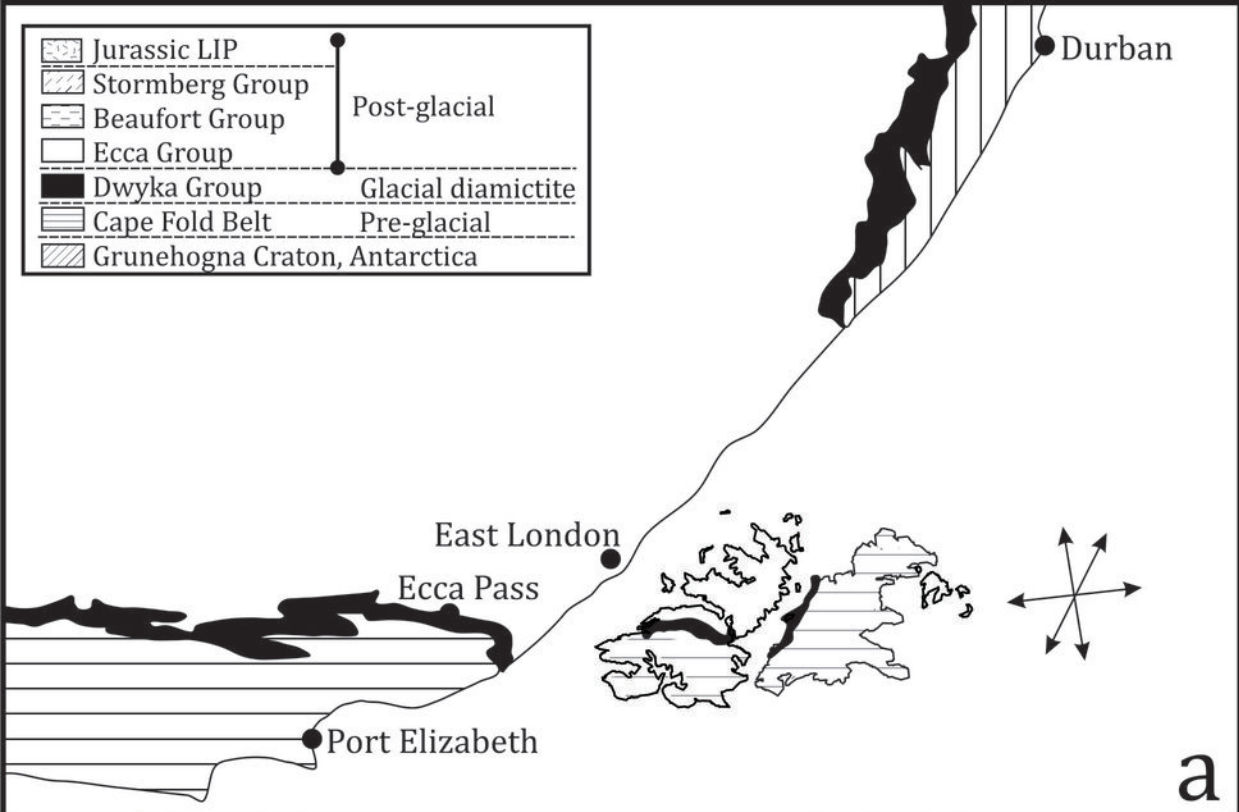
785 Figure 14. Ti/Zr *versus* SiO<sub>2</sub>, for a) Karoo LIP volcanic rocks and b) Ferrar LIP volcanic rocks.  
786 Note that the KwaZulu-Natal and Theron Mountains low TiO<sub>2</sub> samples are considered to be  
787 magma types transitional between Karoo and Ferrar types. Data sources; Transantarctic  
788 Mountains and Theron Mountains; Hergt *et al.* (1989), Brewer *et al.* (1992), Elliot *et al.* (1995),  
789 Fleming *et al.* (1995), Molzahn *et al.* (1996), Wilhelm & Worner (1996), Antonini *et al.* (1999),  
790 Elliot *et al.* (1999), Elliot & Fleming (2004). Karoo (including Dronning Maud Land) Luttinen  
791 *et al.* (1998), Luttinen & Furnes (2000), Heinonen & Luttinen (2008), Heinonen *et al.* (2010;  
792 2013; 2014) Neumann *et al.* (2011). Kirwanveggan; Harris *et al.* (1990).

793 Figure 15. a) Olivine equilibration temperatures (°C) *versus* Mg# of liquid in equilibrium with  
794 olivine for Ahlmannryggen dykes (filled dots; Heinonen & Luttinen 2008), Vestfjella high TiO<sub>2</sub>  
795 ferropicrite (filled triangles; Heinonen *et al.* 2013), Etendeka picrite (open squares; Kieding *et al.*  
796 *et al.* 2011) and Falklands Islands MAT intrusion (star in circle) and E-W intrusions (open  
797 triangles). Olivine equilibration temperatures have been calculated according to the scheme of  
798 Putirka *et al.* (2007). Vertical lines connecting points for Ahlmannryggen samples are calculated  
799 equilibration temperatures for different olivine phenocrysts in individual whole-rock samples.  
800 Figures in italics are T<sub>p</sub> from melt inclusions for Etendeka samples plotted in Fig. 15a (Kieding  
801 *et al.* 2011). b) Inferred temperature-pressure conditions at which fractional melting terminated  
802 for calculated primary magmas from Dronning Maud Land, the Karoo Province of southern  
803 Africa, Ferrar dolerites of Antarctica and picrites of the Etendeka Province of western Africa.  
804 The diagram was constructed following the methods of Herzberg and Gazel (2009) with data for  
805 the Ferrar province and Iceland from Hole (2015). Samples with MgO > 20 weight % are  
806 shown schematically following an adiabatic pathway for T<sub>p</sub> = 1640°C. The diagonally shaded  
807 box on the temperature axis is the range of olivine equilibration temperatures, calculated at 0  
808 GPa, for olivine in ferro-picrite dykes from Dronning Maud following the method of Putirka *et al.*  
809 *et al.* (2007), and the box labelled 'MAT & E-W' is the same calculations for MAT and E-W  
810 intrusions. Adiabatic melting paths are labelled with mantle potential temperature. 2σ error bars  
811 are from Hole (2015). c) T<sub>p</sub> calculated from melt inclusions in ultra-magnesian olivines from the

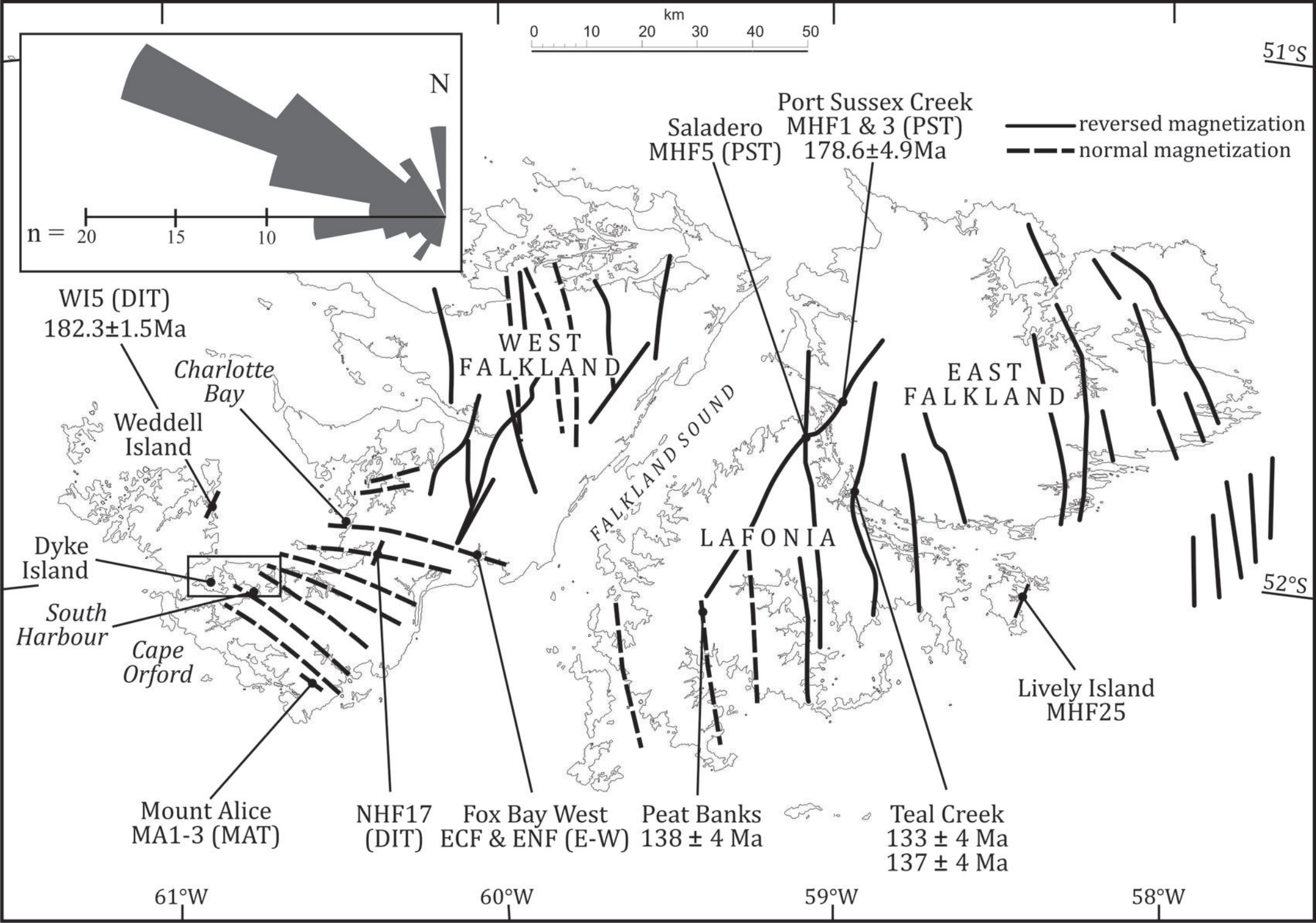
812 Etendeka Province *versus* olivine equilibration temperatures for the same samples. Data from  
813 Keiding *et al.* (2011).

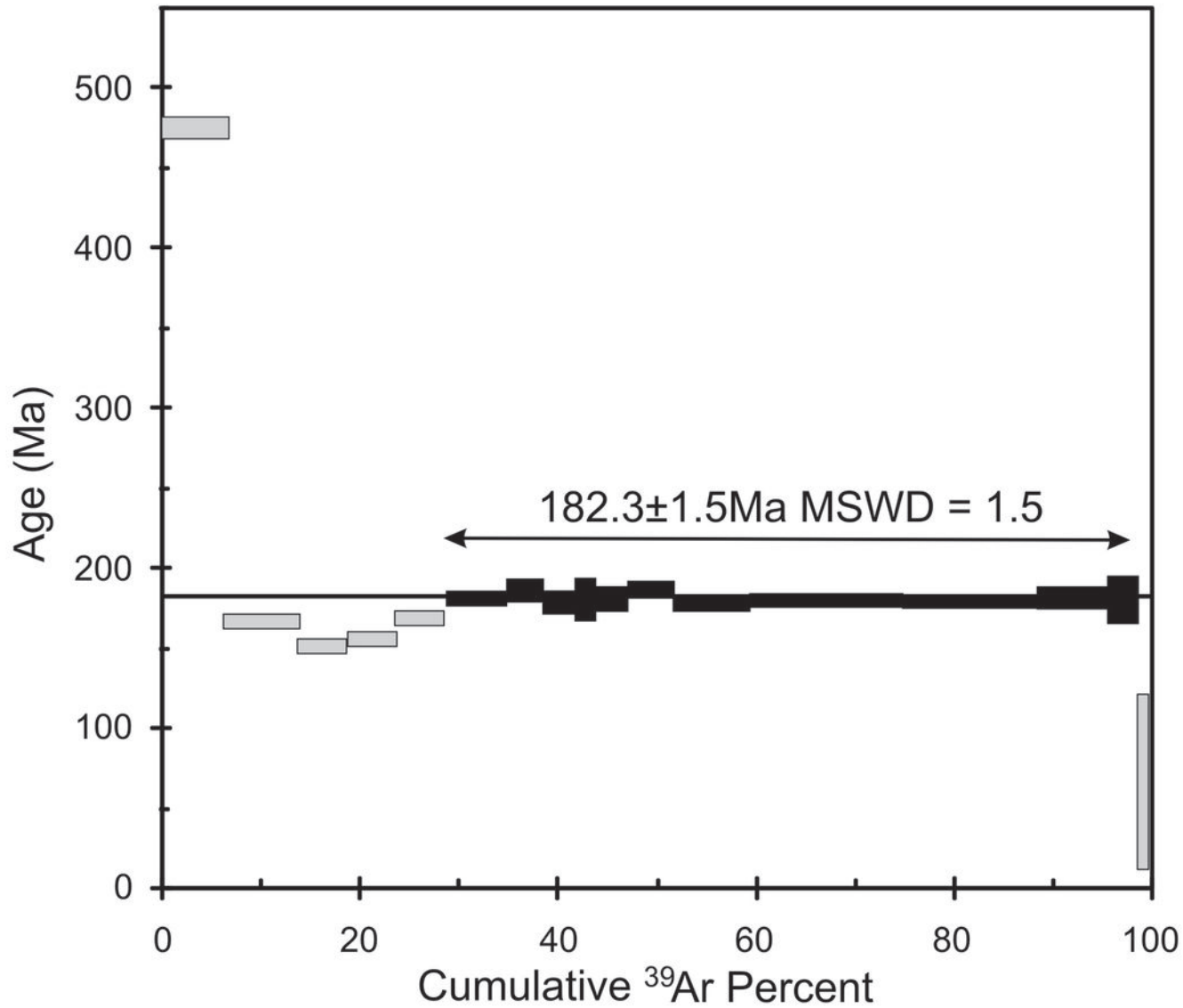
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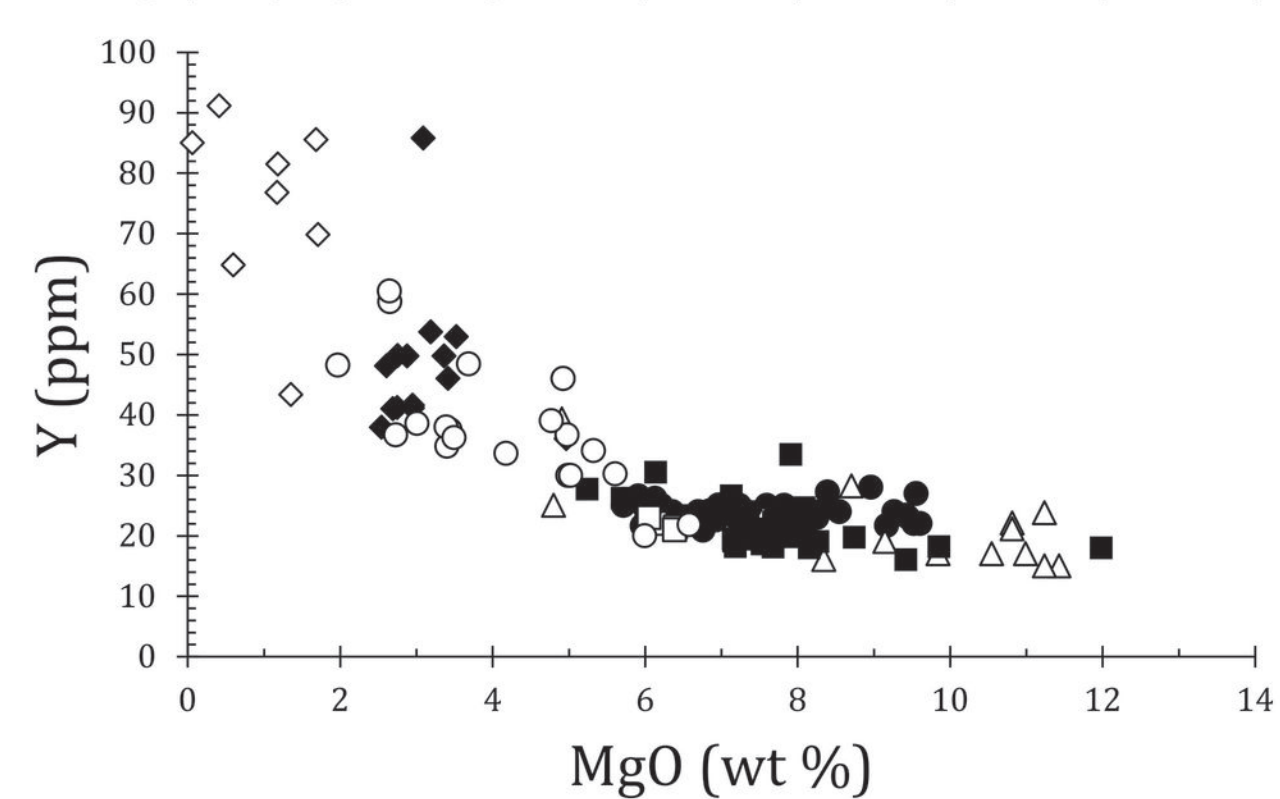
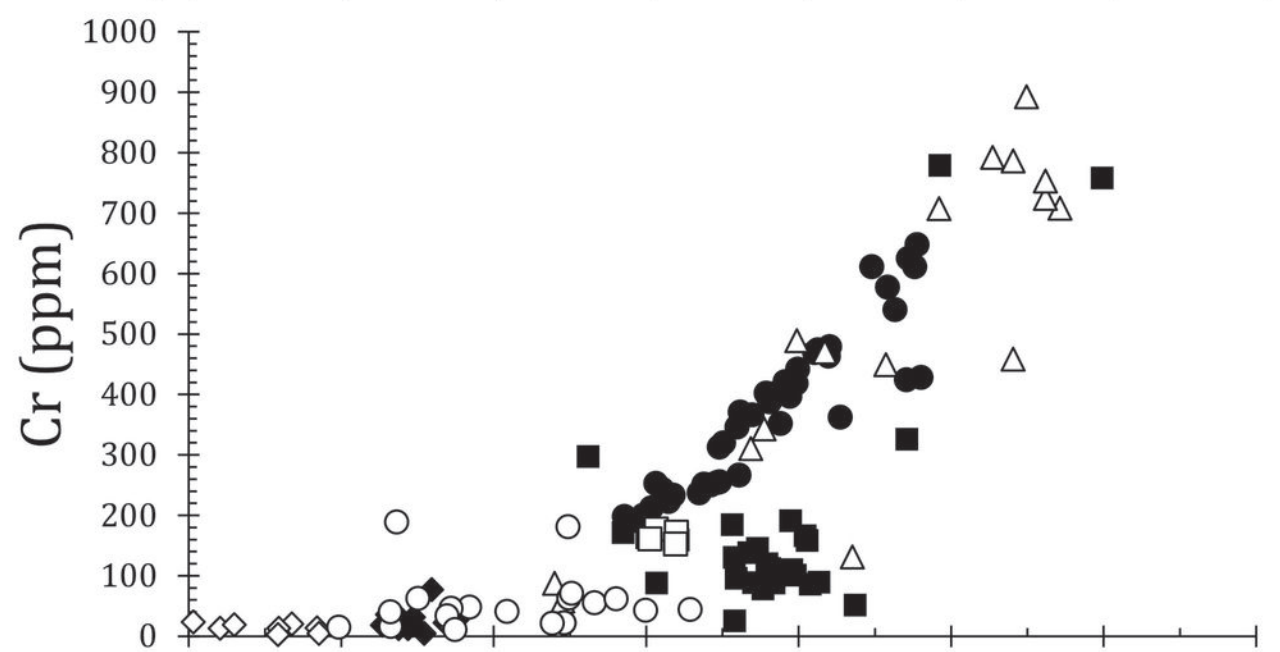
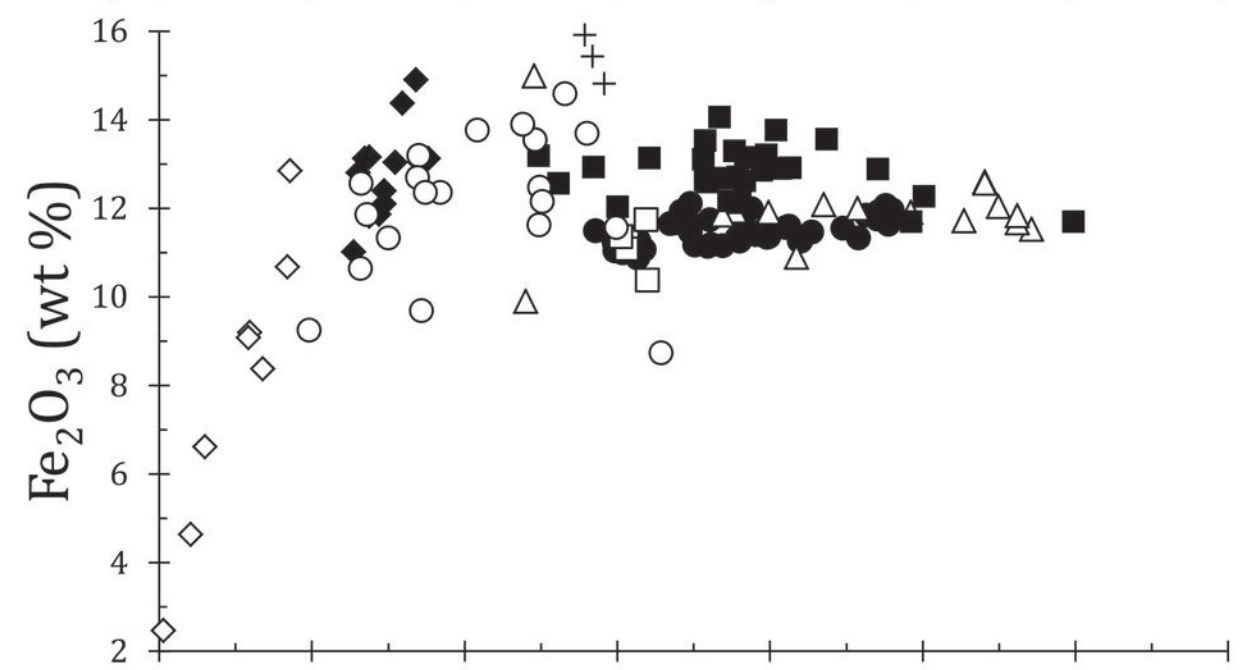
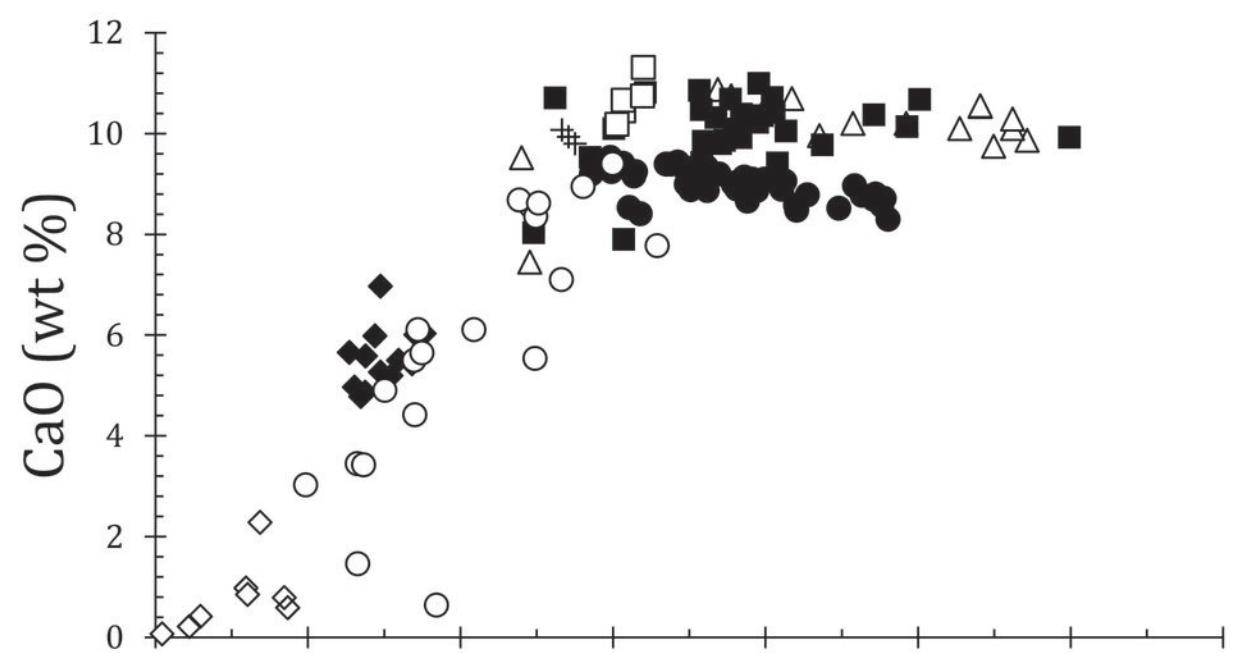
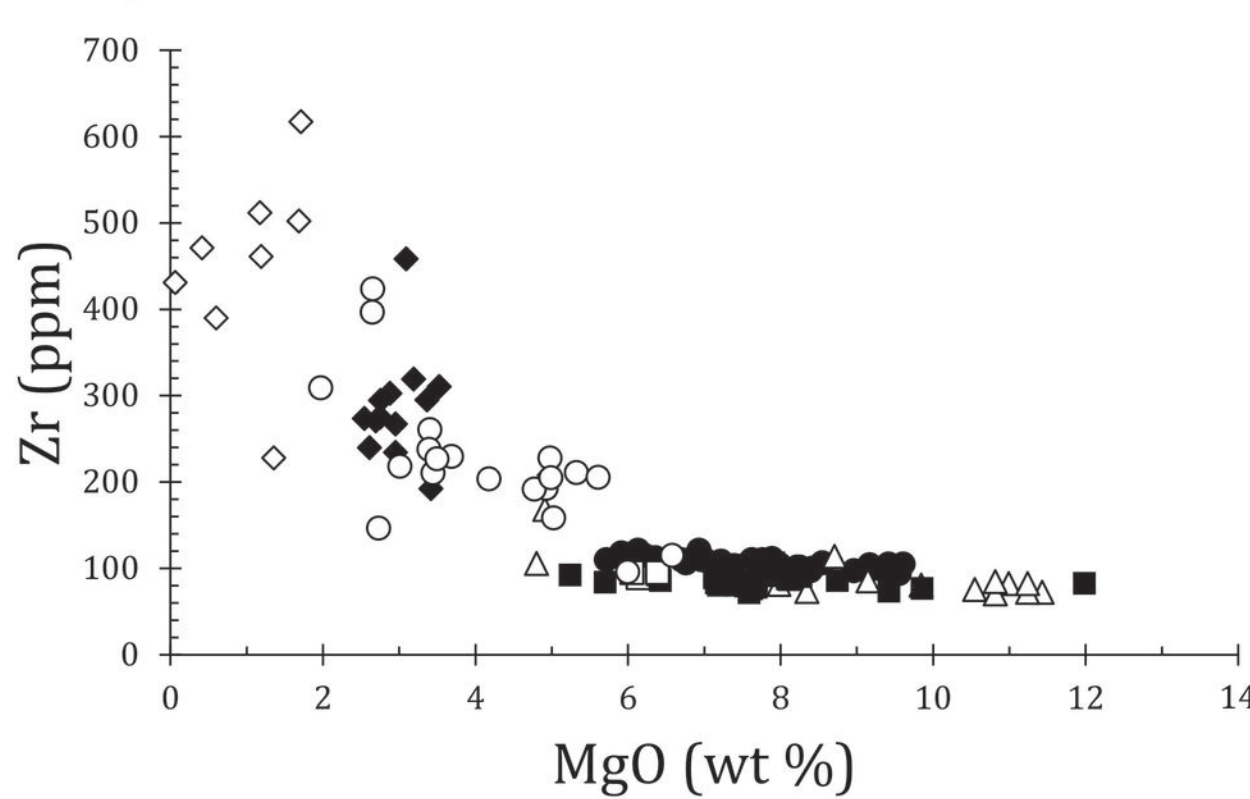
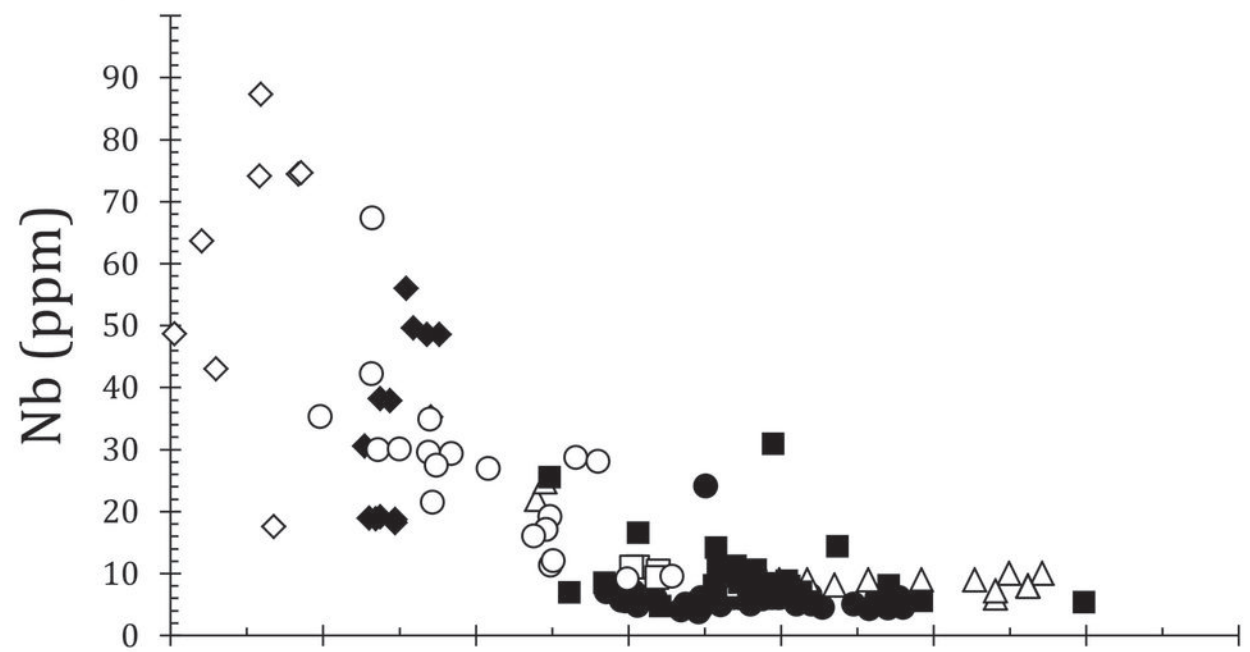
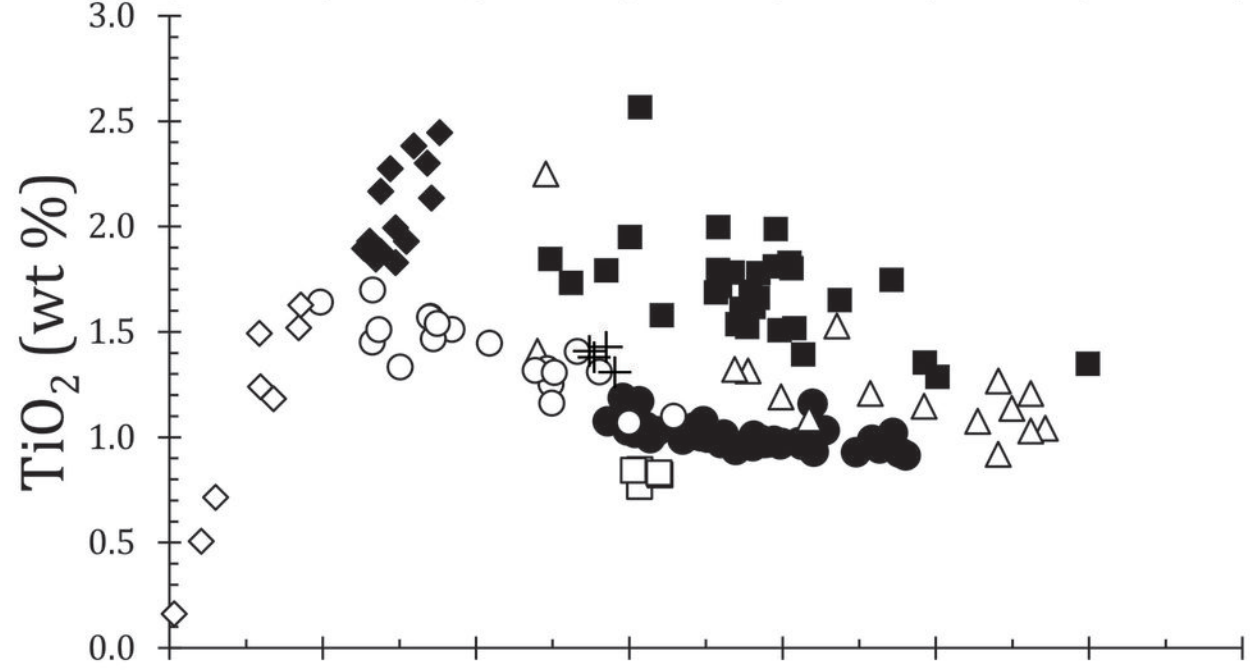
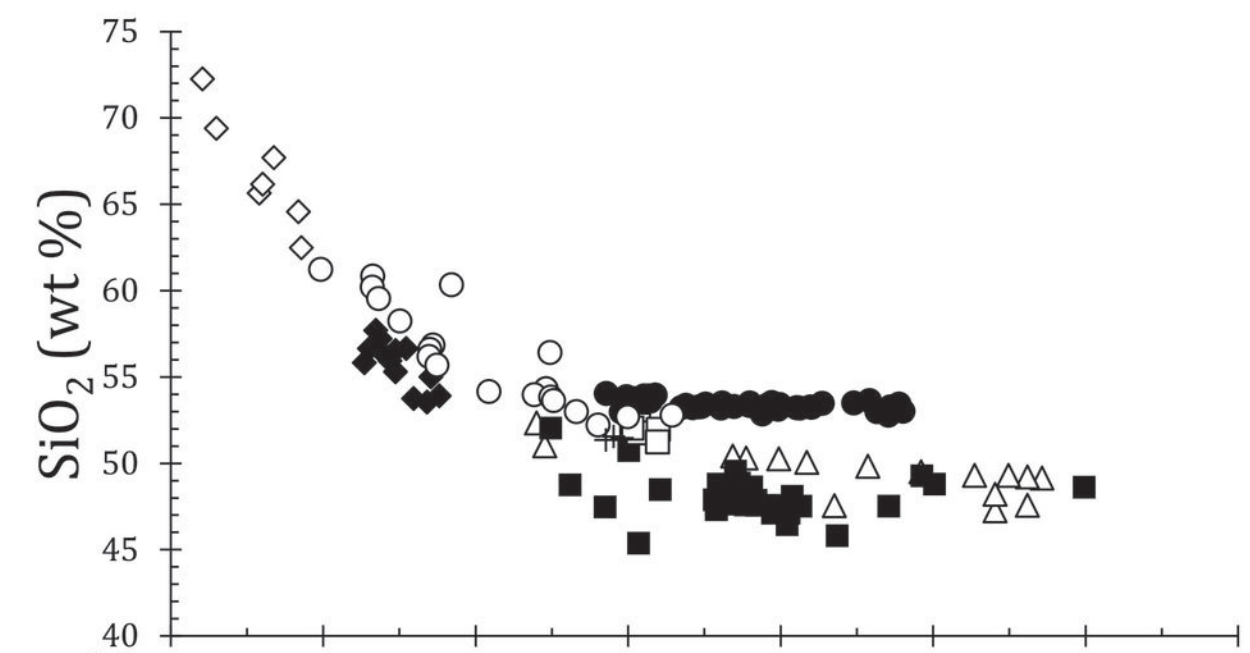












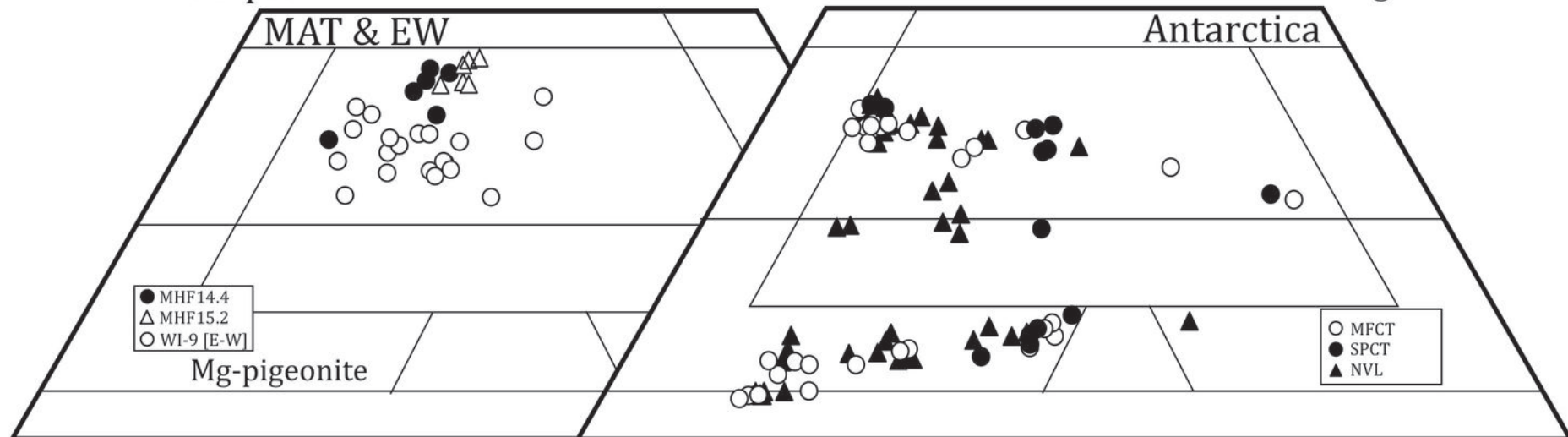


Diopside

Hedenbergite

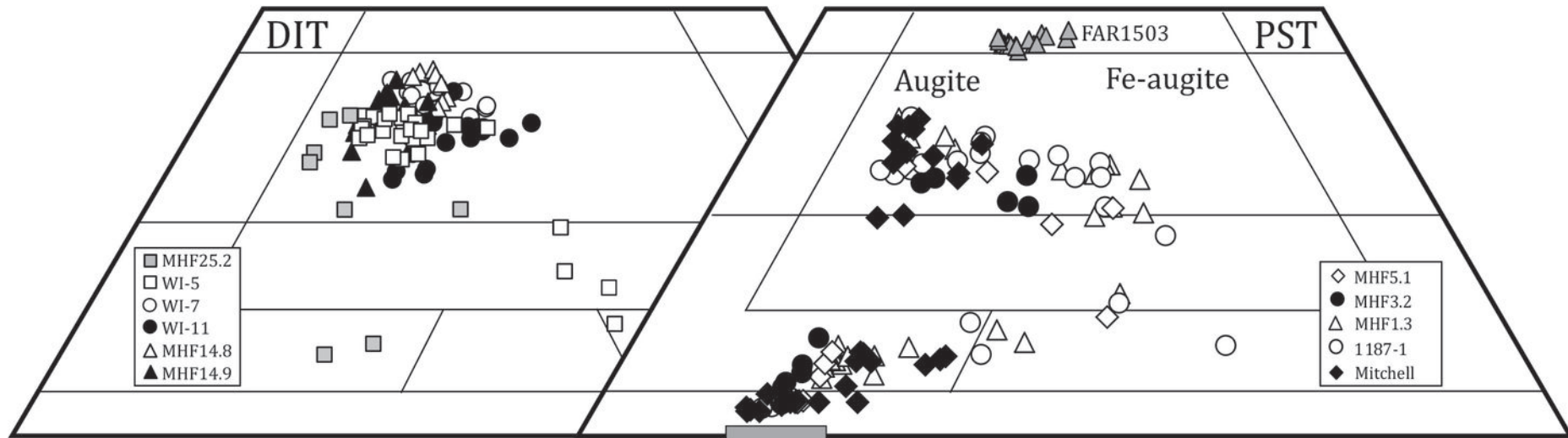
MAT &amp; EW

Antarctica



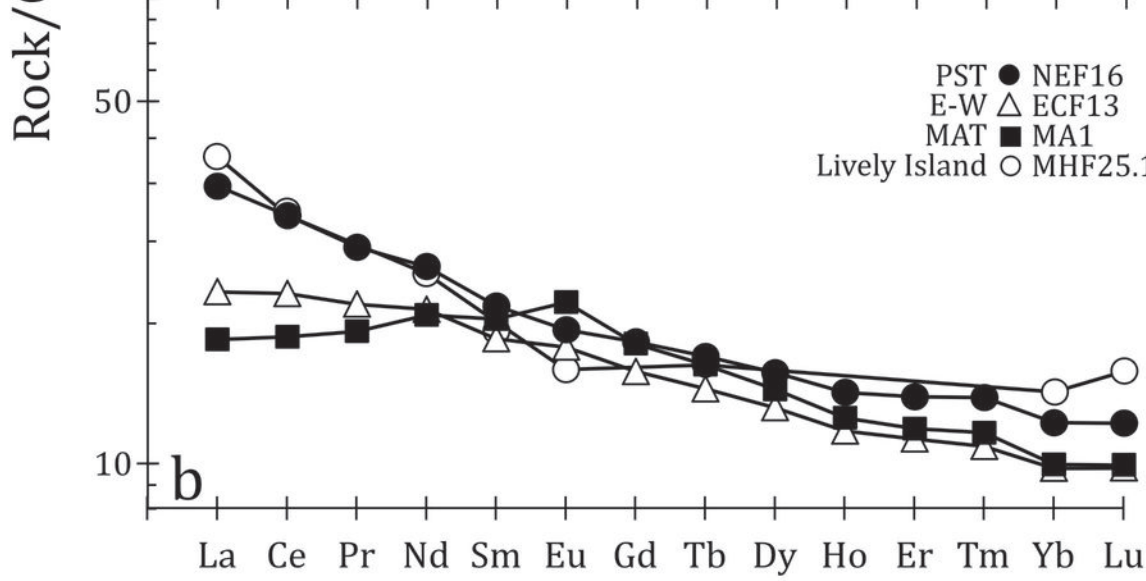
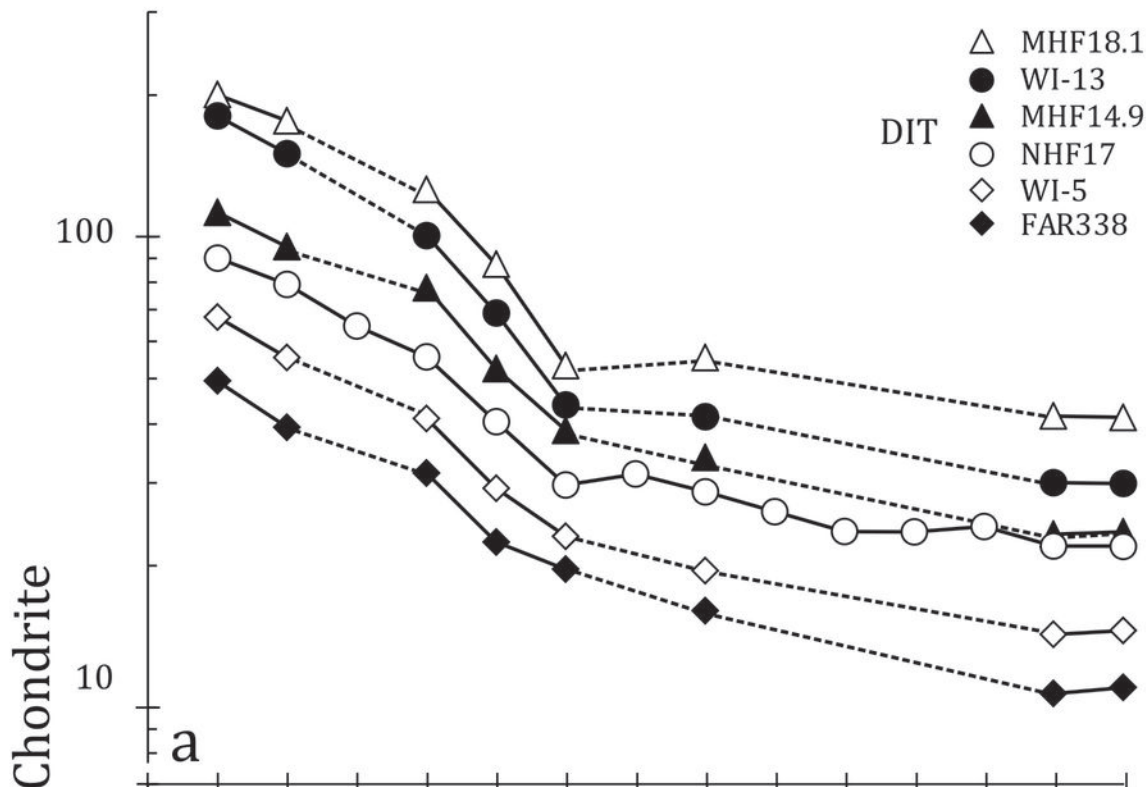
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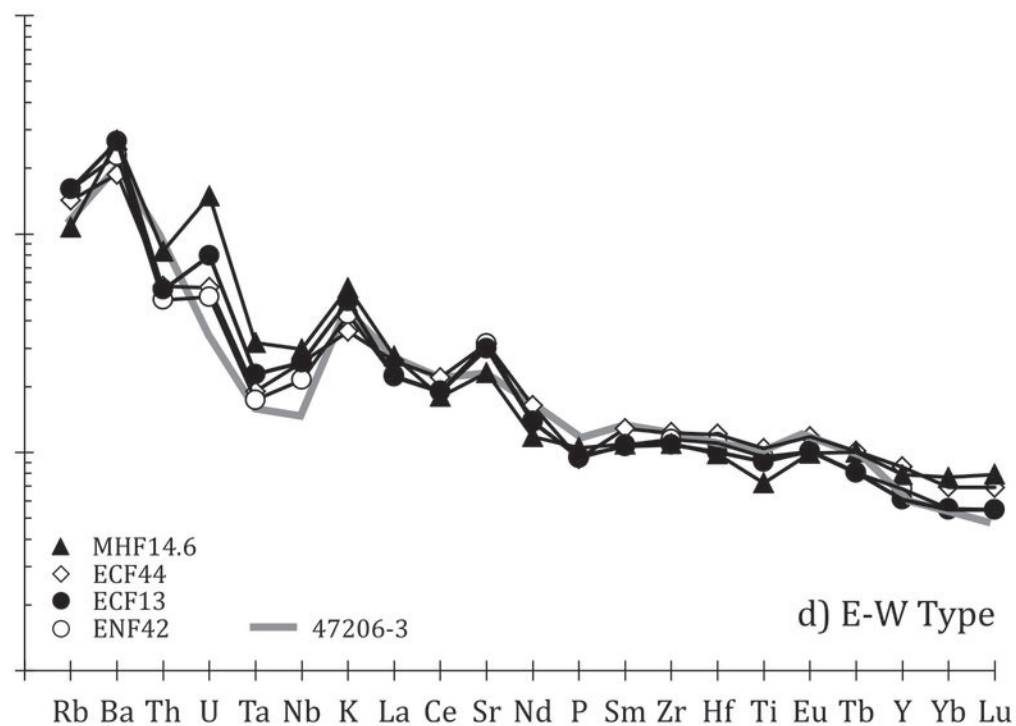
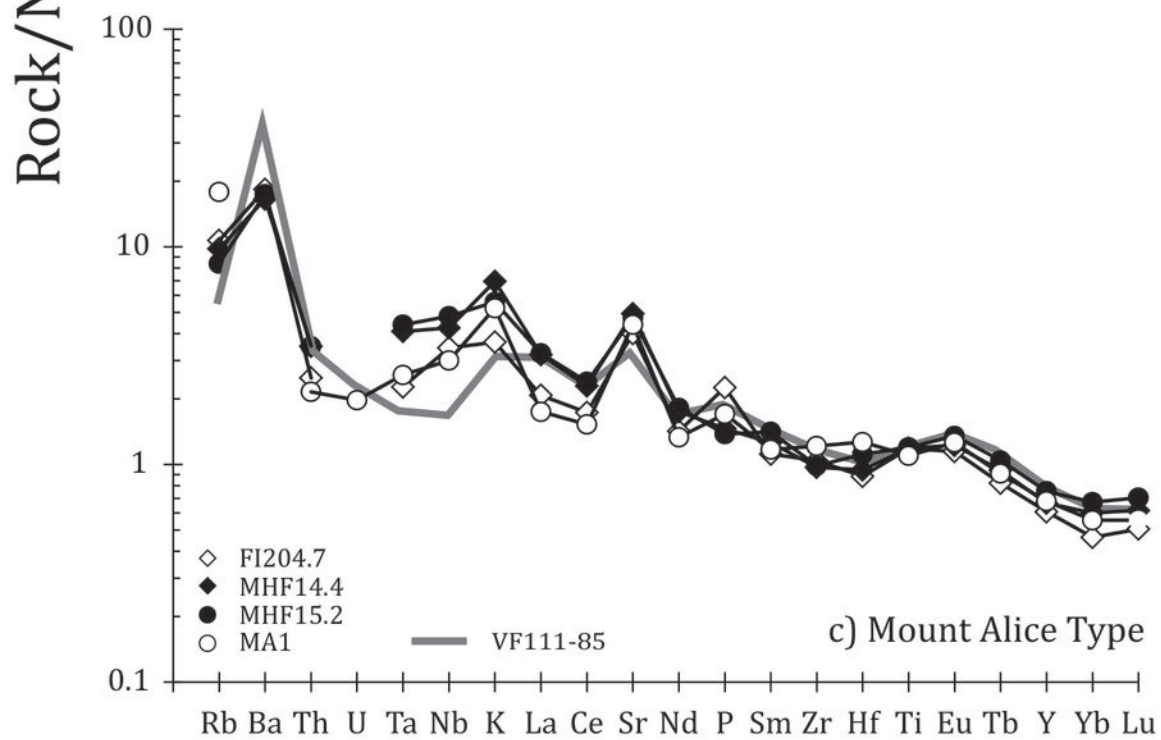
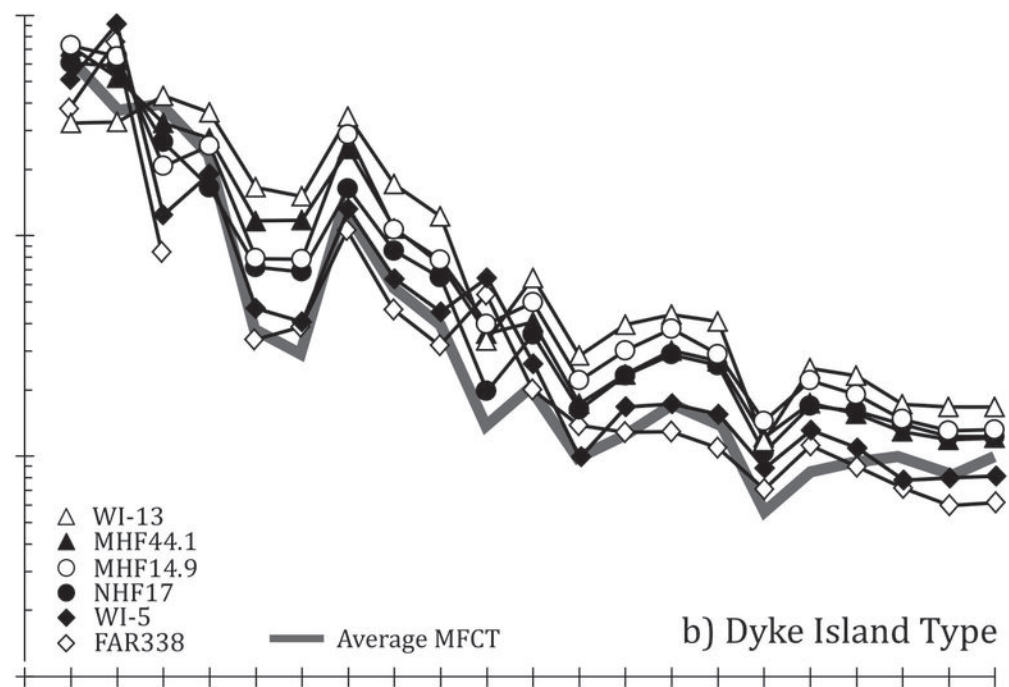
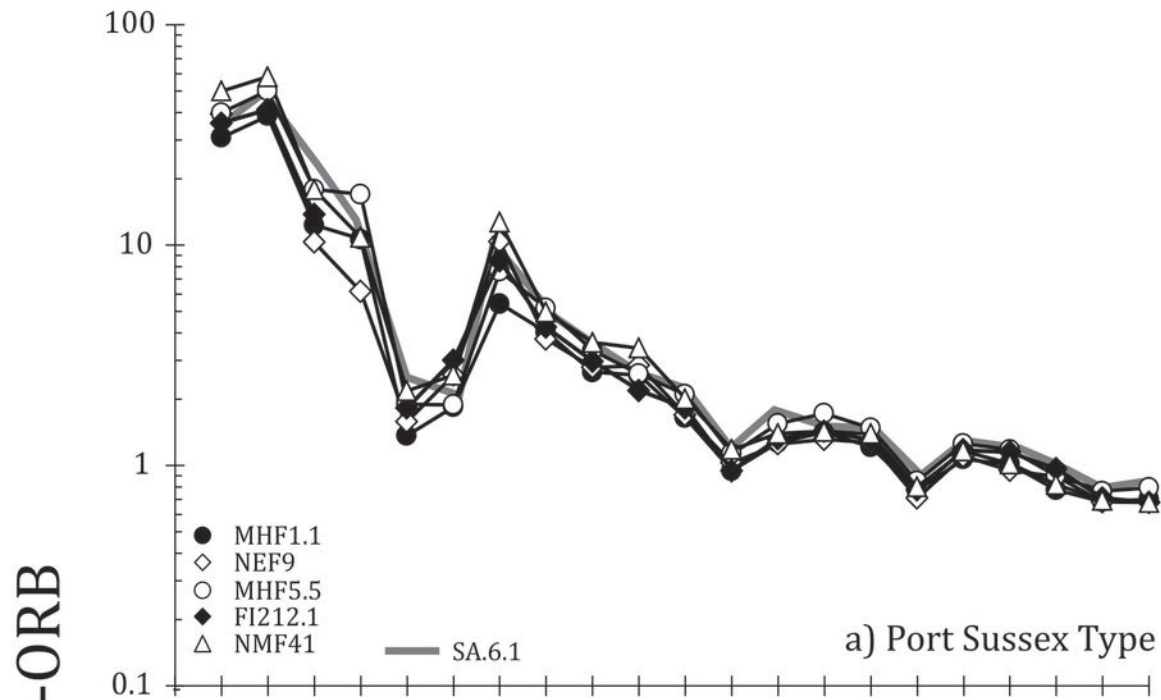
PST

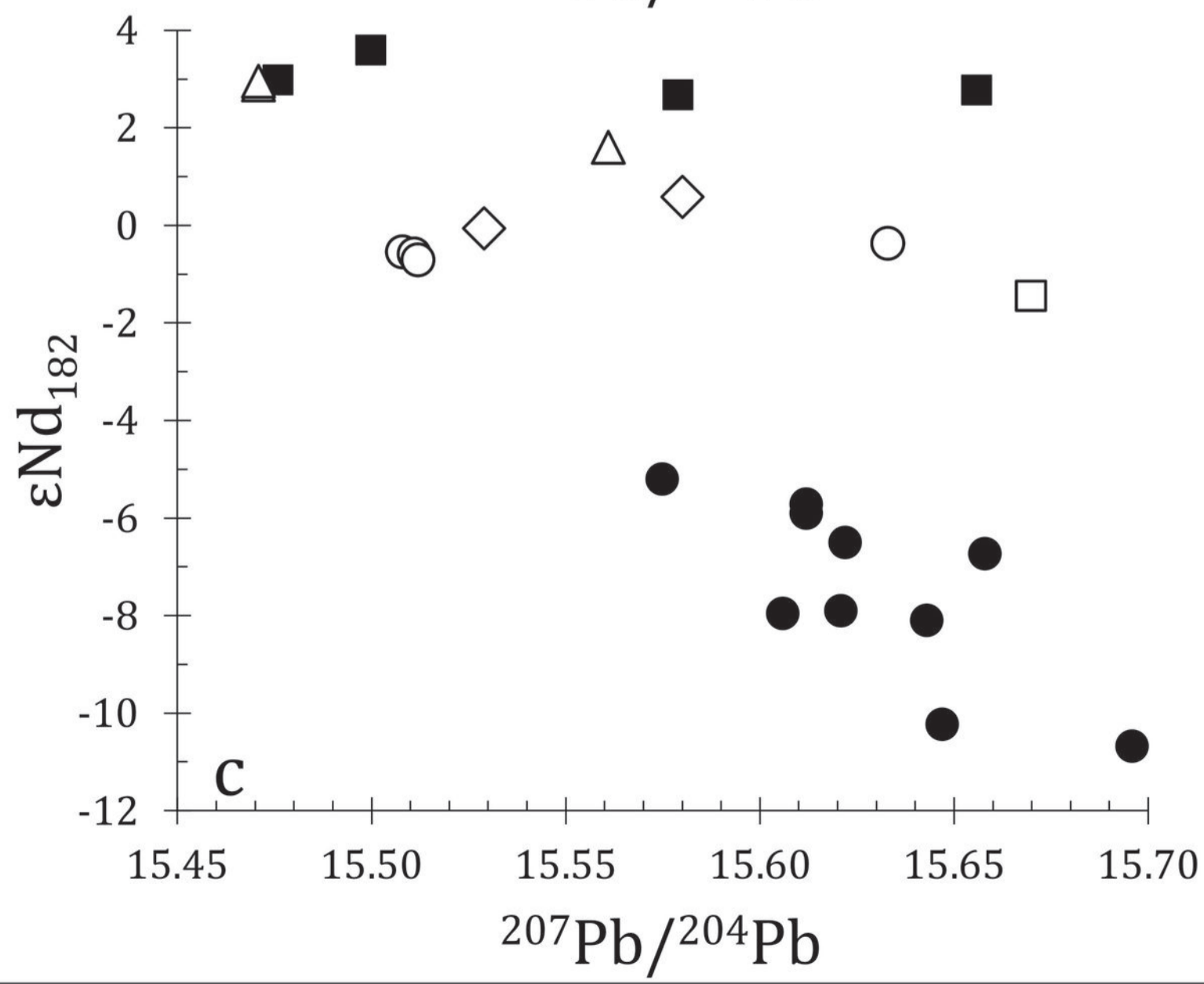
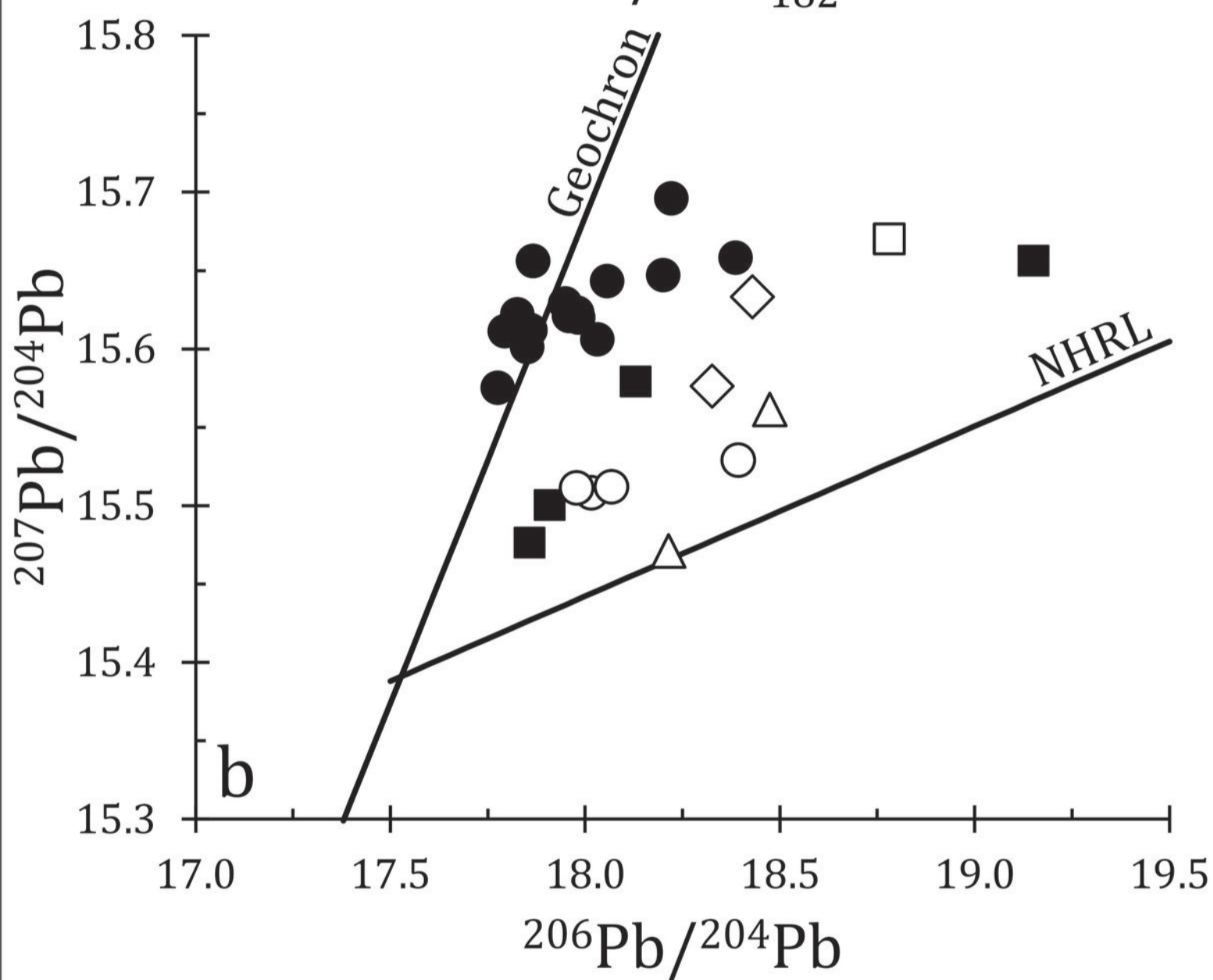
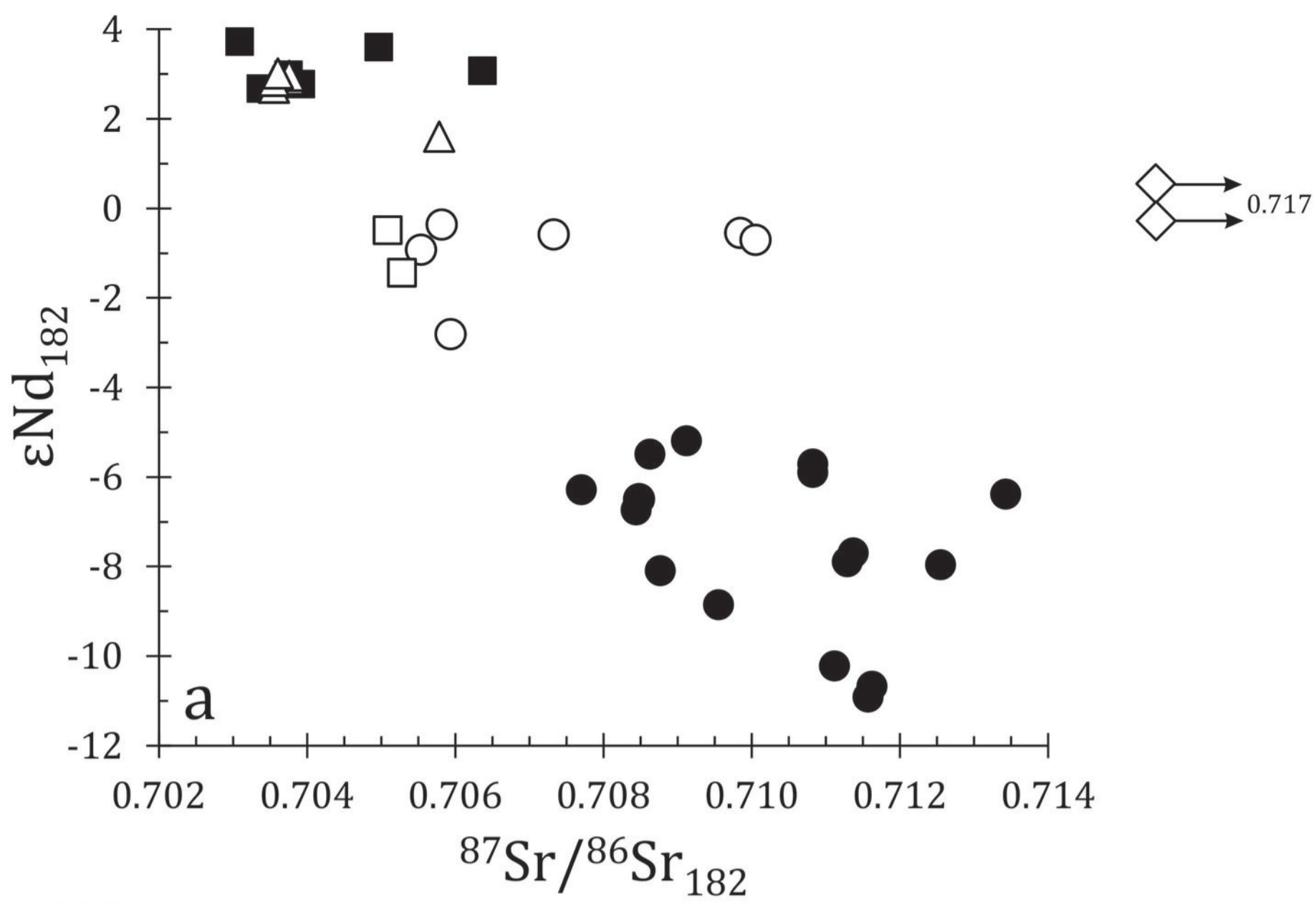
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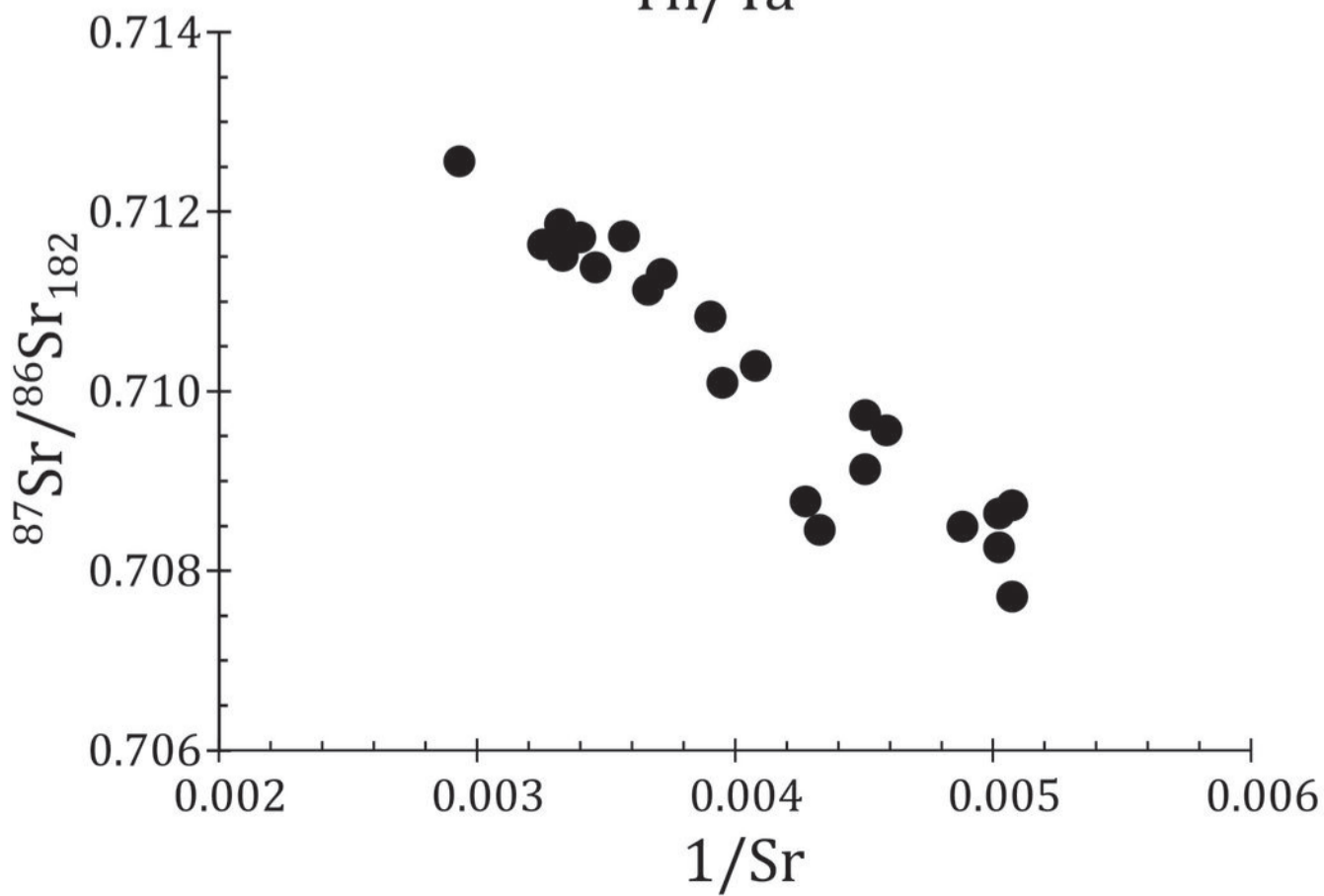
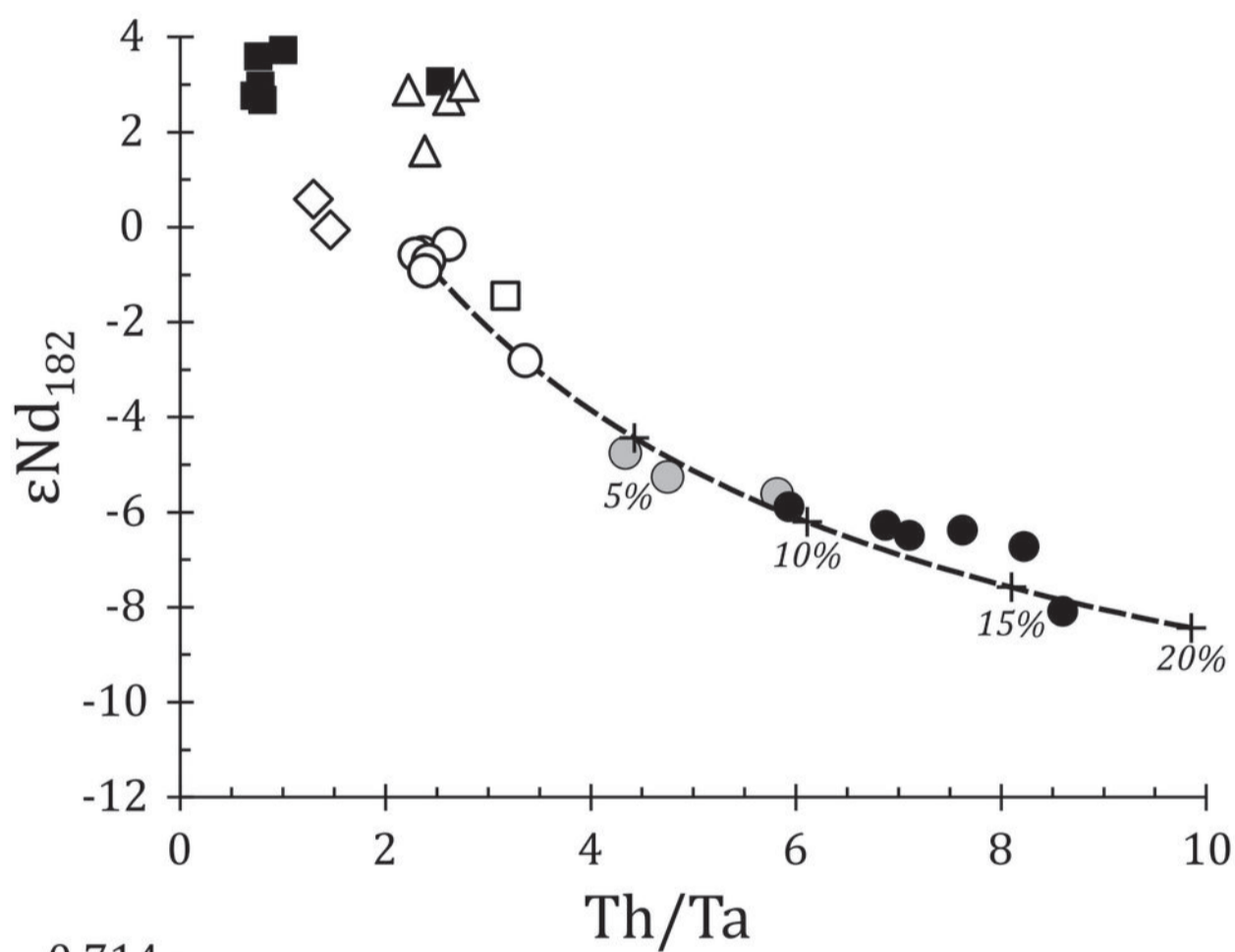
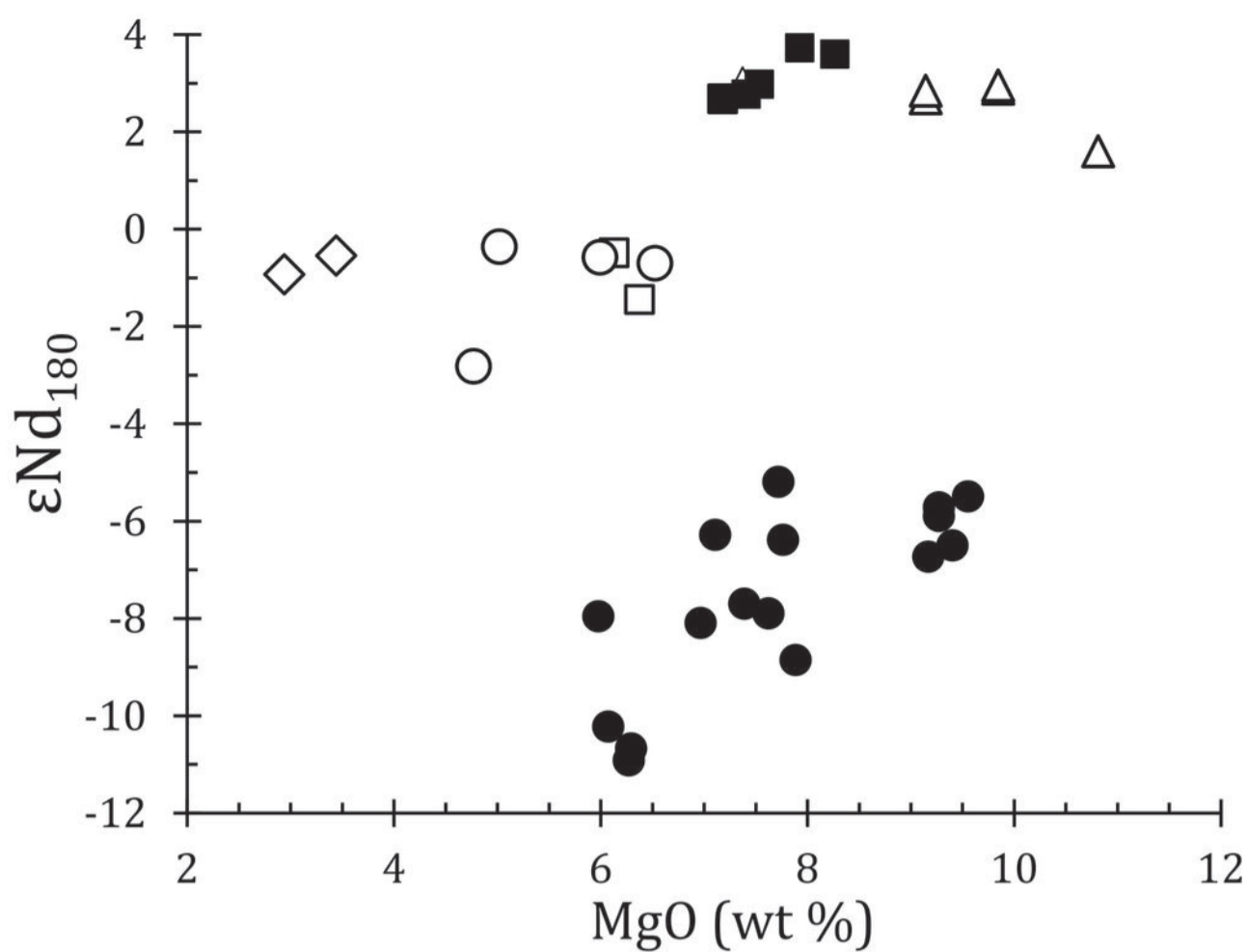
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Enstatite

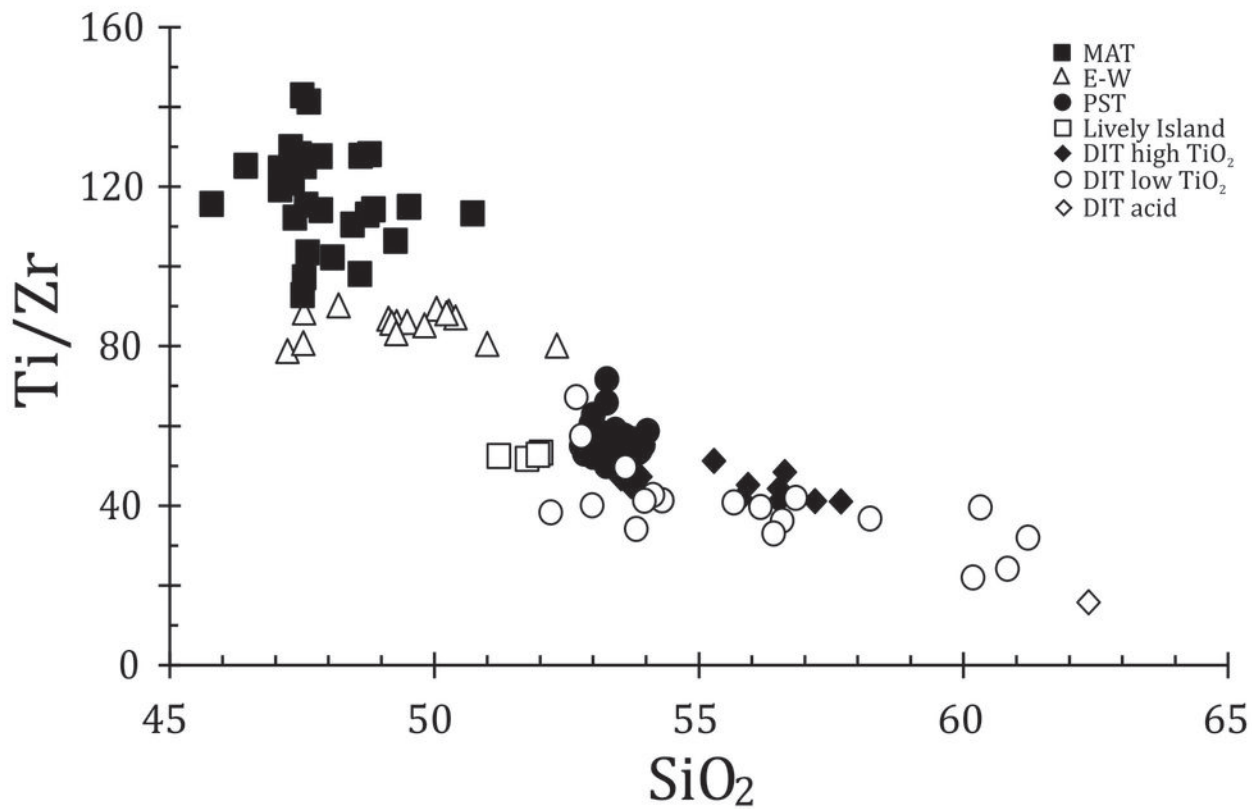


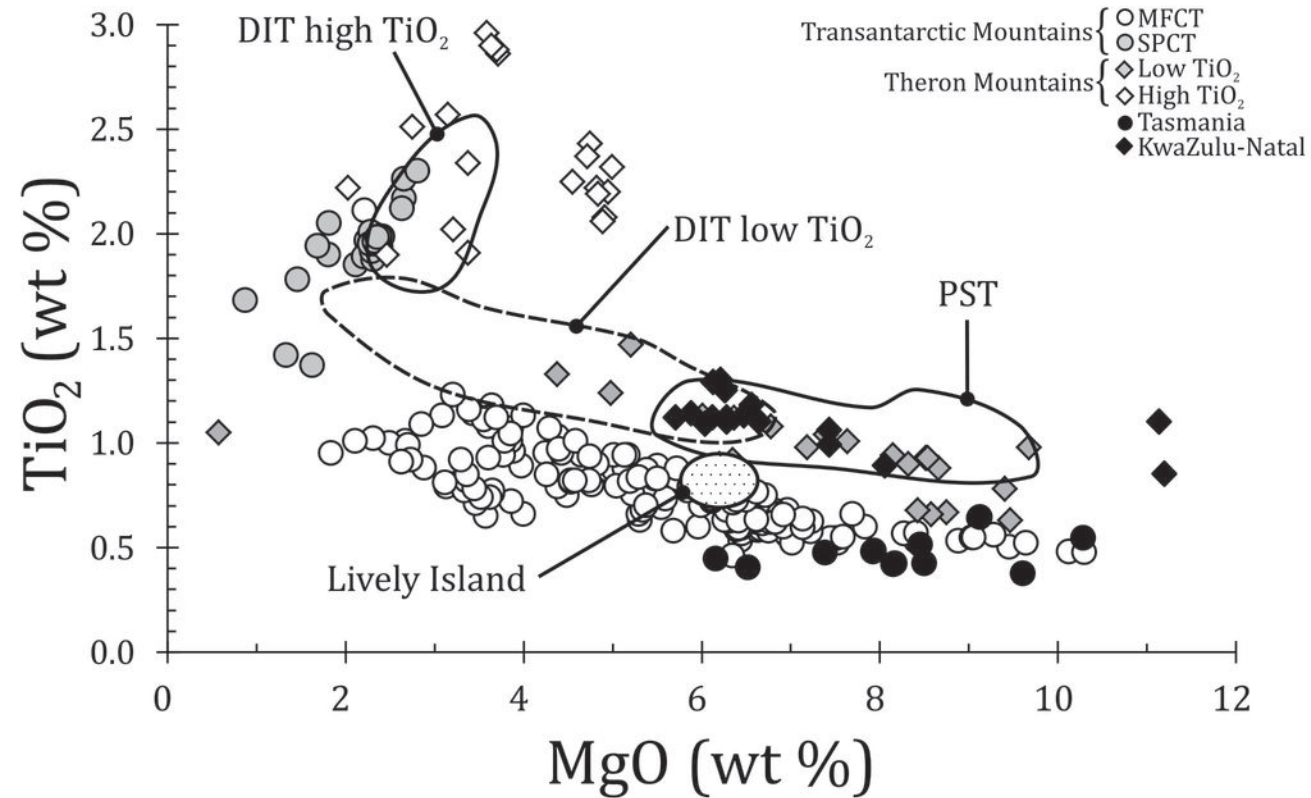


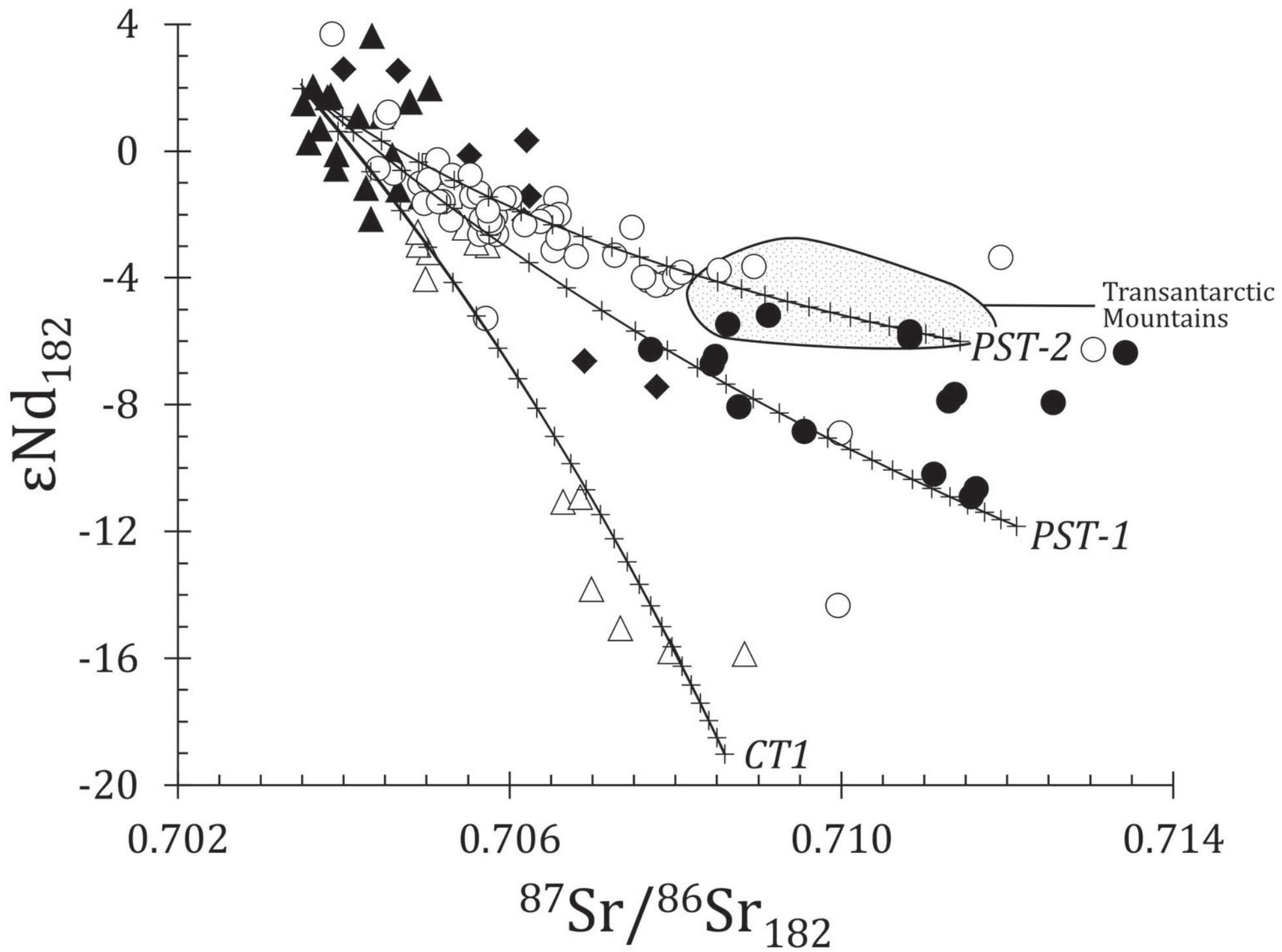


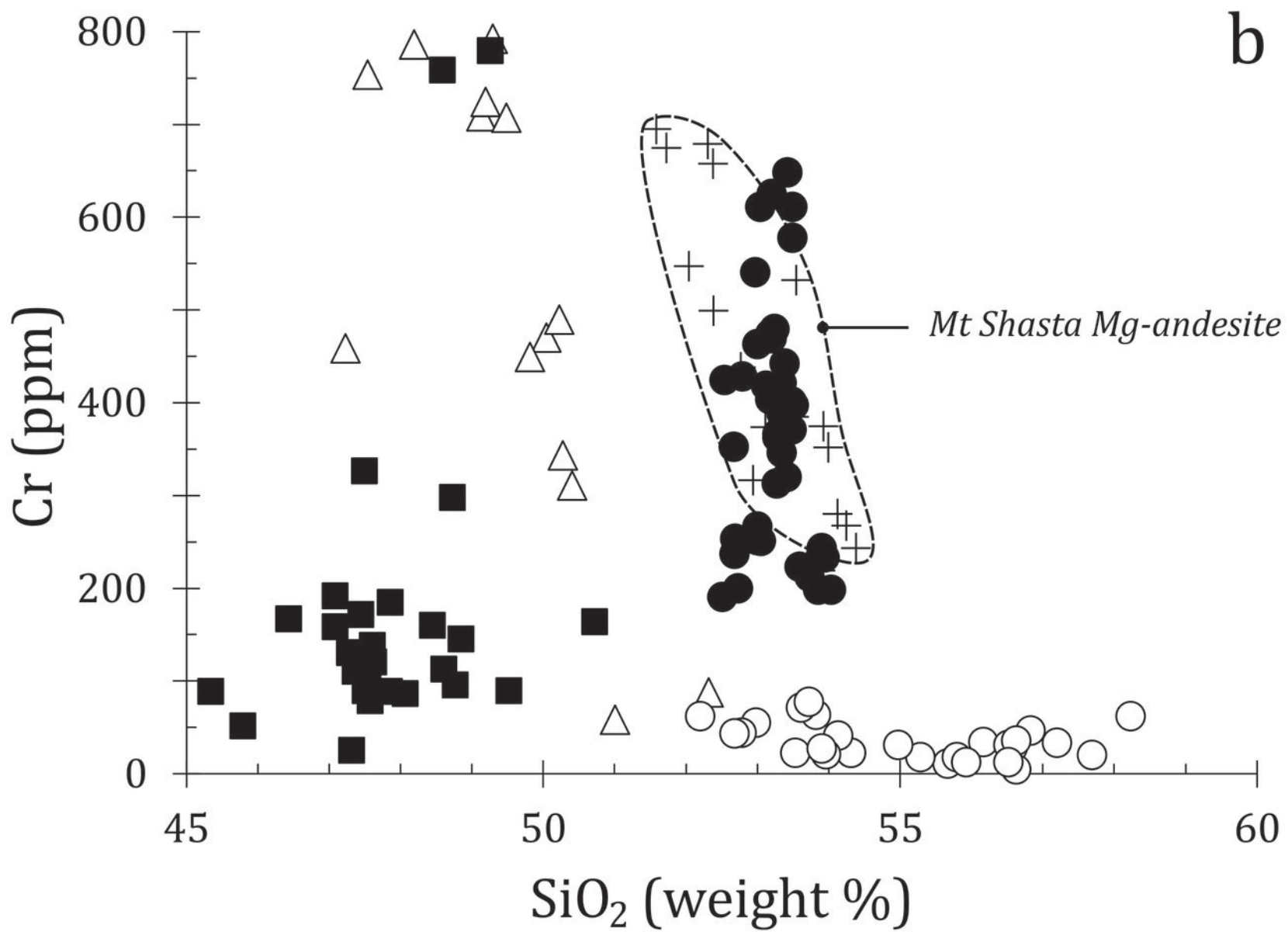
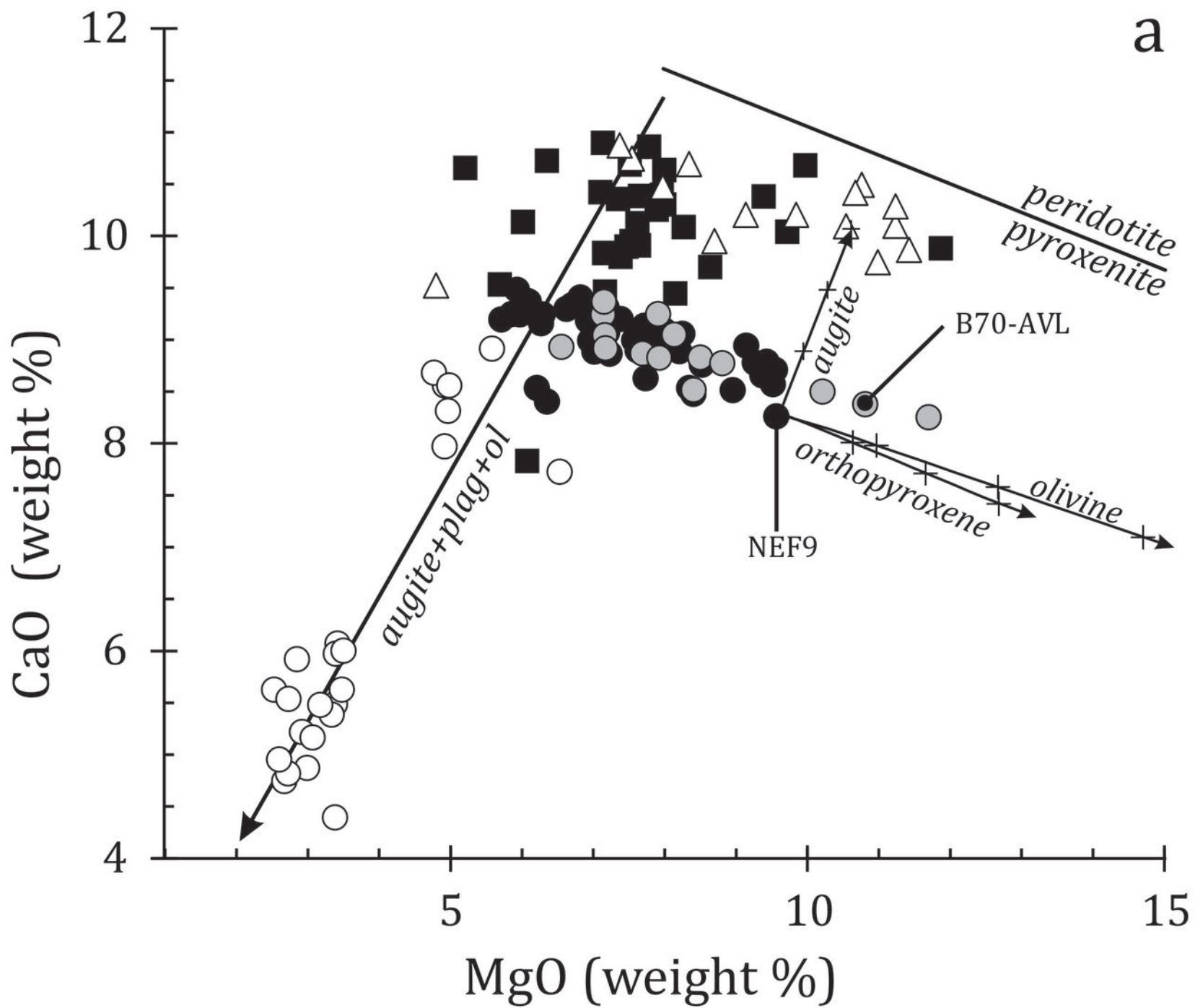


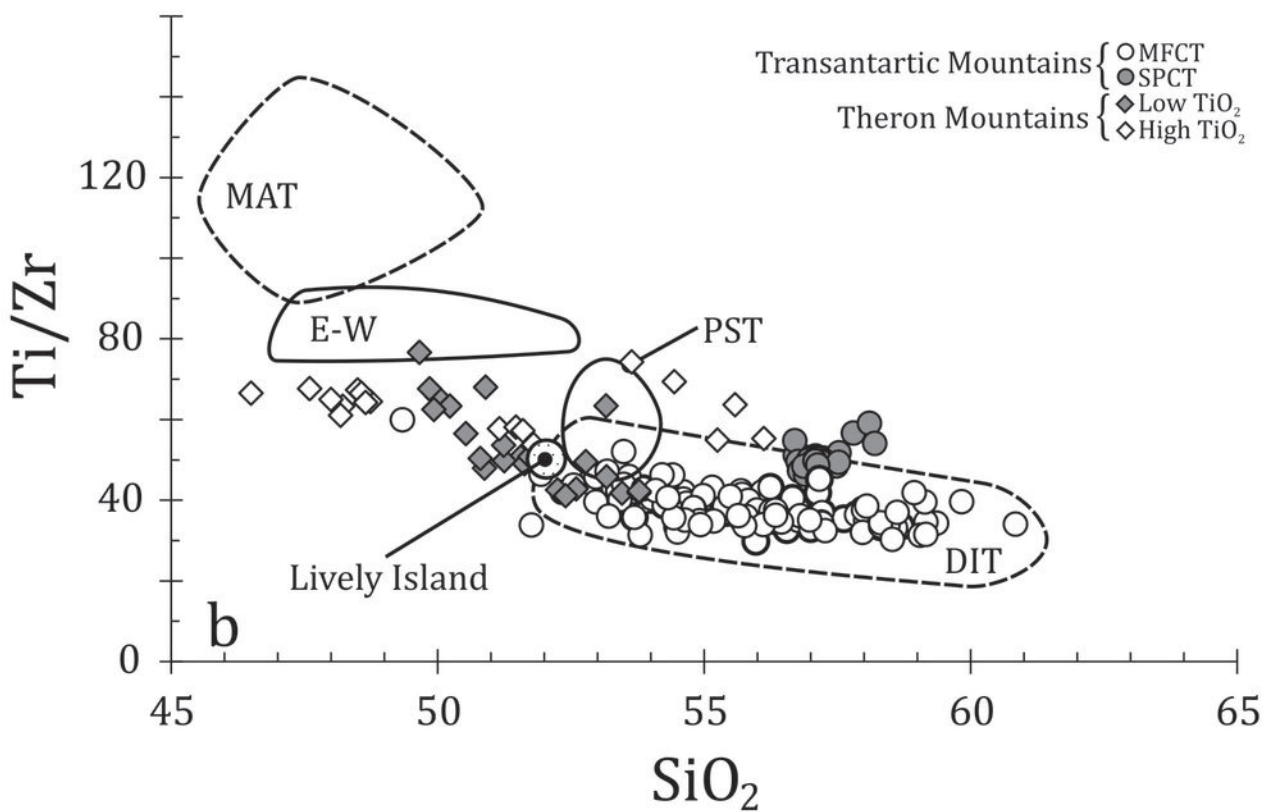
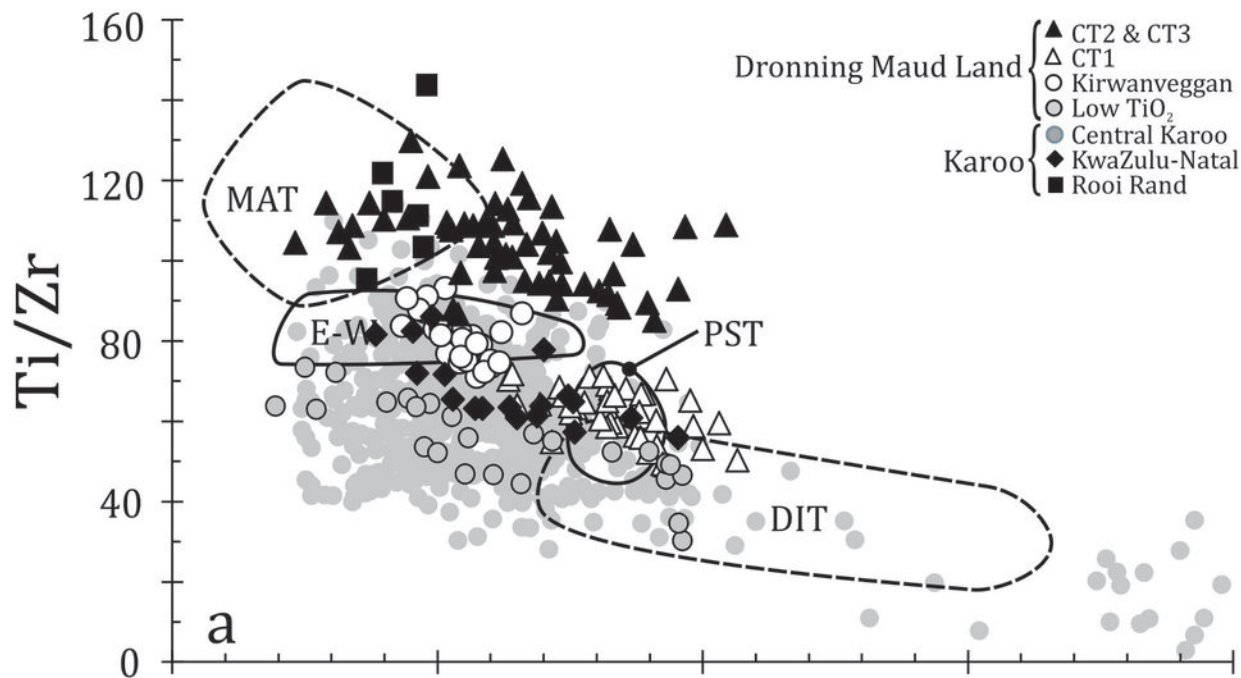




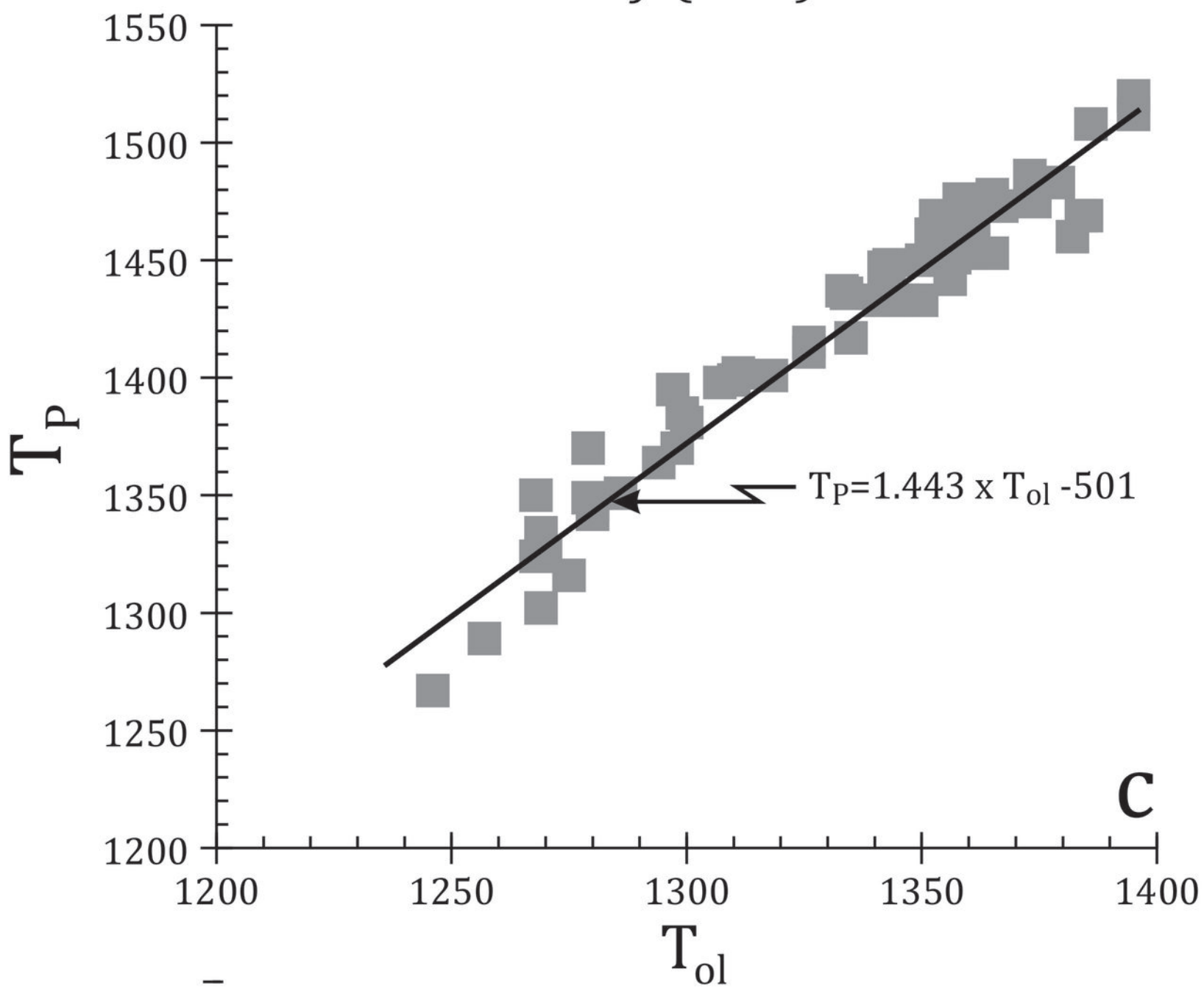
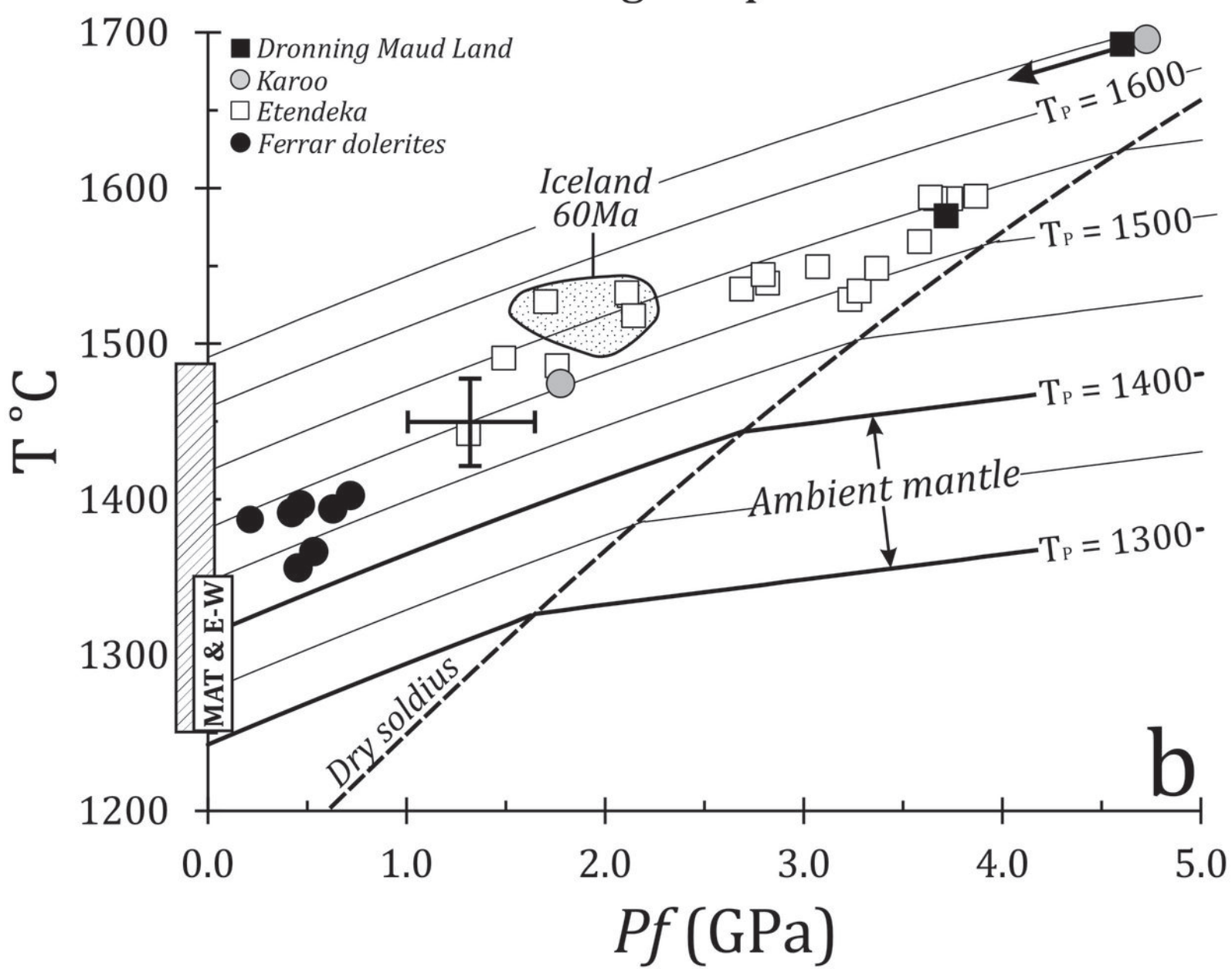
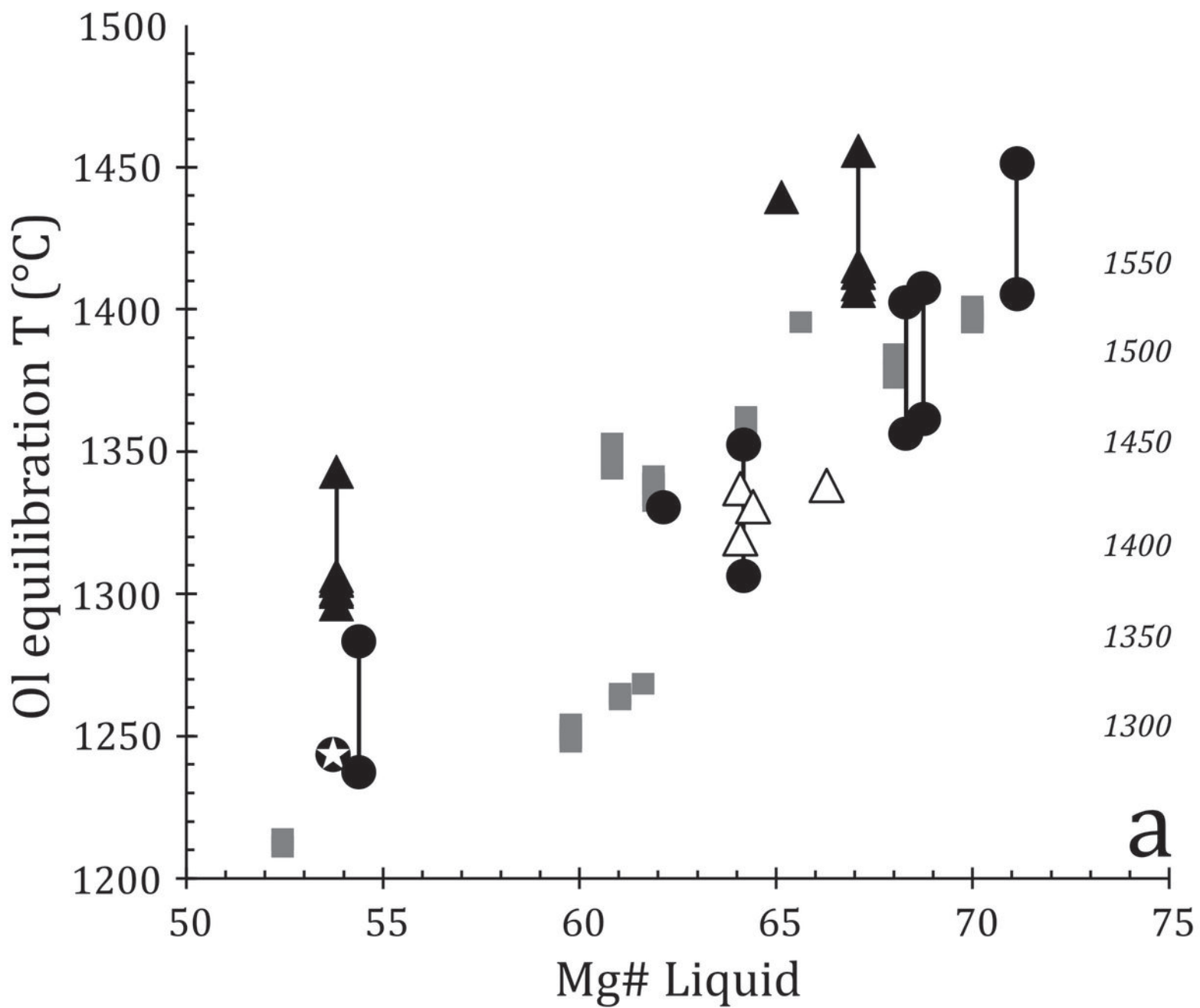












Sample	1187-1	#1503	NEF9	NMF41	MHF1.1	MHF5.5	ECF13	ECF42	ECF44	MHF14.6
Type	PST	PST	PST	PST	PST	PST	EW	EW	EW	EW
SiO <sub>2</sub>	52.67	53.20	52.41	53.59	53.49	53.01	49.49	49.82	50.40	47.02
TiO <sub>2</sub>	0.97	1.02	0.89	1.02	0.99	1.07	1.15	1.21	1.32	0.92
Al <sub>2</sub> O <sub>3</sub>	14.00	12.90	15.69	14.87	13.15	13.90	14.30	14.47	15.22	14.79
Fe <sub>2</sub> O <sub>3</sub>	11.94	11.90	11.50	11.25	11.29	12.03	11.91	11.99	11.85	12.54
MnO	0.14	0.14	0.13	0.17	0.13	0.16	0.17	0.18	0.02	0.19
MgO	7.74	9.43	5.96	6.29	9.15	6.93	9.85	9.14	7.38	10.77
CaO	8.63	8.79	9.35	9.24	8.94	9.18	10.20	10.20	10.87	10.49
K <sub>2</sub> O	2.45	1.90	2.61	2.52	2.10	2.52	2.47	2.58	2.57	2.33
P <sub>2</sub> O <sub>5</sub>	0.98	0.46	0.77	0.92	0.40	0.55	0.36	0.31	0.26	0.41
Na <sub>2</sub> O	0.18	0.16	0.16	0.14	0.11	0.13	0.11	0.11	0.11	0.11
TOTAL	99.70	99.90			99.75	99.48				99.57
Loss	1.50	0.50			1.26	2.62				2.20
Rb	33.1	14.2	22	28	17	22	9	9	8	6
Sr	533	205	256	307	232	234	269	282	283	208
Nb	5.8	4.9	6	6	4.3	4.4	6	5	6	6.9
Zr	110	93	97	106	104	113	80	85	91	81
Y	23.9	23.1	25	23	22	25.1	17	19	24	22.1
Cr	352	625	589	223	577	255	707	449	310	458
Ni	100	170	133	75	146	91	236	174	113	254
Ba	nd	nd	1235	365	250	316	167	144	126	208
La	11.6	9.3	9.35	12.42	10.10	12.90	5.54	5.60	6.55	6.9
Ce	25.1	20.1	20.87	27.19	19.90	24.90	14.19	14.29	16.48	13.5
Pr			2.77	3.44			2.09	2.10	2.45	
Nd	14.7	12.3	12.37	14.64	12.00	15.3	10.01	10.16	11.95	8.6
Sm	3.58	3.15	3.29	3.68	3.34	4.06	2.80	2.84	3.37	2.87
Eu	1.22	1.05	1.12	1.19	1.09	1.28	1.03	1.03	1.20	1.01
Gd			3.76	4.16			3.25	3.29	4.10	
Tb	0.72	0.63	0.64	0.68	0.77	0.79	0.54	0.54	0.68	0.67
Dy			3.98	4.20			3.35	3.35	4.42	
Ho			0.80	0.83			0.66	0.67	0.85	
Er			2.30	2.37			1.87	1.90	2.41	
Tm			0.35	0.36			0.28	0.28	0.36	
Yb	2.14	1.93	2.08	2.12	2.09	2.33	1.66	1.68	2.11	2.34
Lu	0.33	0.28	0.31	0.31	0.32	0.36	0.25	0.25	0.31	0.36
Th	1.83	1.35	1.24	2.14	1.48	2.15	0.67	0.60	0.69	1.0
U	0.80	0.50	0.29	0.51	0.5	0.8	0.37	0.24	0.27	0.7
Ta	0.24	0.19	0.21	0.29	0.18	0.25	0.30	0.23	0.25	0.42
Hf	2.73	2.33	2.6	2.86	2.49	3.03	2.05	2.27	2.48	2.01
Pb			5.55	4.93			2.19	2.65	2.65	
Cs	0.42	0.29			2.24	0.74				0.3
Co	43.30	47.30			50.8	42.9				61.4
Sc	25.70	26.60			27.8	28.2				35
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.71448	0.70900			0.70900	0.70948				0.70600
± 2SE	±1	± 2			± 18	± 17				±2
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>182</sub>	0.71343	0.70848	0.71083		0.70845	0.70877	0.70579	0.70376	0.70355	0.70578
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512252	0.512255			0.512259	0.512181				0.512726
± 2SE	±7	±5			± 6	± 6				±7
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>182</sub>	0.512069	0.512063	0.512367		0.51205	0.511981	0.512561	0.512551	0.512565	0.512475
εNd <sub>182</sub>	-6.4	-6.5	-5.7		-6.7	-8.1	2.9	2.7	3.0	1.6
<sup>206</sup> Pb/ <sup>204</sup> Pb		17.826±11	17.86		18.394±8	18.057± 7		18.21		18.474±7
<sup>207</sup> Pb/ <sup>204</sup> Pb		15.622±8	15.61		15.658± 8	15.643±8		15.47		15.561±8
<sup>208</sup> Pb/ <sup>204</sup> Pb		38.044±19	37.97		38.143±20	38.246±16		37.58		38.111±15
Sample	1187-1	#1503	NEF9	NMF41	MHF1.1	MHF5.5	ECF13	ECF42	ECF44	MHF14.6



Sample	MHF25.1	FAR338	WI-5	MHF18.3	NHF17	MHF14.9	MHF44.1	WI-3	MHF18.1	MHF41.3
Type	Lively	DIT	DIT	DIT	DIT	DIT	DIT Acid	DIT Acid	DIT Acid	DIT Acid
SiO <sub>2</sub>	51.76	52.41	52.44	53.24	53.97	56.06	55.49	60.88	64.11	68.92
TiO <sub>2</sub>	0.82	0.89	1.10	1.30	1.32	1.81	1.54	1.63	1.51	0.71
Al <sub>2</sub> O <sub>3</sub>	15.94	15.69	19.53	15.01	13.27	14.30	15.28	15.24	15.96	15.03
Fe <sub>2</sub> O <sub>3</sub>	10.33	11.5	8.66	12.05	13.88	12.28	12.31	9.19	10.61	6.57
MnO	0.13	0.13	0.12	0.12	0.22	0.18	0.22	0.12	0.06	0.10
MgO	6.38	5.96	6.54	4.99	4.77	2.93	3.48	1.96	1.67	0.60
CaO	11.27	9.35	7.73	8.56	8.68	5.22	5.63	3.01	0.75	0.40
K <sub>2</sub> O	2.46	2.61	2.92	3.21	2.52	4.10	3.76	5.74	3.25	3.51
P <sub>2</sub> O <sub>5</sub>	0.73	0.77	0.22	0.66	1.18	2.06	1.78	1.10	0.75	3.29
Na <sub>2</sub> O	0.17	0.16	0.11	0.17	0.19	0.25	0.20	0.59	0.69	0.21
TOTAL	99.63	99.49	99.36	99.30		99.20	99.69	99.45	99.36	99.34
Loss	1.88	2.10	3.28	3.64		4.12	1.56	2.81	3.30	3.44
Rb	10.9	21.7	29	17.6	34	40.8	40	18.2	20.5	59.6
Sr	231	493.3	582	367	178	358	324	302	1204	715
Nb	10.3	9.1	9.5	12.0	16	18.2	27.4	35.2	74.4	72
Zr	93	96	115	158	192	267	226	309	502	620
Y	21.4	20.0	21.8	30.0	39	41.1	36.2	48.2	85.5	64.8
Cr	172	43	44	71	21	31	11	15	13	19
Ni	63	-	7	95	10	20	6	2.0	8	2.0
Ba	205	494	153	199	371	1409	528	207	2939	862
La	10.8	11.7	16.0	15.7	21.36	26.7	27.0	42.8	47.5	68.9
Ce	21.3	24.1	33.9	32.6	48.50	58.3	55.6	91.9	108	140.0
Pr					6.14					
Nd	11.9	14.7	19.2	20.5	25.93	36.4	29.7	47.0	58.9	69.3
Sm	2.9	3.39	4.41	5.32	6.12	7.92	6.18	10.4	13.2	13.60
Eu	0.92	1.14	1.34	1.63	1.72	2.25	1.77	2.55	3.09	3.41
Gd					6.46					
Tb	0.61	0.60	0.73	0.94	1.07	1.27	1.04	1.55	2.07	2.06
Dy					6.62					
Ho					1.34					
Er					3.91					
Tm					0.62					
Yb	2.41	1.82	2.43	2.81	3.74	3.97	3.62	5.11	7.07	7.47
Lu	0.4	0.28	0.37	0.44	0.056	0.60	0.55	0.76	1.05	1.11
Th	1.68	1.03	1.50	1.49	3.19	2.48	3.90	5.17	5.6	6.54
U	0.8		0.90	0.9	0.77	1.20	1.30	1.70	2.2	2.10
Ta	0.53	0.45	0.62	0.57	0.95	1.04	1.54	2.19	4.32	4.47
Hf	2.32	2.24	3.18	3.78	5.29	5.94	5.53	8.35	11.7	13.80
Pb					13.83					
Cs	0.47		0.28	0.4		0.56	0.38	0.20	0.3	0.20
Co	39.3	21.5	23.20	50.6		35.00	29.4	16.70	12.5	7.40
Sc	37.4	17.1	18.80	15.8		24.70	23.7	18.40	16.3	13.30
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.70564	0.70771±1	0.71042	0.70618		0.70730	0.70730	0.71029	0.71720	0.71745
± 2SE	±15	1	±6	±17		± 17	± 17	± 23	±17	±14
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>182</sub>	0.70510	0.70733	0.71004	0.70581	0.70594	0.70554	0.70637	0.70983	0.71716	0.71714
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512505	0.512540±	0.512533 ±8	0.512572		0.512561	0.512561	0.512535	0.512595	0.512542
± 2SE	±8	9		±5		± 6	± 6	± 7	±7	±7
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>182</sub>	0.512322	0.512374	0.51236	0.512377	0.512259	0.512349	0.512405	0.512369	0.512427	0.512398
εNd <sub>182</sub>	-1.5	-0.6	-0.7	-0.4	-2.8	-0.9	3.1	-0.6	0.6	-0.1
<sup>206</sup> Pb/ <sup>204</sup> Pb	18.781±9	17.980±9	18.067±9	18.430±8				18.016±10	18.326±8	18.393±6
<sup>207</sup> Pb/ <sup>204</sup> Pb	15.670±9	15.510±8	15.512±8	15.633±7				15.508±10	15.576±8	15.529±6
<sup>208</sup> Pb/ <sup>204</sup> Pb	38.300±23	37.361±18	37.425±12	37.944±2				37.492±26	37.880±20	37.583±12

Sample	MHF15.2	MHF14.4	MHF15.1	MA1
Type	MAT	MAT	MAT	MAT
SiO <sub>2</sub>	45.92	47.40	48.56	47.51
TiO <sub>2</sub>	1.81	1.52	1.71	1.39
Al <sub>2</sub> O <sub>3</sub>	15.70	16.39	16.71	16.45
Fe <sub>2</sub> O <sub>3</sub>	13.60	13.23	12.52	12.91
MnO	0.16	0.17	0.21	0.20
MgO	8.00	7.51	7.15	8.27
CaO	10.59	10.65	9.80	10.05
K <sub>2</sub> O	2.59	2.53	2.53	2.65
P <sub>2</sub> O <sub>5</sub>	0.30	0.05	0.17	0.38
Na <sub>2</sub> O	0.20	0.17	0.19	0.20
TOTAL	98.87	99.63	99.56	
Loss	1.5	1.90	2.25	
Rb	4.7	5.5	2.0	10
Sr	424	444	381	396
Nb	11.2	9.9	10.7	7
Zr	80	79	80	90
Y	21.1	18.6	18.2	19
Cr	89	78	95	89
Ni	119	102	121	116
Ba	110	105	103	827
La	8.1	8.0	7.9	4.37
Ce	18.0	17.2	18.3	11.45
Pr				1.83
Nd	13.3	12.8	13.8	9.73
Sm	3.72	3.32	3.78	3.08
Eu	1.38	1.25	1.42	1.29
Gd				3.72
Tb	0.7	0.63	0.72	0.61
Dy				3.66
Ho				0.71
Er				1.97
Tm				0.30
Yb	2.05	1.83	2.04	1.69
Lu	0.32	0.28	0.33	0.23
Th	0.42	0.42	0.48	0.26
U				0.09
Ta	0.58	0.54	0.6	0.3
Hf	2.28	1.94	2.26	2.60
Pb				1.07
Cs	0.3	0.35	0.2	
Co	48.2	42.60	49.3	
Sc	28.9	25.40	28.6	
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.70400	0.70385±17	0.70342	
± 2SE	±14		±15	
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>182</sub>	0.70392	0.70375	0.70338	0.70497
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512747	0.512743	0.512738	
± 2SE	±6	±17	±6	
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>182</sub>	0.512537	0.512548	0.512532	0.512588
εNd <sub>182</sub>	2.8	3.0	2.7	3.6
<sup>206</sup> Pb/ <sup>204</sup> Pb	19.152±6	17.858±7	18.131±7	17.91
<sup>207</sup> Pb/ <sup>204</sup> Pb	15.656±5	15.476±6	15.579±7	15.50
<sup>208</sup> Pb/ <sup>204</sup> Pb	38.464±13	37.425±18	37.762±16	37.57

Table 1. Whole-rock major and trace element and isotopic compositions of Falkland Islands intrusions used in this study. Major elements and isotopic compositions for sample numbers starting with EC, NE, NH, NM and MA are from Mitchell *et al.* (1999) with addition trace elements from this study. See supplementary materials for analytical methods.

Table 2. Geochemical, mineralogical and petrographical characteristics of the different groups of Falkland Islands intrusions.

Type	Type locality	Mitchell <sup>1</sup>	Stone <sup>2</sup>	Petrographic features	Mineralogy	Subgroup	Mg#	SiO <sub>2</sub>	TiO <sub>2</sub>	Ti/Zr	Zr/Y	<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>182</sub>	εNd <sub>182</sub>
Port Sussex Creek (PST)	Port Sussex 51°40'15" S 58°58'41" W	N-S	NE-SW	Coarse-grained dolerite	Pig ± Opx + Aug Rare Ol + Di	none	48-58	52-54	0.9-1.2	50-70	3.6-5.3	0.7077 -0.7134	-5.5 to -10.9
E-W	Fox Bay West 51°57'02" S 60°05'21" W	E-W	E-W	Coarse-grained olivine dolerite	Ol + Plag ± Aug	none	42-64	47-54	1.0-1.9	77-90	3.2-4.8	0.7036-0.7058	-0.4 to +3.0
Lively Island (LI)	Lively Island 52°00'00" S 58°27'47" W	Lively Island	NE-SW	Coarse-grained with accessory biotite	Ol + Plag + Aug ± rare pigeonite	none	48-52	51-52	0.8-0.9	53	4.0-4.54	0.7053	-0.5 to -1.4
Dyke Island (DIT)	Dyke Island 51°59'33" S 60°52'50" W	Not defined	Radial swarm	Fine-grained aphyric, rarely plagioclase ± augite-phyric	Plag + Aug	Acid	<22	62-75	0.2-1.6	<31	5.0-8.8	0.7055-0.7098	-2.8 to -0.5
						Low TiO <sub>2</sub>	27-57	52-61	1.1-1.7	24-67	4.8-7.4		
						High TiO <sub>2</sub>	41-51	53-58	>1.80	25-53	6.8-8.4		
Mount Alice (MAT)	Mount Alice 52°09'12" S 60°35'55" W	Mount Alice	Radial swarm	Fine-grained plagioclase ± olivine phyric	Ol + Plag ± Aug	none	44-64	47-50	1.3-1.9	98-142	3.2-5.2	0.7031-0.7039	0.0 to +3.7

1. Groups described by Mitchell *et al.* (1999); 2. Groups defined by Stone *et al.* (2009)

Calculated extract for fractionation of PST and E-W intrusions

	PST			E-W		
	NGF16	MHF5.1	Calc	ECF12	ECF44	Calc
SiO <sub>2</sub>	54.01	53.81	53.82	49.69	51.03	50.98
TiO <sub>2</sub>	0.94	1.00	1.13	1.05	1.21	1.36
Al <sub>2</sub> O <sub>3</sub>	13.20	14.97	15.02	13.30	15.21	15.20
FeO	10.60	10.62	10.60	10.51	11.11	11.06
MnO	0.20	0.17	0.18	0.17	0.17	0.21
MgO	9.67	6.78	6.78	11.62	6.71	6.70
CaO	8.81	9.50	9.49	9.81	11.51	11.50
Na <sub>2</sub> O	1.96	2.62	2.56	2.29	2.56	2.46
K <sub>2</sub> O	0.48	0.42	0.56	0.32	0.39	0.39
P <sub>2</sub> O <sub>5</sub>	0.12	0.11	0.15	0.11	0.11	0.13
Extract			%			%
Olivine		Fo <sub>83</sub>	0.0			57.0
Plagioclase		An <sub>70</sub>	18.9			40.4
Pyroxene		En <sub>71</sub> Fs <sub>19</sub> Wo <sub>9</sub>	74.7			
Pyroxene		En <sub>51</sub> Fs <sub>13</sub> Wo <sub>33</sub>	6.4			2.6
Σ residuals <sup>2</sup>			0.127			0.038
F			0.79			0.75

Table 4 AFC parameters for the trajectories shown in Fig. 12. R is the ratio of assimilated rock to crystal cumulate. A value appropriate for upper-crustal contamination has been used. F is the total amount of crystallization required to reach the most extreme composition on a particular trajectory.  $T_{\text{CHUR}}$  is the Chondritic Uniform Reservoir model Nd age for the most extreme composition on a particular trajectory, in Ga.

		AFC parameters								$T_{\text{CHUR}}$
		Sr	Nd	$\epsilon\text{Nd}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$D_{\text{Sr}}$	$D_{\text{Nd}}$	R	F	(Ga)
<b>CT1</b>	Source	50	5	2	0.7035	0.5	0.1	0.40	$\leq 0.2$	3.0
	Crust	400	20	-50	0.7120					
<b>PST-1</b>	Source	60	4	2	0.7035	0.5	0.1	0.40	$\leq 0.2$	2.2
	Crust	350	40	-20	0.7200					
<b>PST-2</b>	Source	100	5	2	0.7035	0.5	0.1	0.40	$\leq 0.2$	1.8
	Crust	350	60	-10	0.7250					